

Search for new physics in Mu2e experiment

JINR team

Artikov A.M., Atanov N.V., Atanova O.S.,
Azaryan N.S., Baranov V.Yu., Batusov
Yu.A., Budagov J.A., Chokheli D.,
Davydov Yu.I., Demin D.L., Flyagin V.B.,
Glagolev V.V., Guskov D.S., Kazakov
D.I., Kharzheev Yu.N., Kozlov G.A.,
Kolomoets V.I., Kolomoets S.M.,
Korenkov V.V., Kulchitsky Y.A.,
Kuchinsky N.A., Lyablin M.V., Romanov
V.M., Sazonova A.V., Shalyugin A.N.,
Simonenko A.V., Studenov S.N.,
Suhanova A.K., Suslov I.A., Tarasov
O.V., Tereschenko V.V., Tereschenko
S.V., Titkova I.V., Usubov Z.U., Uzhinsky
V.V.,



Mu2e

**Search for
neutrinoless conversion of a
muon into an electron in the
field of a nucleus**

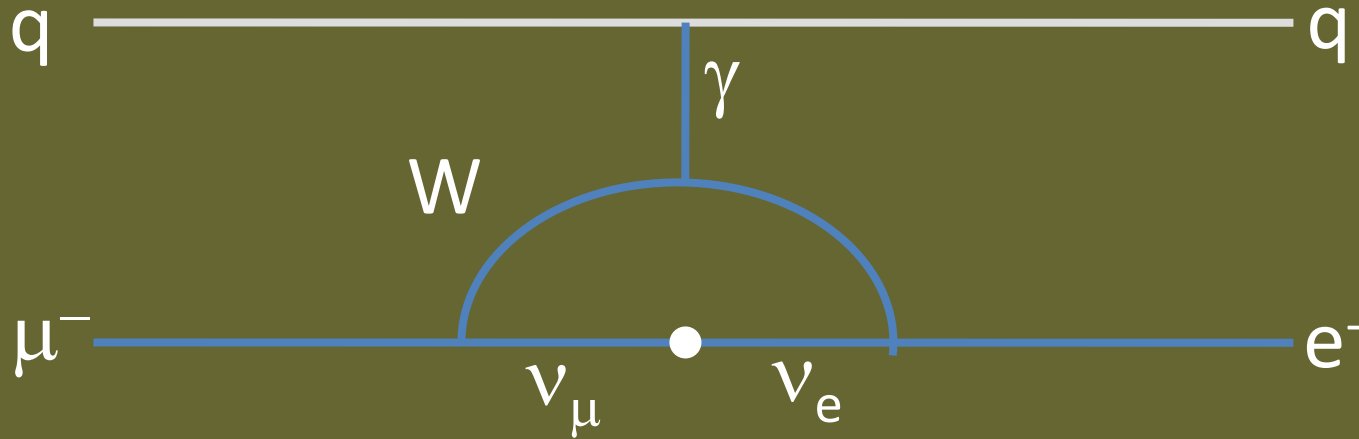
$$\mu^- N \rightarrow e^- N$$

**Charged Lepton Flavor
Violation**

Flavor Violation

- We've known for a long time that quarks mix → (Quark) Flavor Violation
 - Mixing strengths parameterized by CKM matrix
- In last 15 years we've come to know that neutrinos mix → Lepton Flavor Violation (LFV)
 - Mixing strengths parameterized by PMNS matrix
- Why not charged leptons?
 - Charged Lepton Flavor Violation (CLFV)

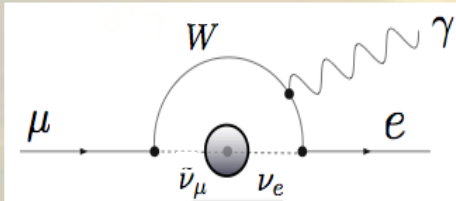
CLFV in the Standard Model



- Strictly speaking, forbidden in the SM
- Even in ν -SM, extremely suppressed
(rate $\sim \Delta m_\nu^2 / M_w^2 < 10^{-50}$)
- However, most all NP models predict rates observable at next generation CLFV experiments

Mu2e : SM prediction and New Physics

The BR of CLFV processes in the Standard Model



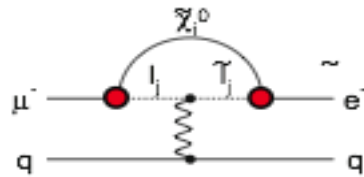
$$\text{BR}(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{ei} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 < 10^{-54}$$

Mu2e sensitivity
is $6 \cdot 10^{-17}$

NP

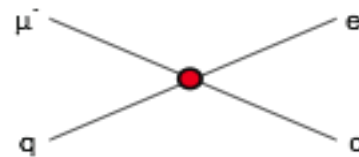
Supersymmetry

rate $\sim 10^{-15}$



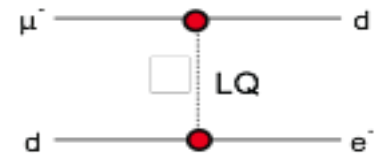
Compositeness

$\Lambda_c \sim 3000 \text{ TeV}$



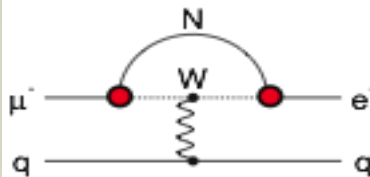
Leptoquark

$M_{LQ} = 3000 (\lambda_{\mu d} \lambda_{e d})^{1/2} \text{ TeV}/c^2$



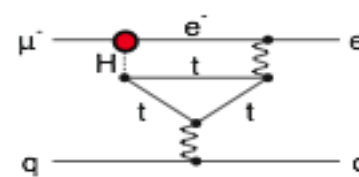
Heavy Neutrinos

$|U_{\mu N} U_{e N}|^2 \sim 8 \times 10^{-13}$



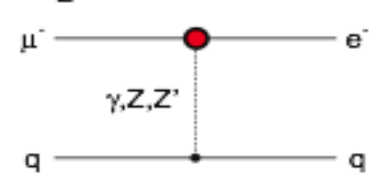
Second Higgs Doublet

$g(H_{\mu e}) \sim 10^{-4} g(H_{\mu \mu})$



Heavy Z' Anomal. Z Coupling

$M_{Z'} = 3000 \text{ TeV}/c^2$

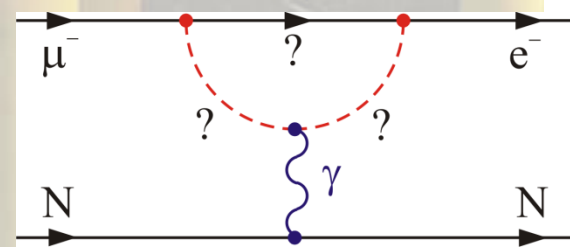


$$\mathcal{L}_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1 + \kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{u}_L \gamma^\mu u_L + \bar{d}_L \gamma^\mu d_L)$$

Mu2e : Sensitivity to High Mass Scales

$$L_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1 + \kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{u}_L \gamma^\mu u_L + \bar{d}_L \gamma^\mu d_L)$$

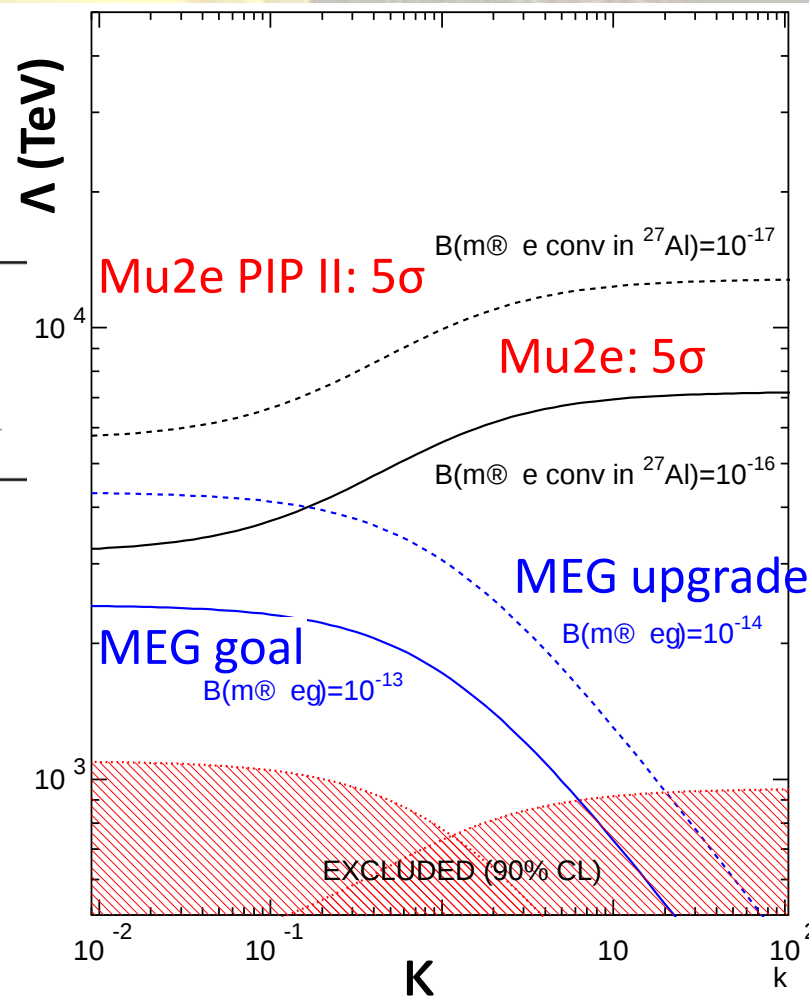
Loops dominate
for $\kappa \ll 1$



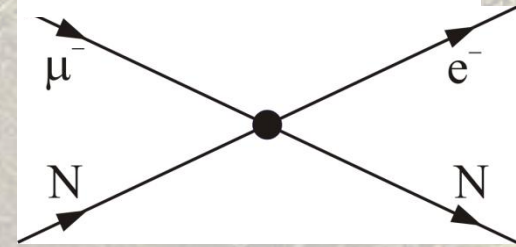
$\mu \rightarrow e \gamma$

$\mu N \rightarrow e N$

$\mu \rightarrow e e e$



Contact terms
dominate for $\kappa \gg 1$



~~$\mu \rightarrow e \gamma$~~

$\mu N \rightarrow e N$

$\mu \rightarrow e e e$

**The great-grandparents of the Mu2e
(MELC, 1992; MECO, 1997)
are INR scientists
V.M. Lobashev and R.M. Djilkibaev**



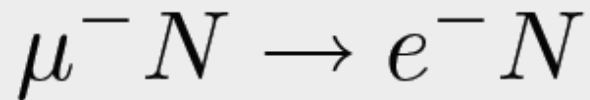
**Владимир Михайлович Лобашев
(29.07.1934—03.08.2011)**



Mu2e Muon-to Electron Conversion

Mu2e will measure the ratio of the coherent neutrinoless muon-to-electron conversion rate to muon capture rate

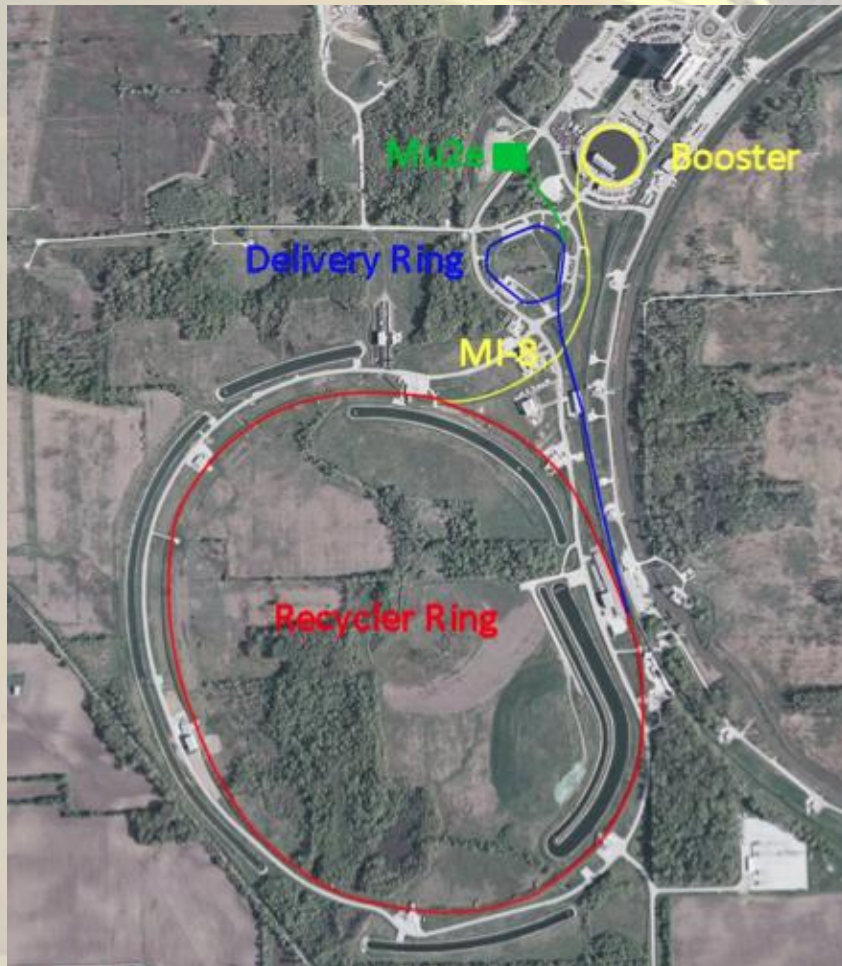
muon converts to electron in the field of a nucleus



$$R_{\mu e} = \frac{\Gamma(\mu^{-} + N(A, Z) \rightarrow e^{-} + N(A, Z))}{\Gamma(\mu^{-} + N(A, Z) \rightarrow \text{all muon captures})}$$

- Charged Lepton Flavor Violation (CLFV)
 - manifest Beyond-Standard-Model physics
 - SES of 2.3×10^{-17} , 0.4 evt bkg; 6×10^{-17} at 90% CL
Requires about 10^{18} stopped muons; about 10^{20} protons on target

The Mu2e Proton Beam



- Mu2e begins by using protons to produce pions
- Mu2e will repurpose much of the Tevatron anti-proton complex to instead produce muons.
- Mu2e can (and will) run simultaneously with NOvA.

The Measurement Method

- Stop negative muons in an **aluminum** target

- The stopped muons form muonic atoms

- 207x smaller radius than inner e^- in Al $\rightarrow$

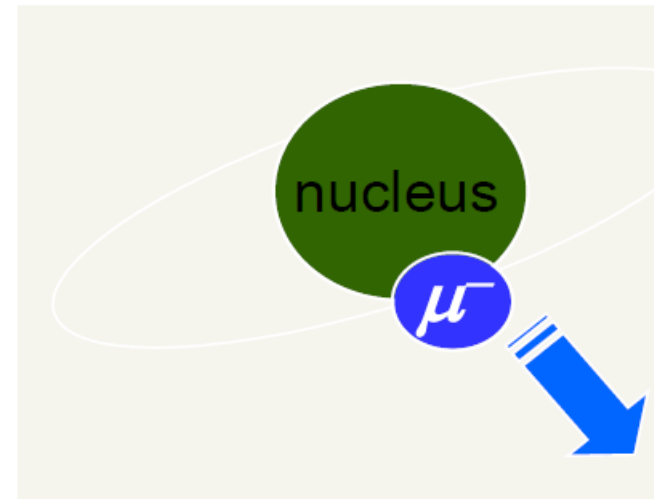
- well inside electron orbits $\rightarrow$

muon forms a hydrogen-like atom, unaffected by e^- 's

- hydrogenic 1S : Bohr radius ~ 20 fm, BE ~ 500 keV

- Nuclear radius ~ 4 fm $\rightarrow$

muon and nuclear wavefunctions overlap significantly



- Three main things can happen (numbers for case of Al):

- Muon decays (40%): $\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$

- Muon captures on the nucleus (60%): $\mu^- + {}_{13}^{27}\text{Al} \rightarrow X + \nu_\mu (\text{capture})$

(capture is roughly sum of reactions with protons in nucleus: $\mu^- + p \rightarrow \nu_\mu + n$)

- Muon to electron conversion: $\mu^- + {}_{13}^{27}\text{Al} \rightarrow {}_{13}^{27}\text{Al} + e^-$

- Muon lifetime in 1S orbit of aluminum ~ 864 ns

(40% decay, 60% nuclear capture), compared to 2.2 μsec in vacuum

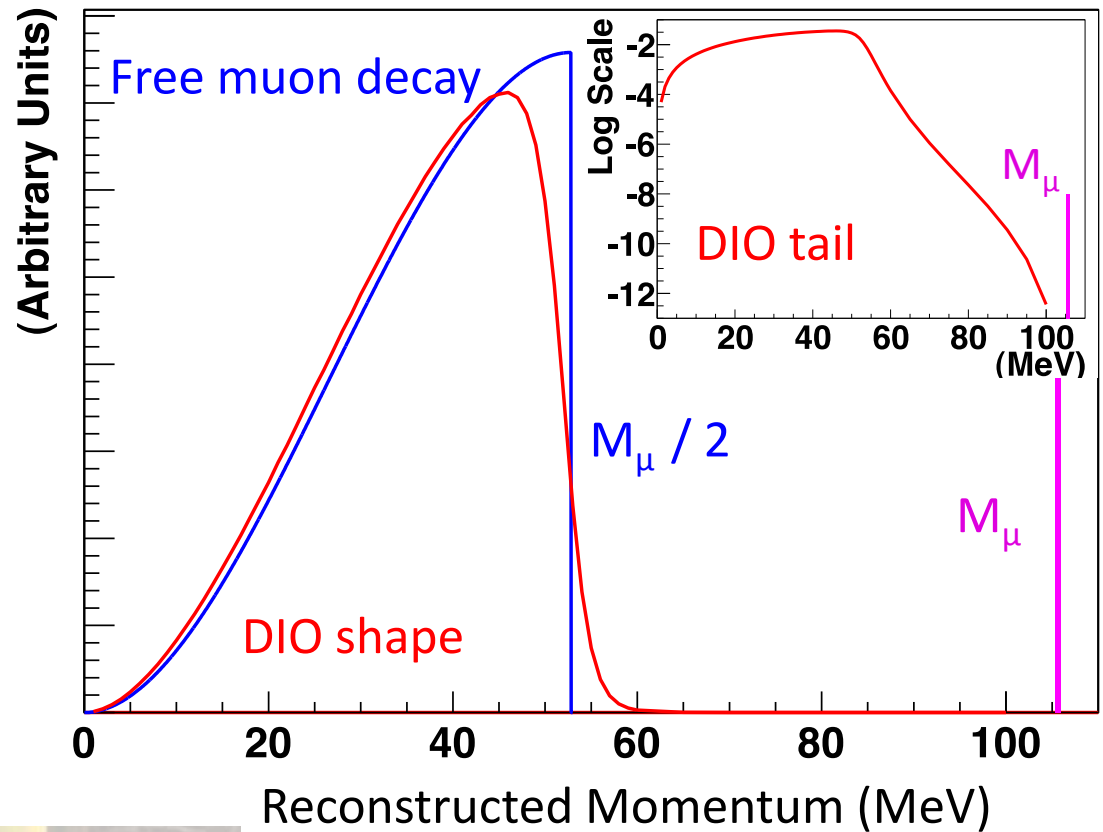
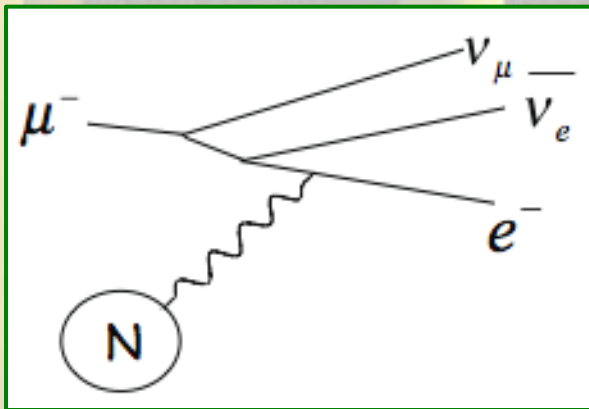
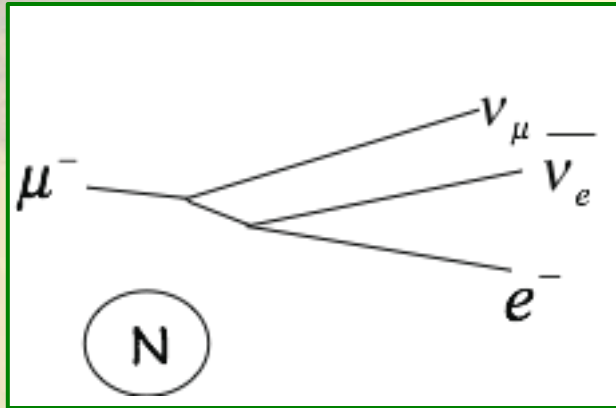
- Look for 105 MeV conversion electron signal $E_e = m_\mu - E_{\text{recoil}} - E_{1\text{S-B.E.}}$ $E_e = 104.96$ MeV

Backgrounds

- Stopped Muon induced
 - Muon decay in orbit (DIO)
 - Out of time protons or long transit-time secondaries
 - Radiative pion capture; Muon decay in flight
 - Pion decay in flight; Beam electrons
 - Anti-protons
 - Secondaries from cosmic rays
-
- Mitigation:
 - Excellent momentum resolution
 - Excellent extinction plus delayed measurement window
 - Thin window at center of TS absorbs anti-protons
 - Shielding and veto

Decay-in-Orbit: Dominant Background

DIO: Decay in orbit

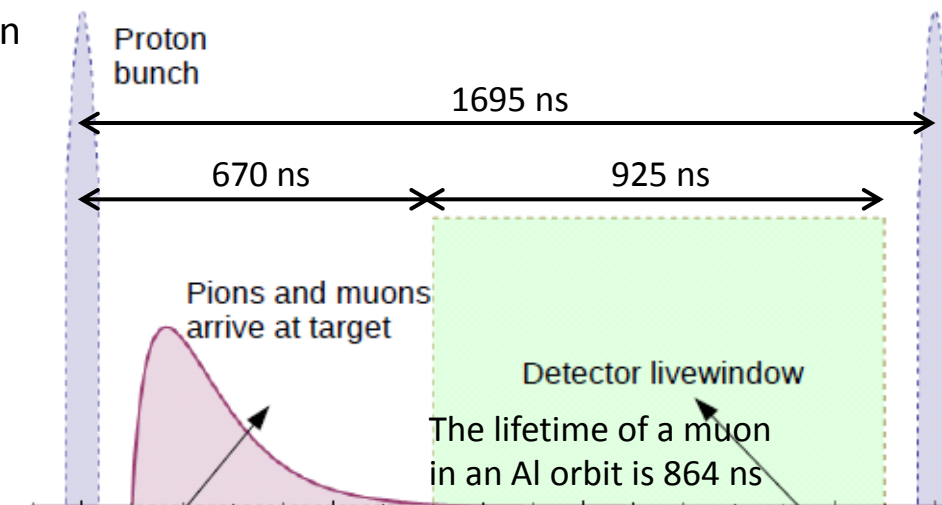


Prompt Background Suppression

- Prompt background

- Happens around the time, when the beam arrives at the target.
- Sources
 - beam electrons,
 - muon decay in flight,
 - pion decay in flight,
 - radiative pion capture
- May create electrons with energies in the signal region

- Prompt background can be suppressed by not taking data during the first 670 ns after the peak of the proton pulse.



Prompt: Radiative Pion Capture with pair production

Delayed: Muon Decay-in-Orbit

- However, this prompt background cannot be eliminated entirely, since some of the protons arrive “out of time”.
 - A ratio of 10^{-10} is required for the beam between pulses vs. the beam contained in a pulse.

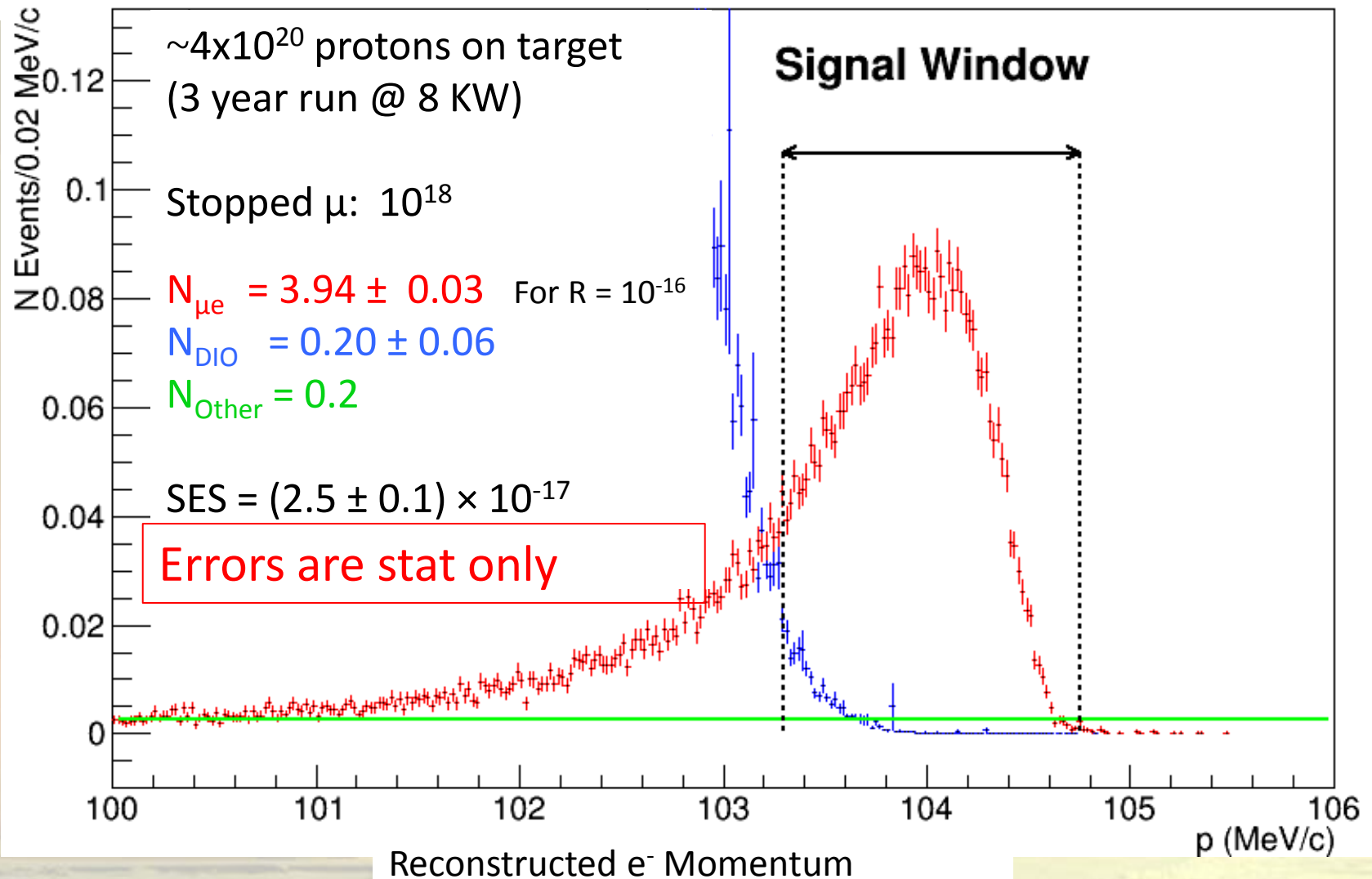
Backgrounds for 3 Year Run

Source	Events	Comment
μ decay in orbit (DIO)	0.20 ± 0.06	
Anti-proton capture	0.10 ± 0.06	
Radiative π^- capture*	0.04 ± 0.02	From protons during detection time
Beam electrons*	0.001 ± 0.001	
μ decay in flight*	0.010 ± 0.005	With e^- scatter in target
Cosmic ray induced	0.050 ± 0.013	Assumes 10^{-4} veto inefficiency
Total	0.4 ± 0.1	

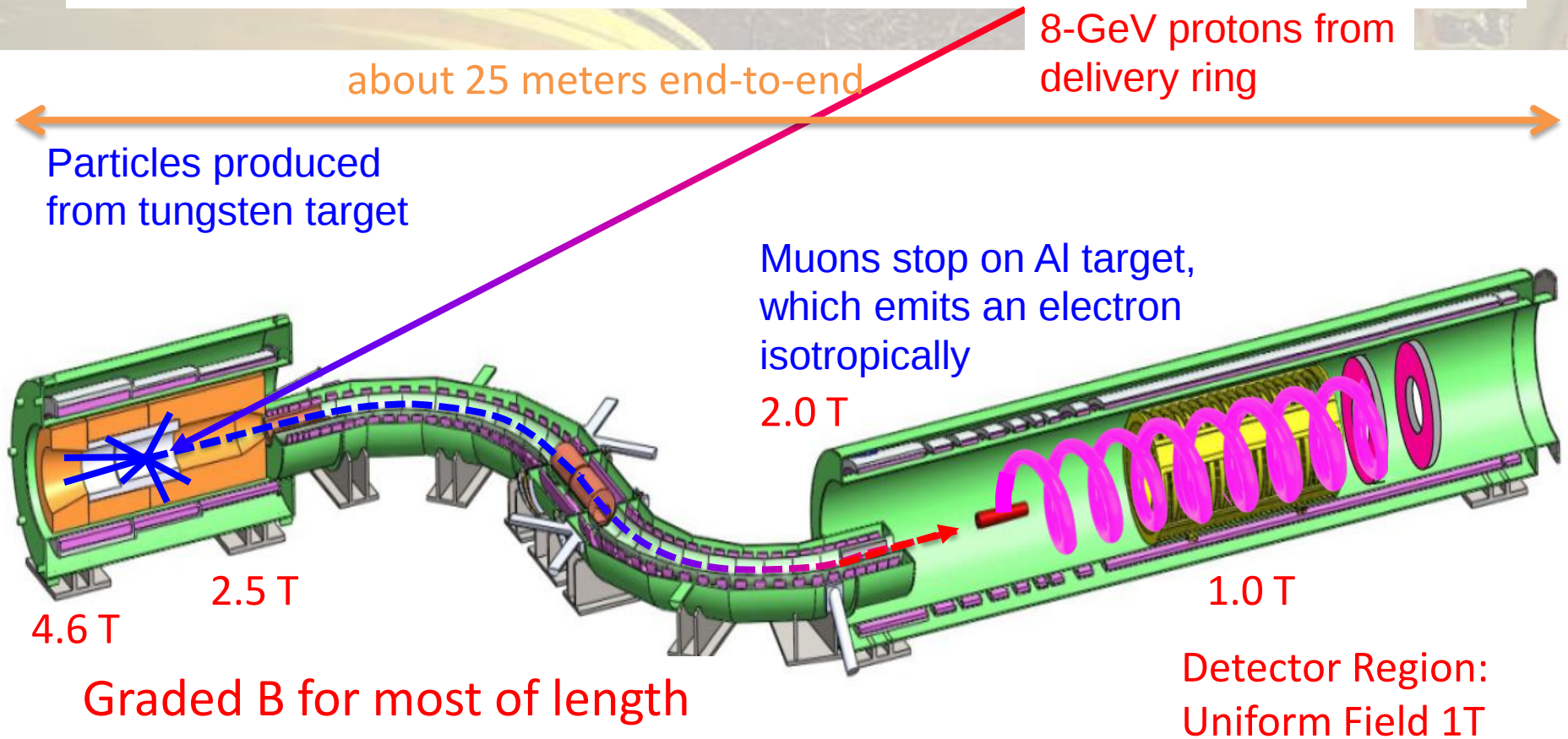
All values preliminary; some are stat error only.

- * scales with extinction: values in table assume extinction = 10^{-10}

Signal Sensitivity for 3 Year Run



Baseline Mu2e Apparatus



S-shaped solenoid:

- transports particles to detector area, and
- allows remaining pions to decay to muons
- collimator selects negatively-charged particles

Tracker/calorimeter detect electron signature

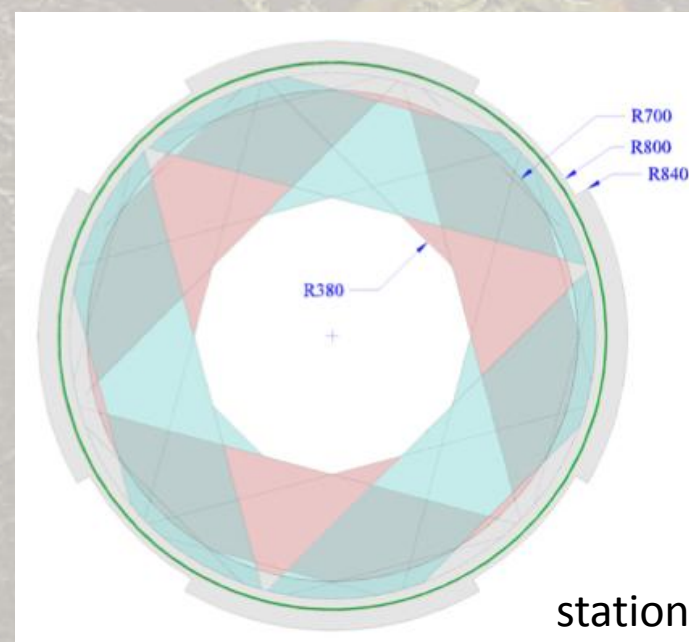
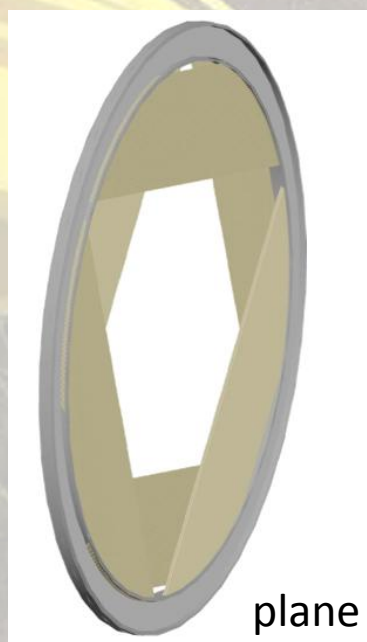
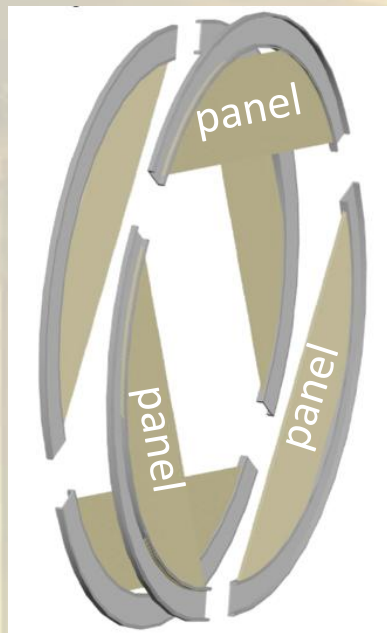
The Mu2e Tracker

- Will employ straw technology
 - Low mass
 - Can reliably operate in vacuum
 - Robust against single-wire failures



- 5 mm diameter straw
- Spiral wound
- Walls: 12 μm Mylar + 3 μm epoxy
 - + 200 $\text{\AA}$ Au + 500 $\text{\AA}$ Al
- 25 μm Au-plated W sense wire
- 33 – 117 cm in length
- 80/20 Ar/CO₂ with HV < 1500 V

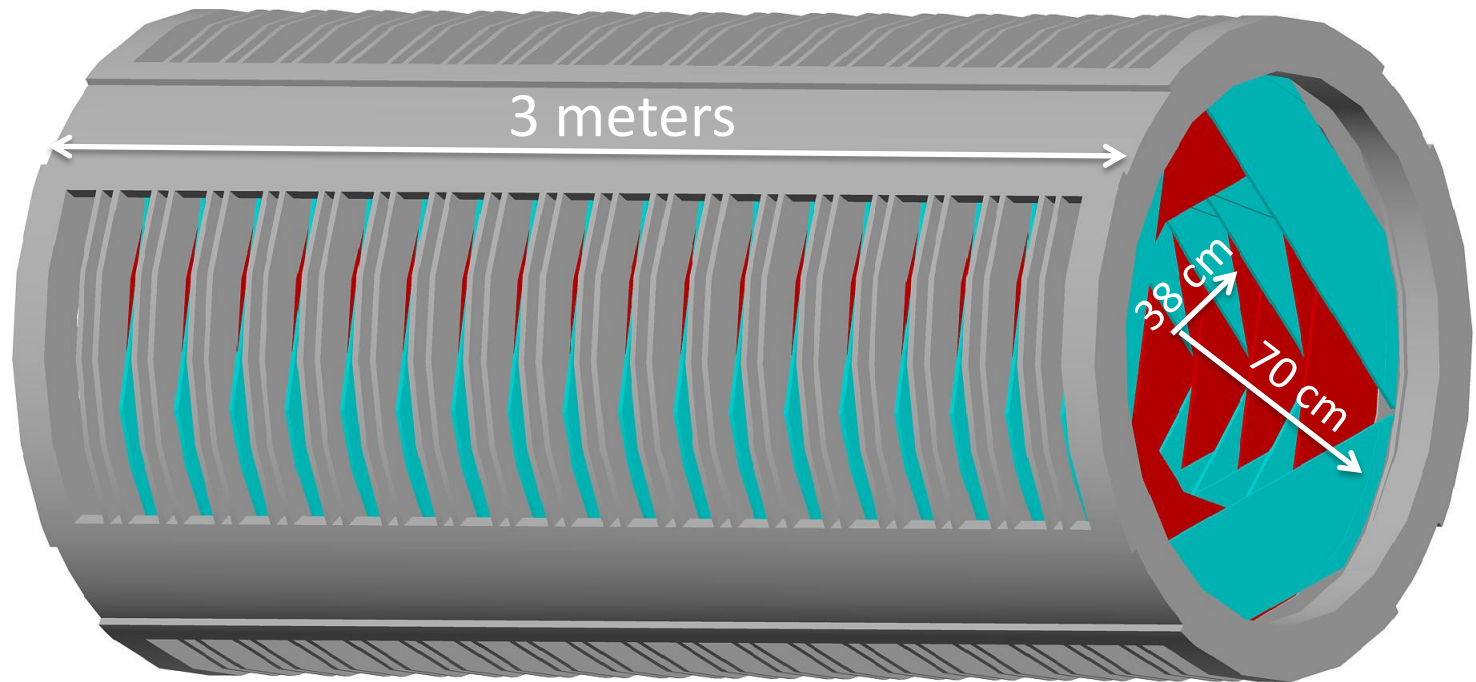
The Mu2e Tracker



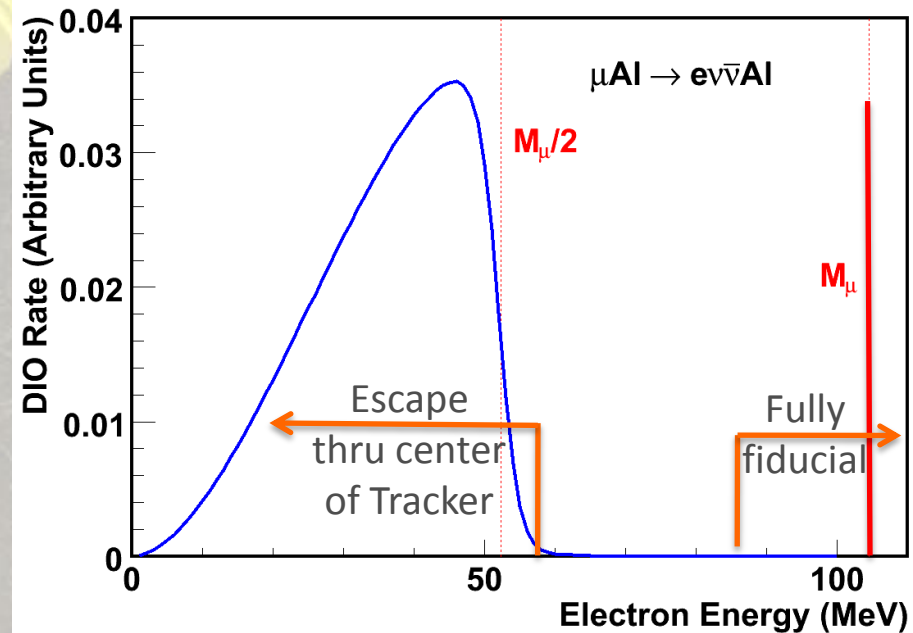
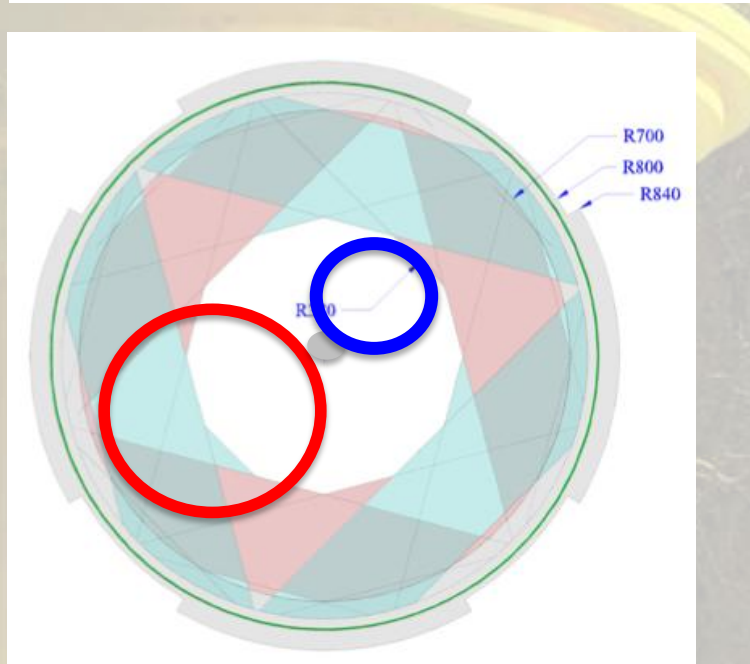
- Self-supporting “panel” consists of 100 straws
- 6 panels assembled to make a “plane”
- 2 planes assembled to make a “station”
- Rotation of panels and planes improves stereo information
- >20k straws total

The Mu2e Tracker

- 18-20 “stations” with straws transverse to beam
- Naturally moves readout and support to large radii, out of active volume

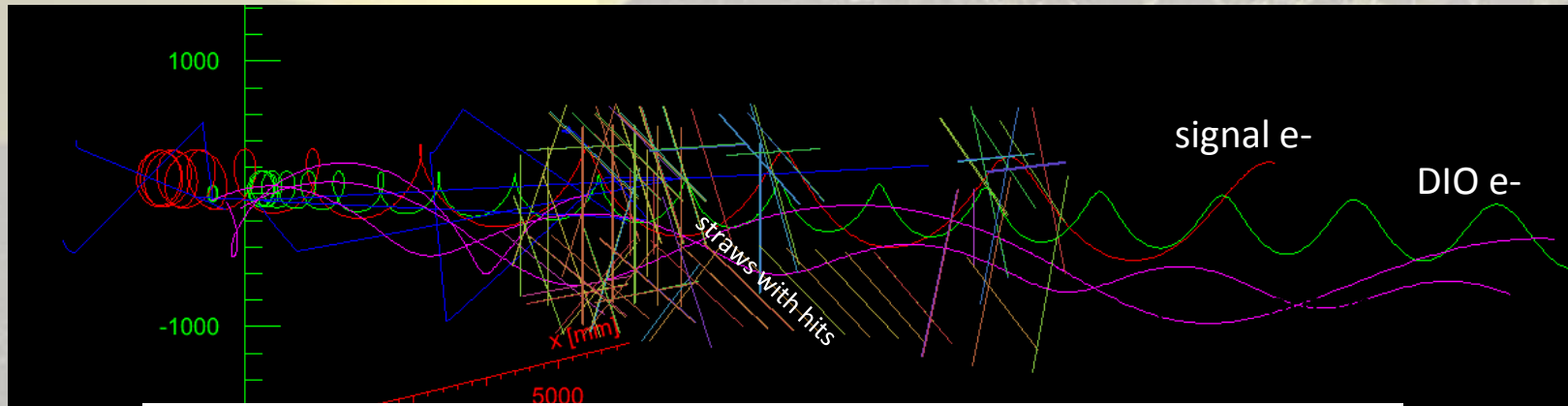


The Mu2e Tracker



- Inner 38 cm is purposefully un-instrumented
 - Blind to beam flash
 - Blind to >99% of DIO spectrum

Mu2e Pattern Recognition



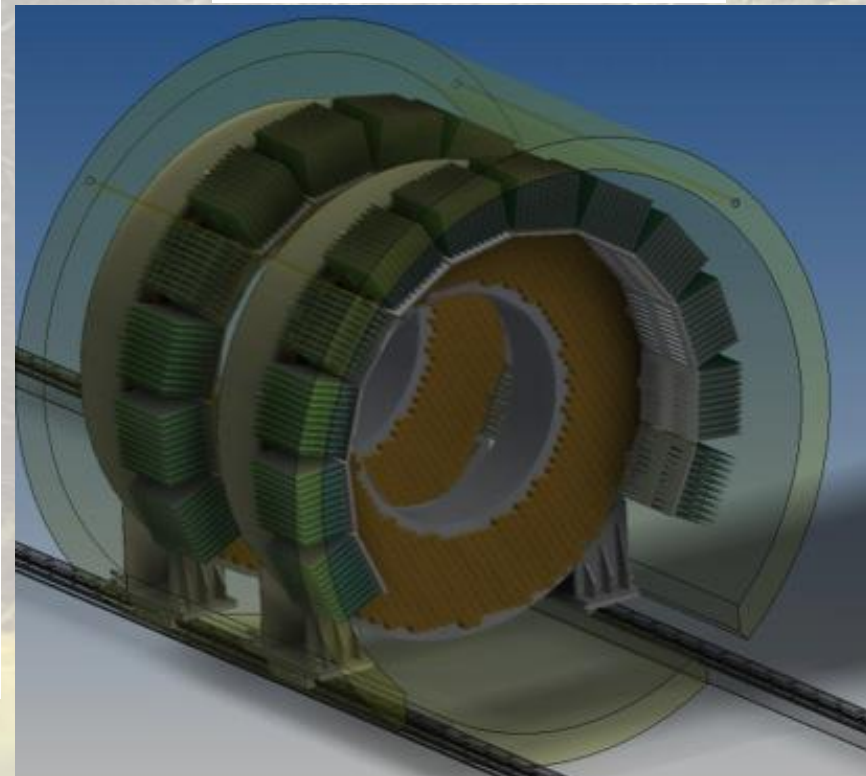
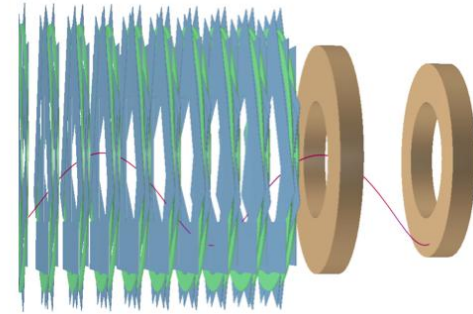
(particles with hits within ± 50 ns of signal electron t_{mean})

- We use timing information to look in ± 50 ns windows – significant reduction in occupancy and significant simplification for Patt. Rec.

Mu2e Calorimeter

The Calorimeter consists of two disks with 1650 BaF₂ hexagonal crystals (30 mm x 200 mm):

- $R_{\text{inner}} = 351 \text{ mm}$, $R_{\text{outer}} = 660 \text{ mm}$, depth = $10 X_0$ (200 mm)
- The distance between disks is optimized at $\frac{1}{2}$ wavelength (70 cm)
- Each crystal is readout by two large area APD's (9x9 mm²) (3300 total)
- Analog FEE and digital electronics are located in near-by electronics crates
- Radioactive source and laser systems provide absolute calibration as well as fast and reliable monitoring capability



Mu2e Calorimeter

	LSO:Ce/LYSO:Ce	BaF ₂	CsI
Density (g/cm ³)	7.40	4.89	4.51
Melting point (°C)	2050	1280	621
Radiation Length (cm)	1.14	2.03	1.86
Molière Radius (cm)	2.07	3.10	3.57
Interaction Length (cm)	20.9	30.7	39.3
Z _{eff}	64.8	51.6	54.0
dE/dX (MeV/cm)	9.55	6.52	5.56
Emission Peak ^a (nm)	420	300 220	310
Refractive Index ^b	1.82	1.50	1.95
Relative Light Yield ^a	100	42 4.8	4.2
LY in 1 st ns (photons)	740	960	100
Decay Time ^a (ns)	40	650 0.9	26
d(LY)/dT ^c (%/°C)	-0.2	-1.9 0.1	-1.4

- a. Top line: slow component, bottom line: fast component.
b. At the wavelength of the emission maximum.
c. At room temperature (20°C)

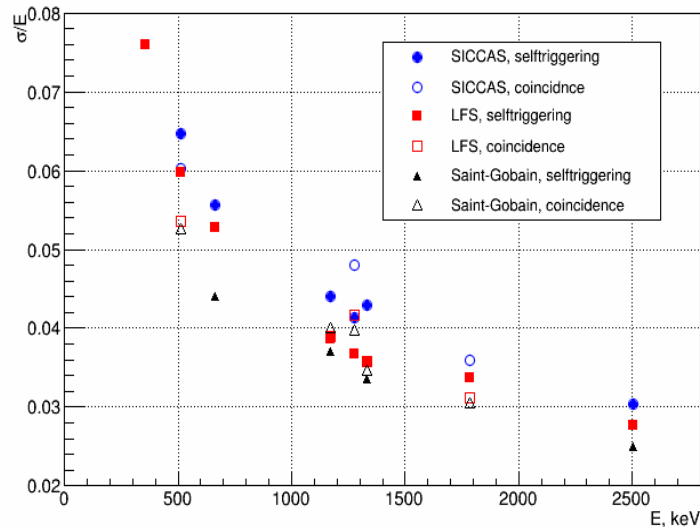
Mu2e Calorimeter

Barium Fluoride (BaF₂)

- Radiation hard, non-hygroscopic
- very fast (220 nm) scintillating light
- Larger slow component at 300 nm. should be suppress for high rate capability
- Photo-sensor should have extended UV sensitivity and be “solar”-blind
- Crystal dimension: hexagonal faces of 33 mm across flats, 200 mm length (10 X₀)

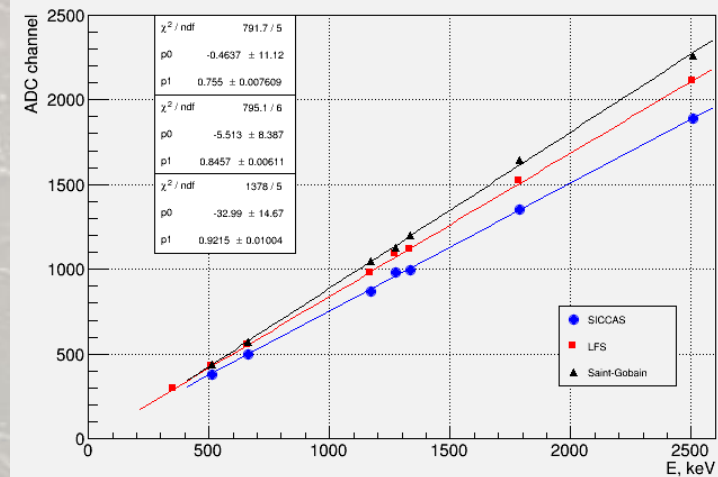
	BaF ₂
Density (g/cm ³)	4.89
Radiation length (cm)	2.03
Moliere Radius (cm)	3.10
Interaction length (cm)	30.7
dE/dX (MeV/cm)	6.52
Refractive index	1.50
Peak luminescence (nm)	220 (300)
Decay time (ns)	1 (650)
Light yield (rel. to NaI)	5% (42%)
Variation with temperature	0.1% (-1.29)% / °C

JINR R&D



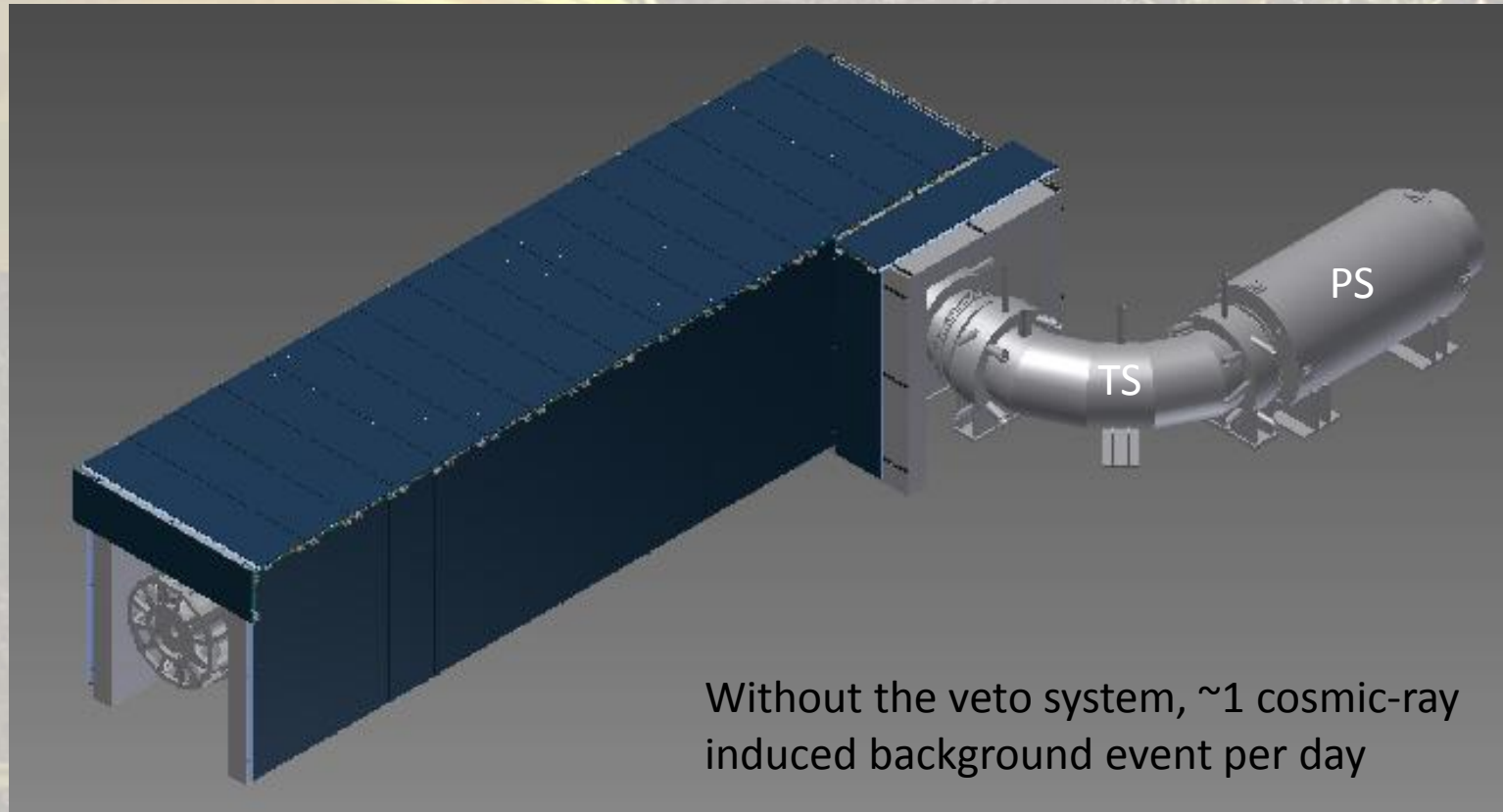
Energy resolutions of LYSO:Ce crystals from Saint-Gobain and SICCAS and LFS crystal from Zecotek.

- R.J. Abrams et al., “Mu2e Conceptual Design Report”, [arXiv:1211.7019](https://arxiv.org/abs/1211.7019) (2012).
- J. Budagov et al., “The calorimeter project for the Mu2e experiment”, Nucl. Instr.&Meth. A718(2013) 56-59.
- O. Sidletskiy et al., “Evaluation of LGSO:Ce scintillator for high energy physics experiments”, Nucl. Instr.&Meth. A735(2014) 620-623.
- K. Afanaciev et al., “Response of LYSO:Ce scintillation crystals to low energy gamma-rays”, JINR preprint E13-2013-141, Dubna, 2013. Submitted to Nucl. Instr.&Meth. A.
- Z. Usubov, “Electromagnetic calorimeter simulation for future $\mu \rightarrow e$ conversion experiments”, [arXiv:1212.4322](https://arxiv.org/abs/1212.4322) (2012).
- Z. Usubov, “Light output simulation of LYSO single crystal”, [arXiv:1305.3010](https://arxiv.org/abs/1305.3010) (2013).



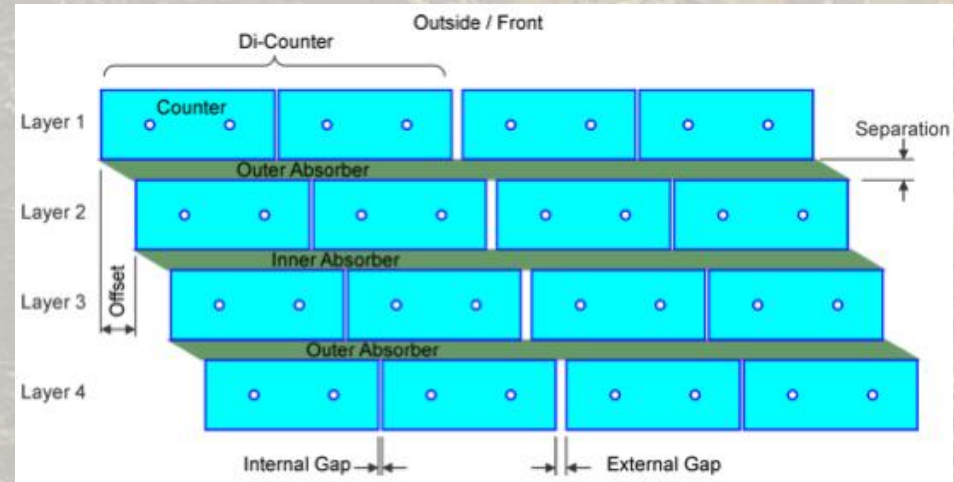
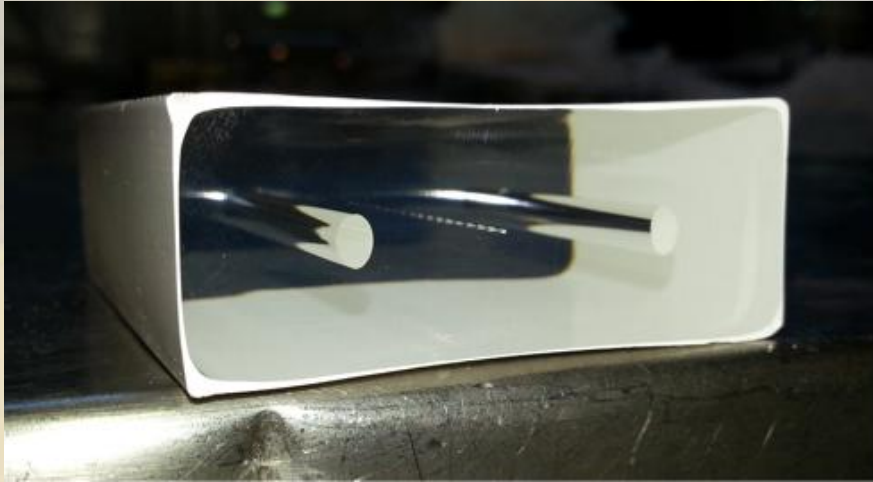
Energy response linearity of the same three crystals in the 511-2500 keV energy range.

Mu2e Cosmic-Ray Veto



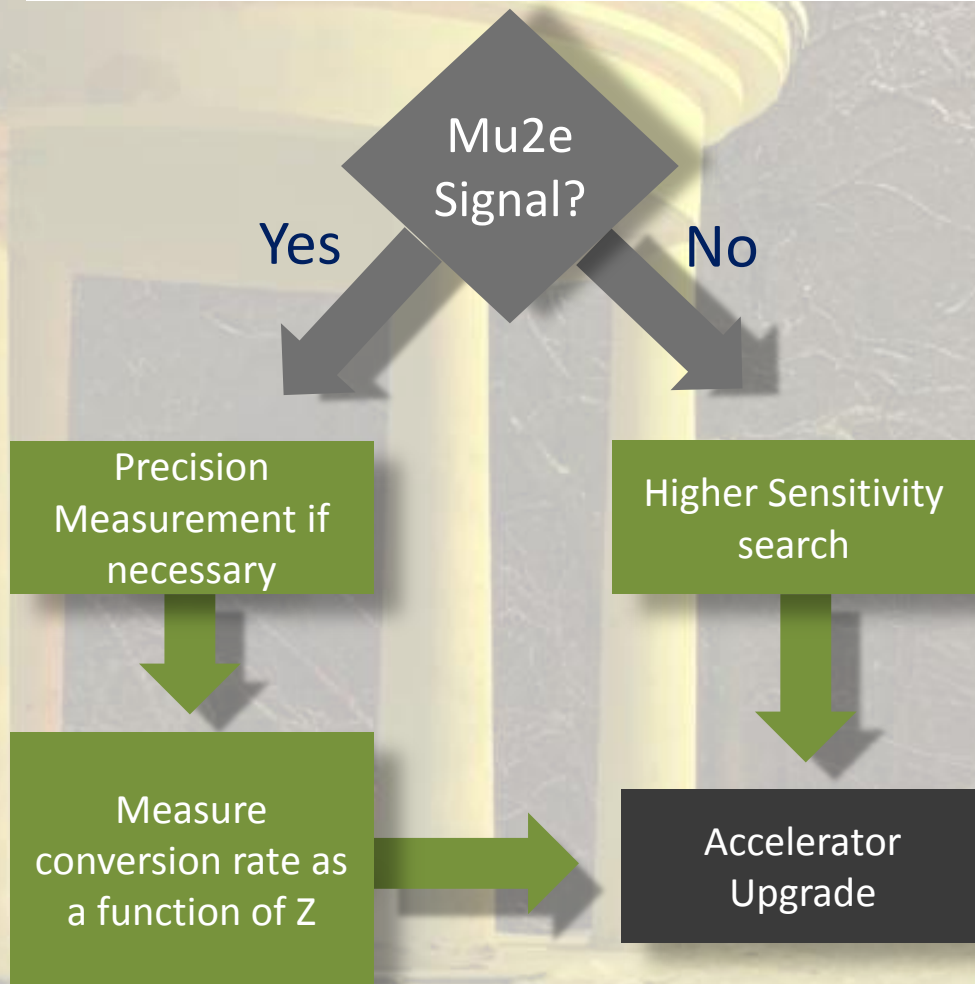
- Veto system covers entire DS and half TS

Mu2e Cosmic-Ray Veto



- Will use 4 overlapping layers of scintillator
 - Each bar is $5 \times 2 \times \sim 450 \text{ cm}^3$
 - 2 WLS fibers / bar
 - Read-out both ends of each fiber with SiPM
 - Have achieved $\varepsilon > 99.4\%$ (per layer) in test beam

What next?



- A next-generation Mu2e experiment makes sense in all scenarios
 - Push sensitivity or
 - Study underlying new physics
 - Will need more protons → upgrade accelerator

Summary

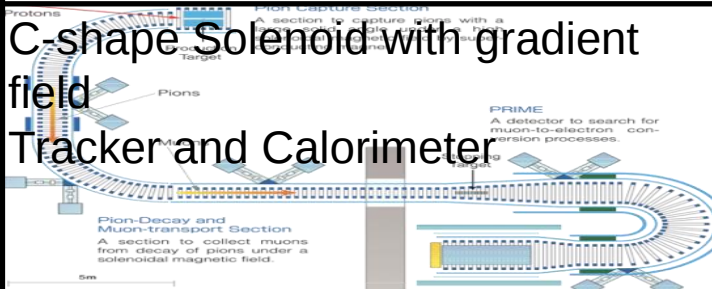
The Mu2e experiment:

- Improves sensitivity by a factor of 10^4
- Provides *discovery capability* over wide range of New Physics models
- Is complementary to LHC, heavy-flavor, and neutrino experiments



BACKUP

Mu2e / COMET comparison

	Mu2e	COMET
approval/ funding	ranked among the very top priorities for the U.S. HEP program P5 Report 2014 “Complete the Mu2e and g-2 projects.” Mu2e is <i>fully funded in all budget scenarios</i> The DOE is committed to completing Mu2e	COMET phase-II funding has not been identified and Japan has clearly stated that their top priorities are Belle-II, long baseline neutrinos, and ILC
Operation Condition	Mu2e will run simultaneously with NOvA and the short-baseline neutrino program at Fermilab	COMET cannot run simultaneously with the JPARC neutrino program. It forces COMET to plan for higher beam power in order to minimize the amount of required beam time
Detector	Straight Solenoid with gradient field Tracker and Calorimeter 	C-shape Solenoid with gradient field Tracker and Calorimeter 

Mu2e is equally sensitive to e^- and e^+ and thus have an additional physics channel $\mu^- N \rightarrow e^+ N'$ and will be able to measure in situ some background components (e.g. $\pi^- N \rightarrow \gamma N' \rightarrow e^- e^+ N'$)

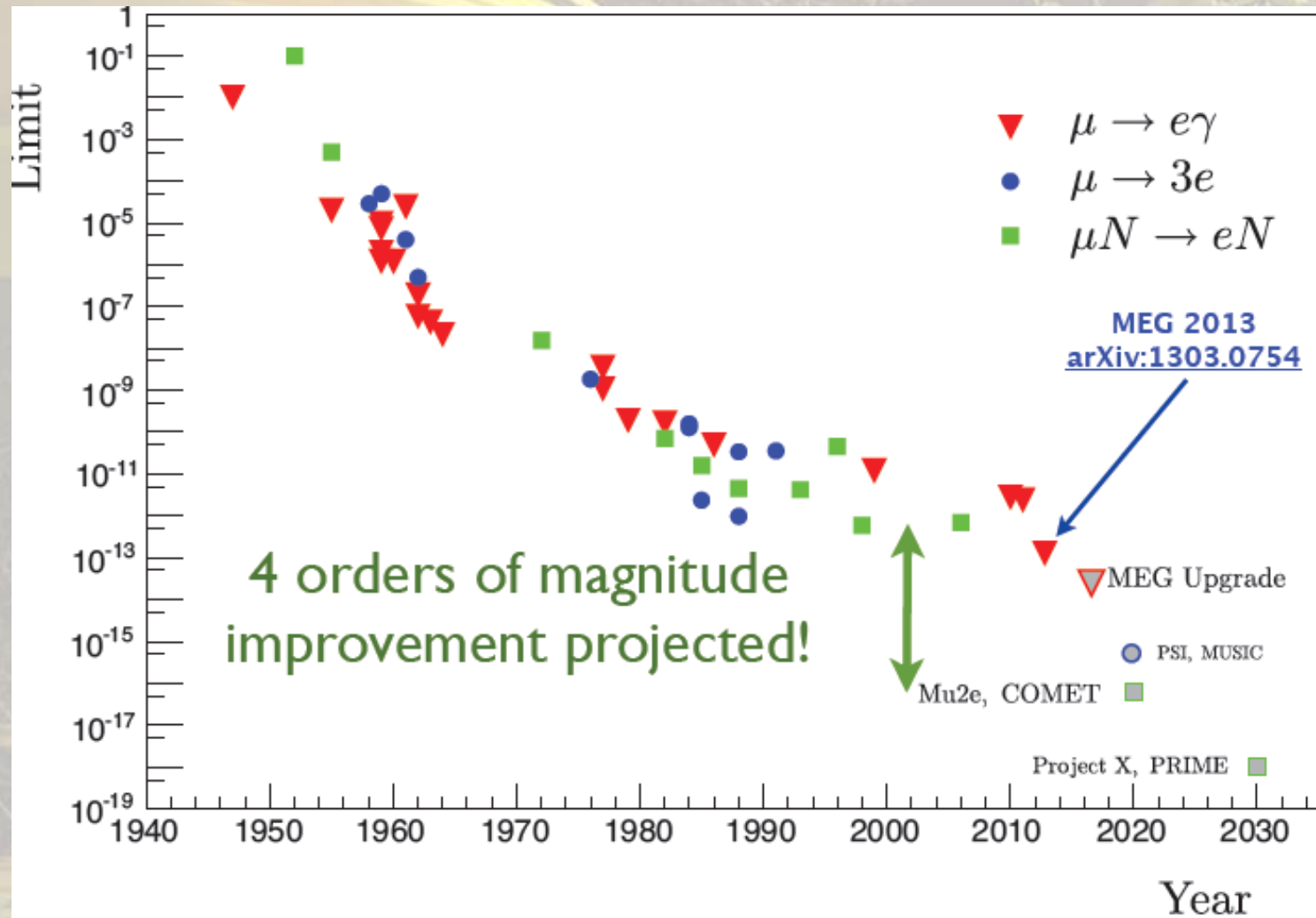
the COMET solenoids, particularly their production solenoid, is more technically risky (e.g. they will use a 9 (!) layer coil, Mu2e is 3 layer).

Comparison of COMET Phase-I / Phase-II and Mu2e

90% C.L. upper limit is 7×10^{-13} (SINDRUM)

	S.E. sensitivity	BG events at aimed sensitivity	running time (sec)	Year	Comments
COMET Phase-I	3×10^{-15}	0.03	1.5×10^6	~2016	Proposal (2012)
COMET Phase-II	3×10^{-17}	0.34	2×10^7	~2019	CDR (2009)
Mu2e	3×10^{-17}	0.4	3x (2×10^7)	~2019	J. Miller's talk at SSP2012

Status of CLFV Searches



Mu2e / COMET comparison

- Mu2e employs Booster batches left unused by the Fermilab neutrino program
 - **Mu2e will run simultaneously with NOvA and the short-baseline neutrino program at Fermilab**
 - All these program can simultaneously get the protons they need to meet their physics goals
 - **Mu2e can run at lower beam power**
 - **Saves significant money, reduces detector rates, simplifies solenoids, strengthens physics program**
- **COMET cannot run simultaneously with the JPARC neutrino program**
 - Either one or the other can take data
 - **Forces COMET to plan for higher beam power in order to minimize the amount of required beam time**
 - More complicated solenoid system, need to mitigate detector rates with a C-shaped detector solenoid, which significantly reduces their acceptance for e+

Mu2e and Muon g-2 work together

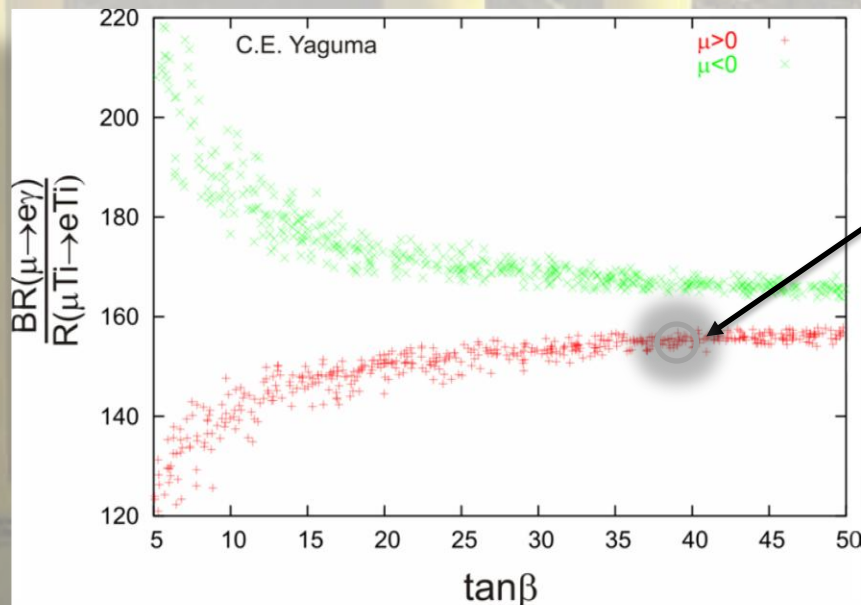
Example:

SUSY contributes to $a_\mu = (g-2)/2$

- From g-2 we know $\tan \beta$
- From g-2 we know also know $\mu > 0$
- From Mu2e we measure $R(\mu N \rightarrow e N)$ and take the ratio to the MEG result

$$a_\mu^{SUSY} \approx 130 \times 10^{-11} \left(\frac{100 \text{ GeV}}{M_{SUSY}} \right)^2 \tan \beta \text{ sign}(\mu)$$

We use this match to prediction as a way to disentangle, or validate, or interpret manifestations of SUSY



g-2 selects which curve we should be on, and gives us the value of $\tan \beta$

$\tan \beta$ - the ratio of the vacuum expectation values of the two Higgs doublets
 $\text{Sign}(\mu)$ - the sign of the higgsino mass parameter

Mu2e and Muon g-2 work together

Gives us access to
Slepton Mixing
Matrix elements

Slepton Mixing Matrix

$$\begin{pmatrix} m_{\tilde{e}\tilde{e}}^2 & \Delta m_{\tilde{e}\tilde{\mu}}^2 & \Delta m_{\tilde{e}\tilde{\tau}}^2 \\ \Delta m_{\tilde{e}\tilde{\mu}}^2 & m_{\tilde{\mu}\tilde{\mu}}^2 & \Delta m_{\tilde{\mu}\tilde{\tau}}^2 \\ \Delta m_{\tilde{\tau}\tilde{e}}^2 & \Delta m_{\tilde{\tau}\tilde{\mu}}^2 & m_{\tilde{\tau}\tilde{\tau}}^2 \end{pmatrix}$$

