



High Precision Measurement of Energy Spectra of CR Species with LHAASO

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For LHAASO collaboration

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Outline

- LHAASO introduction
- LHAASO science in cosmic ray measurement
- The measurements of knees of cosmic rays for individual species.
- Summary





Large High Altitude Air Shower Observatory



KM2A: • 5195 EDs: 1 m² each, 15m spacing • 1171 MDs: 1146 MDs, 36 m² each, 30m spacing



WFCTA: 18 Cherenkov telescopes (1024 pixels/telescope)

WCDA: three pools



Daochen, Sichuan, China (29°21' 31" N, 100°08'15" E, 4410 m a.s.l., 600 g/cm)



Hybrid Measurements of Showers 5m 7m3200 cells WCDA WFCTA 1m 150 m See and

LHAASO science in cosmic ray spectra measurment

Measure individual cosmic ray spectra from 10TeV to EeV

• Multi-stages, Multi-parameters



3. 10PeV-100PeV: knee for iron



- 1. 10TeV-100TeV: energy scale;
- 2. 100TeV-10PeV: H, He knees



4. >100PeV: second knee





Prospects of P, He knees from 100TeV to 10PeV

- ¼ LHAASO array is expected to be operated by the end of this year
 - 6 WFCT telescopes
 - WCDA++, namely the pool with higher dynamic range.
 - 300 muon detectors







Measurement and Reconstruction

WCDA

- Energy flux near the core
- Core reconstruction: 3m
- Arrival direction reconstruction: 0.3°

WFCTA

- SIZE (total PE in image)
- Width, Length
- Angular offset between arrival directions to the image center
- KM2A MD







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Particle Identification





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Particle identification Results

¼ LHAASO



	Proton	Proton+Helium
Aperture	~900	~1800
Contamination	~10%	<4%





Prospects for cosmic ray spectra

- One year statistics, 10% duty cycle
- **6** Cherenkov Telescopes



Horandel model

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Prospects of iron knee from 10PeV to 100PeV

• KM2A:

- Geometry reconstruction
 - 3m, 0.3°
- No. of Muons
- No. of electromagnetic particles

• WFCTA:

- Px: angular offset between shower arrival direction and the image center Xmax
- Size: Number of pe. in the Cherenkov image

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Selection criteria:

1: The cores should be in the array of KM2A

2: The arrival diections should be in the FOV of WFCTA ($|Y| < 7^{\circ}$)

3: image completeness N_{trigger}>100



with full-moon duty cycle >30%



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Lateral distributions of Log10(N_{tot}/N_{μ})



Normalized to 1









The correlation is less then 0.9







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Energy Resolution









Summary

The knees of proton and proton + helium spectra can be obtained by 6 CTs, MD, and the first pool of WCDA.

The knees of iron or iron + MgAlSi spectra can be obtained by 18 CTs (45 in elevation)+ ED+ MD array

 With high statistics , high purity energy resolution is better than 20%.





Thank you !





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Water Cherenkov Detector Array

3 water ponds:

- 78,000 m² in total;
- 4 m effective depth;
- 3120 cells, with an 8"/9" PMT in each cell;
- Cells are partitioned with black curtains.
- Detect Cherenkov photons produced by secondary particles in the water.
- WCDA++: 1"PMTs enhance dynamic range







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KiloMeters² Array







Wide Field of View Cherenkov Telescope Array

- 5m² spherical mirror
- ◆ 32×32 SiPMs array
- Pixel size 0.5 $^{\circ}$
- \bullet FOV: 16 $^{\circ}$ imes 16 $^{\circ}$
- Detect the Cerenkov and fluorescent composition in EAS
- Duty cycle: 30%

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Sub-Cluster imes 64 ightarrow





