

Particles and Cosmology  
16th Baksan School on Astroparticle Physics

# Final results of the search for $\nu_\mu \rightarrow \nu_e$ oscillations with the OPERA detector in CNGS beam

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# The OPERA experiment



## Goals

- ▶ The observation of the  $\nu_\mu \rightarrow \nu_\tau$  oscillations in the appearance mode through the detection of the  $\tau$ -lepton [PRL 120, 211801 \(2018\)](#)
- ▶ The tracking capabilities of the emulsion allow to identify electrons produced in CC interactions of  $\nu_e$  and, hence, to study  $\nu_\mu \rightarrow \nu_e$  oscillations in appearance mode [JHEP06\(2018\)151](#)

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## Experimental requirements:

- ▶ long baseline
- ▶ high energy neutrinos
- ▶ high intensity beam
- ▶ low background
- ▶ large active target mass
- ▶  $\mu\text{m}$  space resolution

# The OPERA experiment

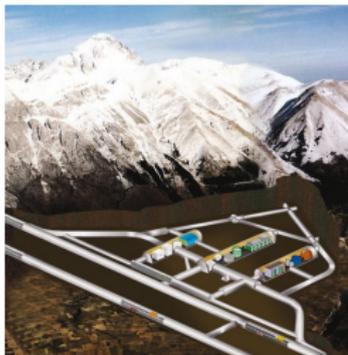


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baseline  $\sim 732$  km

## LNGS

$\sim 3\,800$  m w.e.

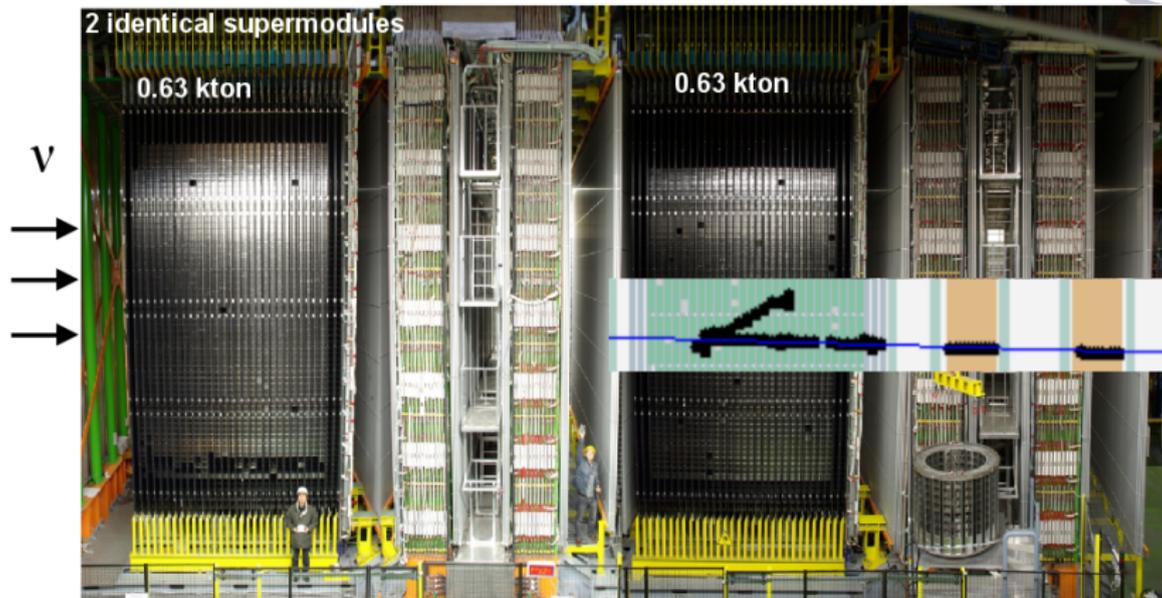
$\sim 1$  cosmic muon per  
( $m^2 \times h$ )

## CNGS beam

$\langle E_{\nu_\mu} \rangle$ (GeV)	17
$(\nu_e + \bar{\nu}_e)/\nu_\mu$	0.87%
$\bar{\nu}_\mu/\nu_\mu$	2.1%
$\nu_\tau$ prompt	Negligible

interactions rates at LNGS site

# The OPERA experiment



Target and Target Tracker  
( $6.7\text{ m} \times 6.7\text{ m}$ )  
~75000 bricks

Muon  
Spectrometer  
( $8\text{ m} \times 10\text{ m}$ )

Target and Target Tracker  
( $6.7\text{ m} \times 6.7\text{ m}$ )  
~75000 bricks

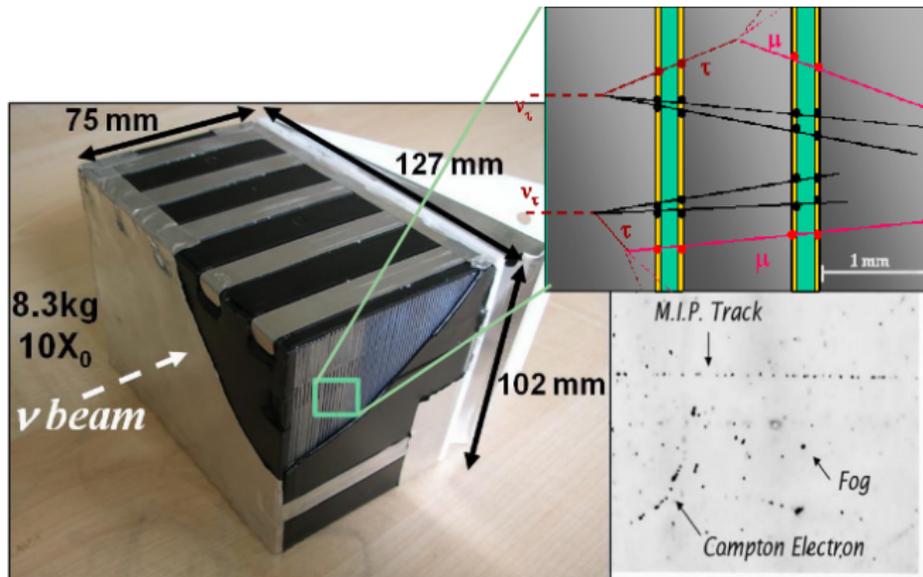
Brick  
Manipulator  
System

# The OPERA experiment



Emulsion Cloud Chamber technique provides large target mass and high spatial resolution:

- ▶  $\sim 150\,000$  ECC, 56 lead plates and 57 emulsions each
- ▶  $\sim 9$  million films in total (sensitivity 30 grains per  $100\ \mu\text{m}$ )
- ▶  $\sim 1.25$  kton total target mass



# Collected data sample

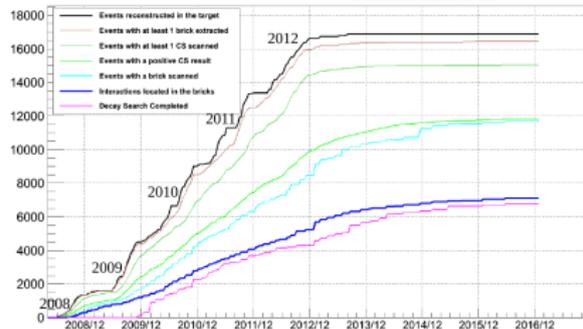
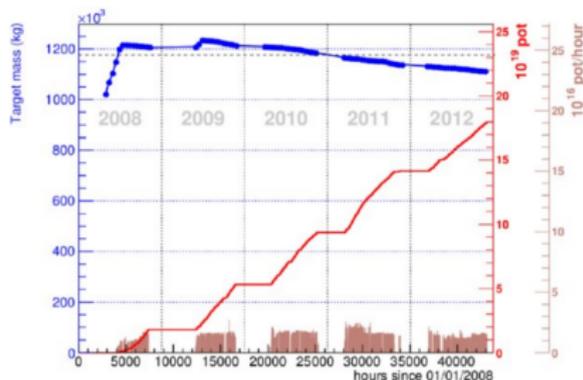


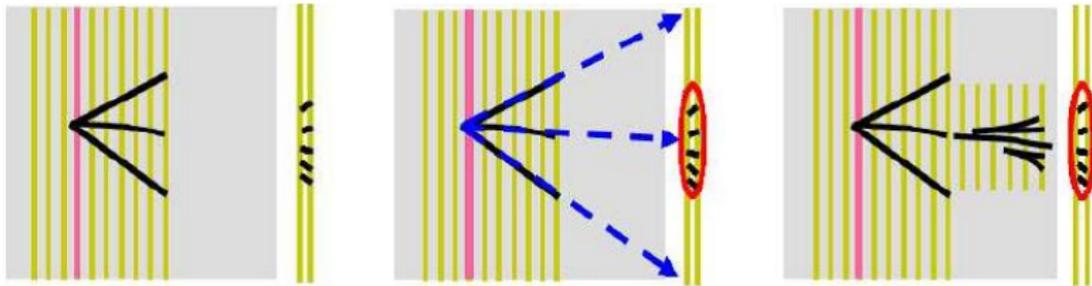
## 2008-2012 CNGS run

- ▶  $17.97 \times 10^{19}$  p.o.t.
- ▶ 1.18 kt average detector mass
- ▶ 19505 on-time interactions in detector
- ▶ 6785 decay searched events\*

\* Decay searched events:

- ▶ vertex is located in the lead/emulsion target
- ▶ event topology is reconstructed





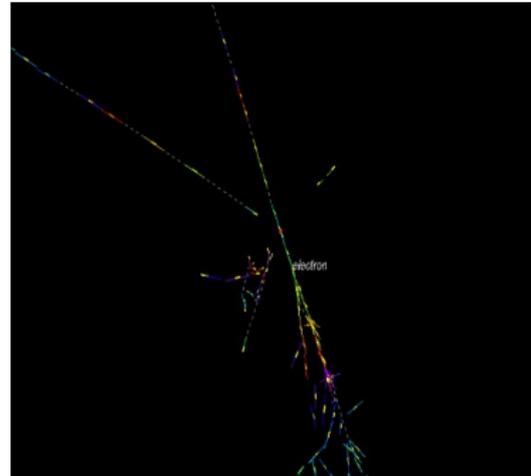
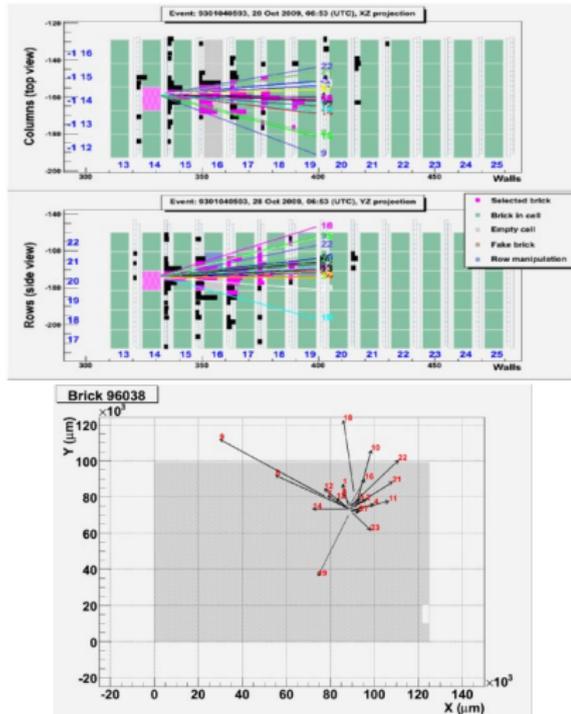
Search for  $\nu_e$  candidates [JHEP 1307 \(2013\) 004](#)

- ▶ The electron identification is based on the search of associated electromagnetic shower. Primary tracks extrapolated to the changeable emulsion doublets. The tracks with angles and positions similar to projection ones are searched ( $\Delta\theta < 150\text{mrad}$ , 2 mm).
- ▶ If 3 or more tracks found, an additional volume along the candidate track is scanned.

# Search for $\nu_e$ CC interactions

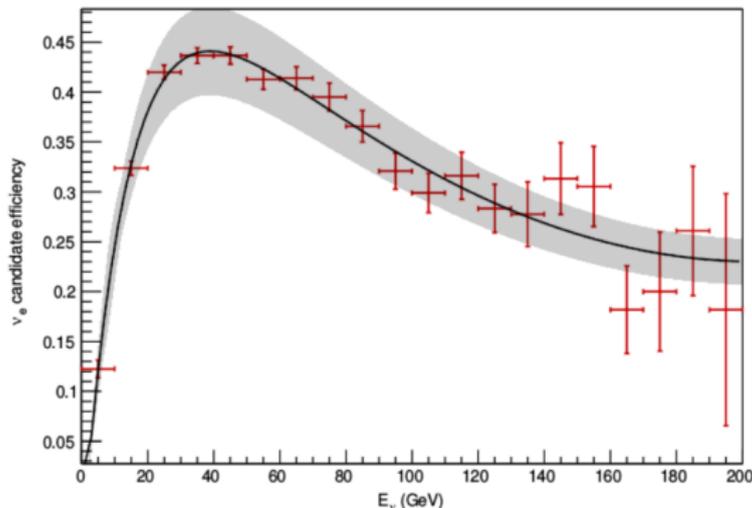


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one  $\nu_e$  candidate event

# Search for $\nu_e$ CC interactions



The  $\nu_e$  CC candidates selection efficiency as a function of neutrino energy (red error bars — MC statistical errors; grey area represents the systematical error.)

## Analysed data sample

- ▶  $17.97 \times 10^{19}$  p.o.t.
- ▶ 19505 on-time events
- ▶ 5868 vertexes in 1st or 2nd most probable brick
- ▶ 1281 events tagged as  $0\mu$
- ▶ 35  $\nu_e$  candidates were found

# Background and $\nu_e$ beam contamination



- ▶ Intrinsic  $\nu_e(\bar{\nu}_e)$  beam components
- ▶  $\nu_\tau$  CC with the decay of  $\tau \rightarrow e$
- ▶  $\pi^0 \rightarrow \gamma \rightarrow e^+ + e^-$  in  $\nu$  interaction without a reconstructed  $\mu$

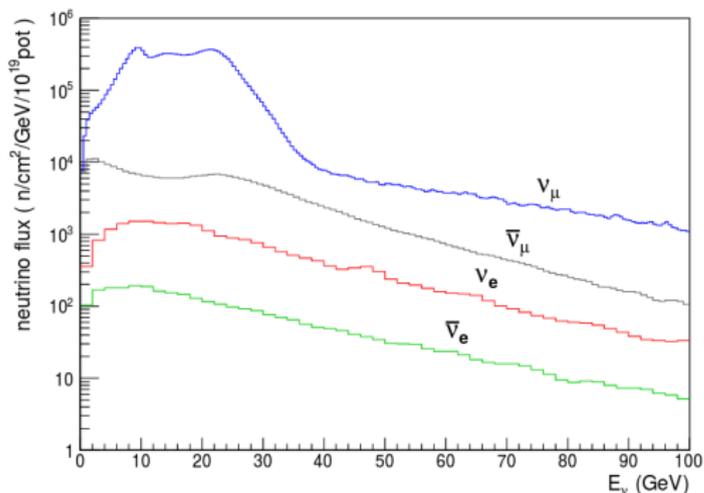
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Neutrino fluxes at Gran Sasso JHEP 1307  
(2013) 004



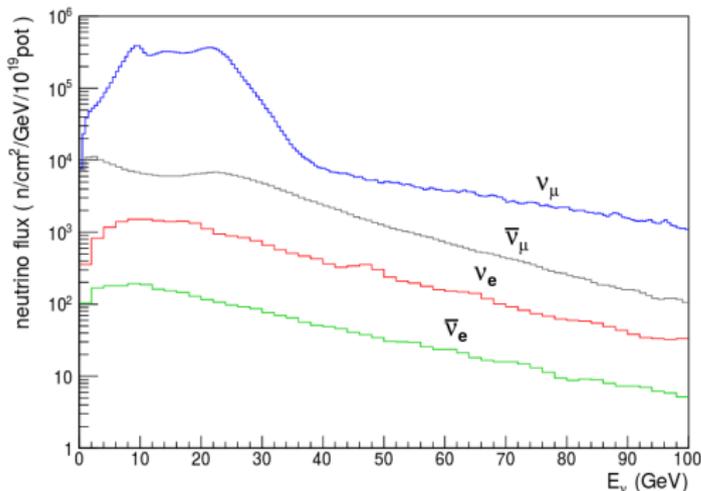
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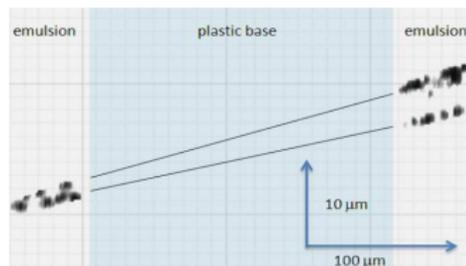
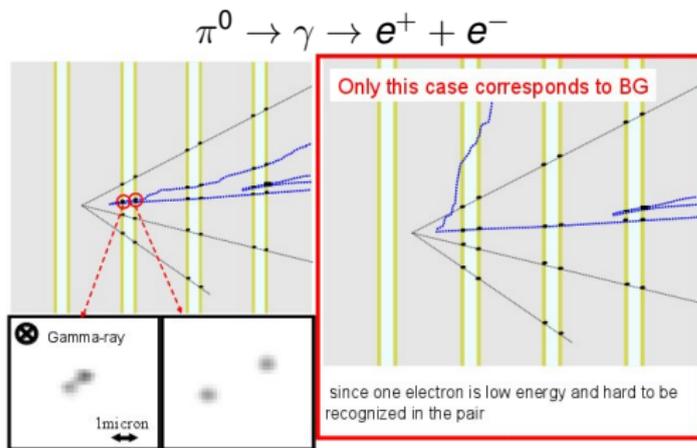
- ▶  $\tau \rightarrow e$  in the same lead plate as the primary vertex with  $\text{IP} < 10 \mu\text{m}$  (an impact parameter of the  $e$  to the 1ry vertex).
- ▶ An undetected kink ( $\theta_{\text{kink}} < 20 \text{ mrad}$ ) from a  $\tau \rightarrow e$  in further downstream material.

**$0.7 \pm 0.2(\text{sys.})$  bg events**

# Background and $\nu_e$ beam contamination



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Side view of  $e^+e^-$  pair detected in the emulsion film

[JHEP 1307 \(2013\) 004](#)

BG evaluated from the data:

**$0.5 \pm 0.5(\text{stat.})$  events**

- ▶  $e^+e^-$  can not be distinguished from a single particle in the first 2 emulsion films after the vertex
- ▶ one pair component undetected

## No oscillation scenario

- ▶  $\nu_e$  beam contamination  
 $30.7 \pm 0.9(\text{stat.}) \pm 3.1(\text{syst.})$
- ▶ other background  
 $1.2 \pm 0.5(\text{stat.}) \pm 0.2(\text{syst.})$

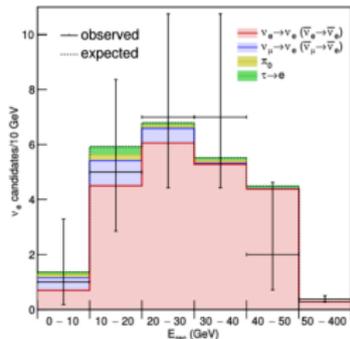
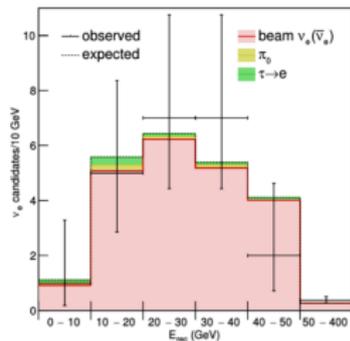
## 3-flavour oscillation scenario

- ▶  $34.3 \pm 1.0(\text{stat.}) \pm 3.4(\text{syst.})$  (including BG)

## 2008-2012 data ( $17.97 \times 10^{19}$ p.o.t.)

- ▶ 35  $\nu_e$  candidates found

**Result:**  $\sin^2(2\theta_{13}) < 0.43$  (90% C.L.)



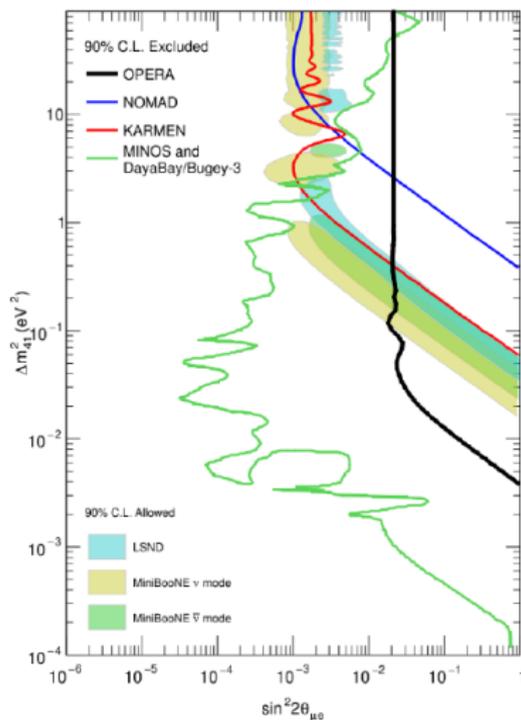
## Sterile neutrino search

The excess of  $\nu_e$  and  $\bar{\nu}_e$  reported by LSND and MiniBooNE can be interpreted as due to the presence of light sterile neutrino.

3+1 model bounds from  $\nu_e$  appearance with profile Likelihood method.

An upper limit on  $\sin^2(2\theta_{\mu e}) = 0.021$  is set for  $\Delta m_{41}^2 > 0.1 \text{ eV}^2$ .

Moreover, OPERA contributes to limit the effective mixing for low  $\Delta m_{41}^2$  and excludes  $\Delta m_{41}^2 \gtrsim 4 \times 10^{-3} \text{ eV}^2$  for maximal mixing.





- ▶ Search for  $\nu_e$  in full data set analysis is completed ( $17.97 \times 10^{19}$  p.o.t.)
- ▶ Number of observed  $\nu_e$  candidates is in agreement with the expected background and the standard oscillation signal
- ▶ Constraint on sterile neutrinos in the 3+1 flavour model
- ▶ Combined  $\nu_\mu \rightarrow \nu_e$  and  $\nu_\mu \rightarrow \nu_\tau$  appearance analysis in progress...

# Thank you for attention!



Belgium  
IIHE-ULB Brussels



Croatia  
IRB Zagreb



France  
LAPP Annecy  
IPHC Strasbourg



Germany  
Hamburg



Israel  
Technion Haifa



Italy  
Bari  
Bologna  
Frascati,  
L'Aquila  
LNGS  
Naples  
Padova  
Rome  
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Japan  
Aichi  
Toho  
Kobe  
Nagoya  
Nihon



Korea  
Jinju



Russia  
INR RAS Moscow  
LPI RAS Moscow  
SINP MSU Moscow  
JINR Dubna



Switzerland  
Bern



Turkey  
METU, Ankara



<http://operaweb.lngs.infn.it>



# Backup slides

$N_{beam}$  was obtained using a data-driven approach from the number of observed events with no charged leptons:

$$N_{beam} = n_{0l} \frac{R_{\nu_e}^{CC} \langle \varepsilon_{CC}^{\nu_e}(\nu_e) \rangle}{\sum_{i=\mu, e} \sum_{j=CC, NC} R_{\nu_i}^j \langle \varepsilon_j^{0l}(\nu_i) \rangle}, \quad (1)$$

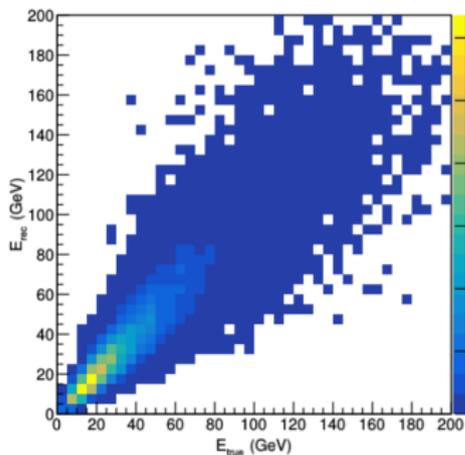
where

$$\langle \varepsilon_j^{\nu_e(0l)}(\nu_i) \rangle = \int \phi_{\nu_i} \epsilon_j^{\nu_e(0l)} \sigma_{\nu_i}^j dE \Big/ \int \phi_{\nu_i} \sigma_{\nu_i}^j dE, \quad (2)$$

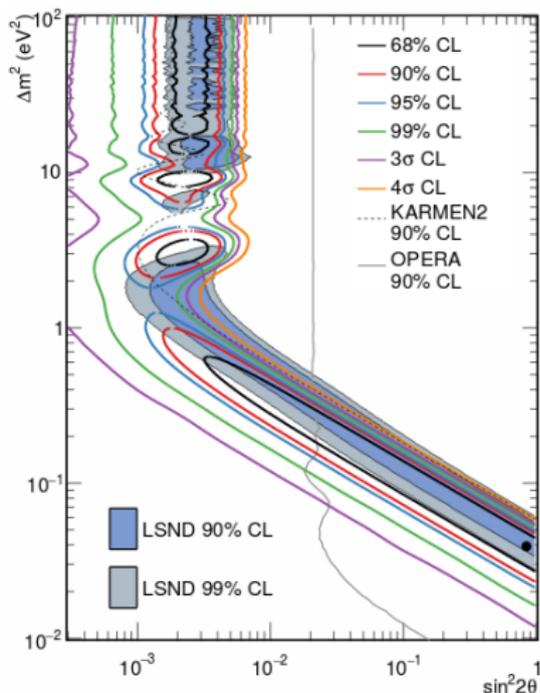
while  $R_{\nu_i}^j$  are the interaction rates of neutrino and antineutrino:

$$R_{\nu_i}^j = \int \phi_{\nu_i} \sigma_{\nu_i}^j dE. \quad (3)$$

# 3+1 analysis



$$\frac{\Delta E}{E} = 0.18 + \frac{0.55}{\sqrt{E(\text{GeV})}}$$



MiniBooNE allowed region  
[arXiv:1805.12028 \[hep-ex\]](https://arxiv.org/abs/1805.12028)

$$P_{\nu_{\mu} \rightarrow \nu_e} \sim \text{standard oscillation} + \text{Exotic oscillation}$$

$$P_{\nu_{\mu} \rightarrow \nu_e} \sim C^2 \sin^2 \Delta_{31} + \sin^2 2\theta_{\mu e} \sin^2 \Delta_{41}$$

$$\text{Interference term} \left\{ \begin{array}{l} + 0.5 C \sin 2\theta_{\mu e} \cos \phi_{\mu e} \sin 2\Delta_{31} \sin 2\Delta_{41} \\ - C \sin 2\theta_{\mu e} \sin \phi_{\mu e} \sin^2 \Delta_{31} \sin 2\Delta_{41} \\ + 2 C \sin 2\theta_{\mu e} \cos \phi_{\mu e} \sin^2 \Delta_{31} \sin^2 \Delta_{41} \\ + C \sin 2\theta_{\mu e} \sin \phi_{\mu e} \sin 2\Delta_{31} \sin^2 \Delta_{41} \end{array} \right.$$

$$C = 2|U_{\mu 3} U_{e 3}^*|$$

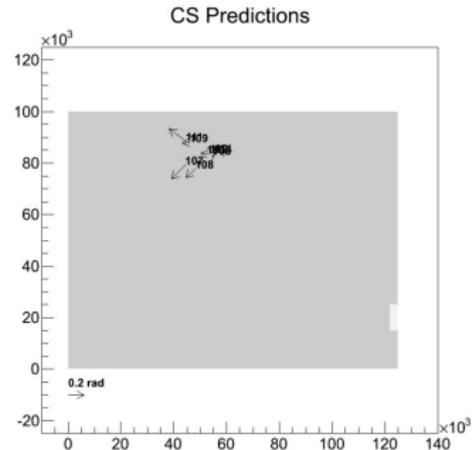
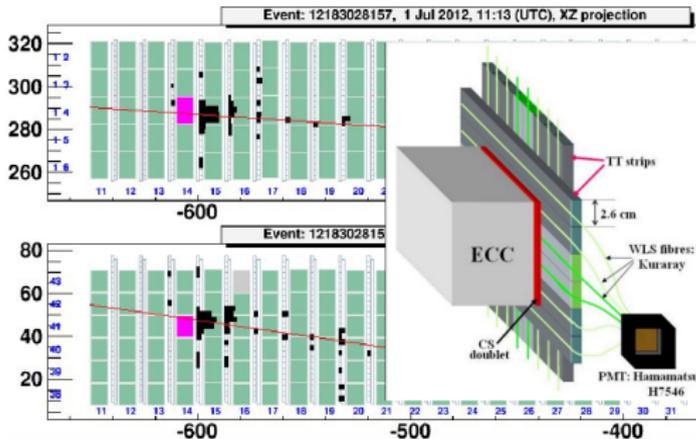
$$\Delta_{ij} = \frac{1.27 \Delta m_{ij}^2 L}{E}$$

$$\phi_{\mu e} = \text{Arg}(U_{\mu 3} U_{e 3}^* U_{\mu 4} U_{e 4}^*)$$

$$\sin^2 2\theta_{\mu e} = 4|U_{\mu 4}|^2 |U_{e 4}|^2$$

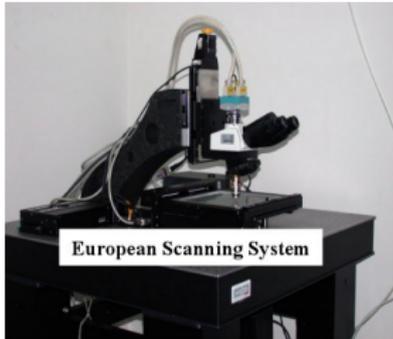
$$-2 \ln L = -2 \sum_i^N (n_i \ln \mu_i - N \mu_i) + \sum_{j=1}^2 \frac{k_j^2}{\sigma_j^2} + \frac{(\Delta m_{31}^2 - \widehat{\Delta m_{31}^2})^2}{\sigma_{\Delta m_{31}^2}^2}. \quad (4)$$

# Event location procedure

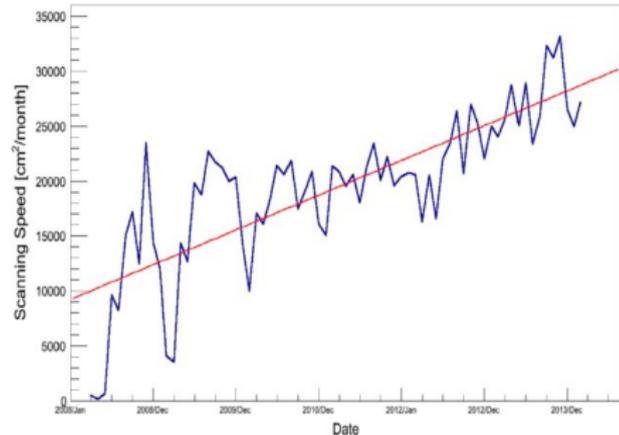


- ▶ TT data is used for a prediction of the bricks which contain the neutrino interactions
- ▶ A large area of the corresponding changeable film is scanned (so far  $2'500'000 \text{ cm}^2$  of CS surface analysed)

## Scanning of Changeable Sheets: two large facilities

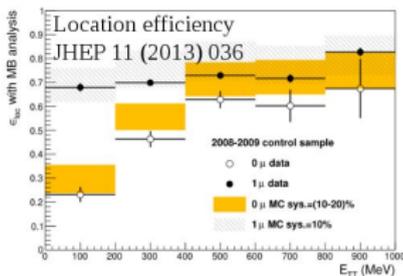
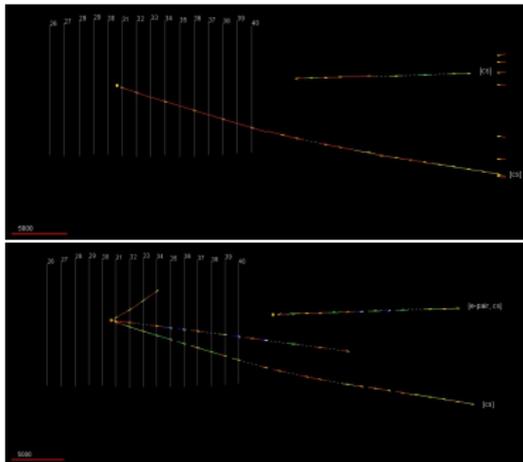


## Scanning speed per facility: improvement during the run



- ▶ LNGS: 10 microscopes, 200  $\text{cm}^2/\text{h}$
- ▶ Nagoya: 5 S-UTS, 220  $\text{cm}^2/\text{h}$

# Event location procedure



- ▶ brick exposure at the surface laboratory to collect cosmic-rays for alignment
- ▶ scan-back: CS-tracks are followed upstream from film to film to find the  $\nu$ -interaction vertex
- ▶ total-scan: scanning of the  $1 \text{ cm}^2$  around the vertex in 15 plates is performed
- ▶ scan-forth: improvement of the momentum measurement of the reaction products [New J. of Phys. 4 \(2012\) 013026](#)
- ▶ decay search [Eur.Phys.J. C74\(2014\) 2986](#)



- ▶ Primary vertex definition
  - ▶ visual inspection of segments on the vertex plate
  - ▶ impact parameter  $< 10(5 + 0.01\Delta z)\mu m$ , if  $\Delta z < 500\mu m$
- ▶ Extra-track search
  - ▶ selection of tracks reconstructed in the volume but not attached to primary vertex
  - ▶ identification of  $e^+e^-$  pairs by visual inspection
- ▶ In-track search
  - ▶ search for small kinks along the tracks attached to the primary vertex
- ▶ Parent search
  - ▶ search for a track connecting the selected extra-track and the primary vertex

(more details: [arXiv:1404.4357 \[hep-ex\]](https://arxiv.org/abs/1404.4357))