

## **Review of the results of the space experiment NUCLEON**

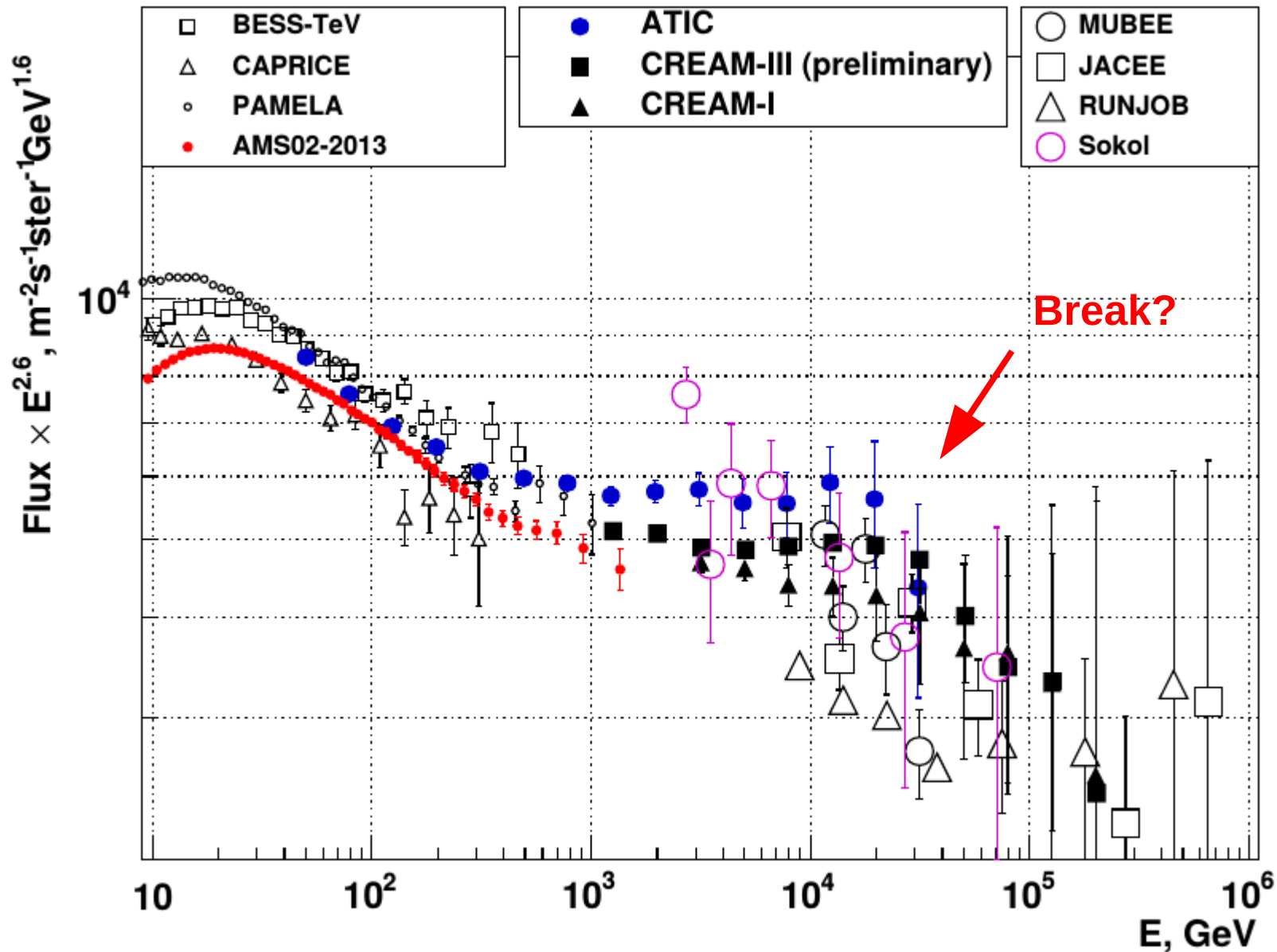


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**WASDHA2018**

# Approaching the knee area.

Problems in cosmic-ray spectra at energies 10 TeV - 1PeV per particle.  
An example: proton spectra before NUCLEON.

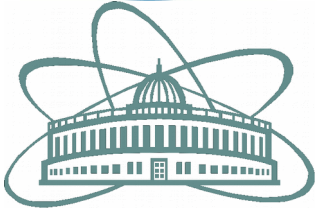


# **The objectives of NUCLEON space experiment**

**Priority: experimental study of cosmic ray spectra in the energy range from a few TeV - 1 PeV per particle with elemental charge resolution.**



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***Joint Institute for Nuclear Research, Dubna, Russia***



***National Research Nuclear University "MEPhI", Moscow***



***SDB Automatika, Ekaterinburg, Russia***



***ROSKOSMOS, Russia***



***JSC SRC "Progress"***



***Russian Academy of Sciences***

# NUCLEON mission

NUCLEON apparatus is placed on board of the **RESURS-P** regular satellite as an additional payload.

Lunched **December 28, 2014.**

The spacecraft orbit is a Sun-synchronous one with inclination  **$97.276^\circ$**  and an average altitude of **475 km.**

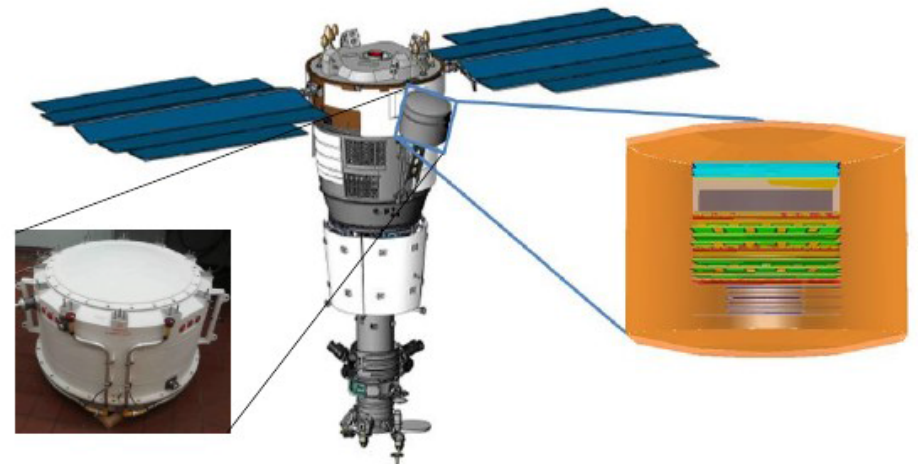
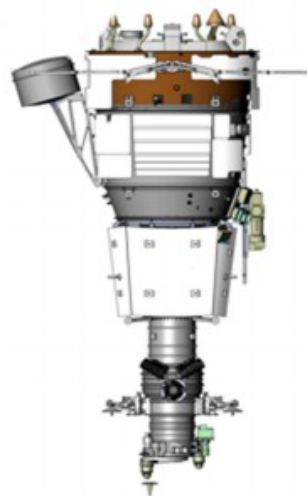


## Vessel:

Weight ~360 kg

Power consumption  
~160 W

Telemetry ~10 GB/day



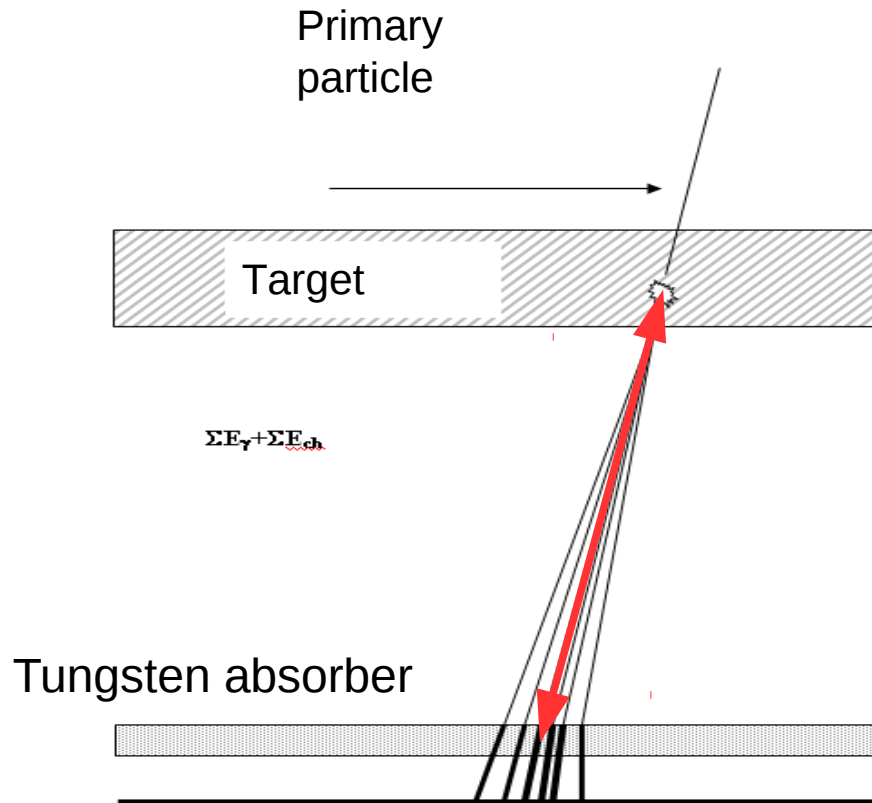
The NUCLEON detector on board of the satellite RESURS-P N2.

## IMPORTANT FEATURE OF THE EXPERIMENT:

Two different methods of measuring of the energy of particles are implemented in the NUCLEON experiment:

1. The kinematic method **KLEM**  
(Kinematic Lightweight Energy Meter)  
-for the first time (**main**)
2. The calorimetric method  
-usual and well studied

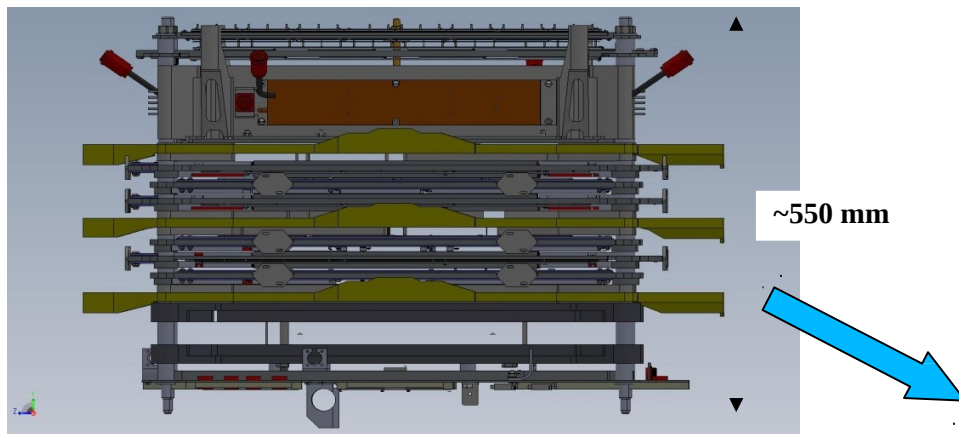
# The kinematic method KLEM + calorimeter



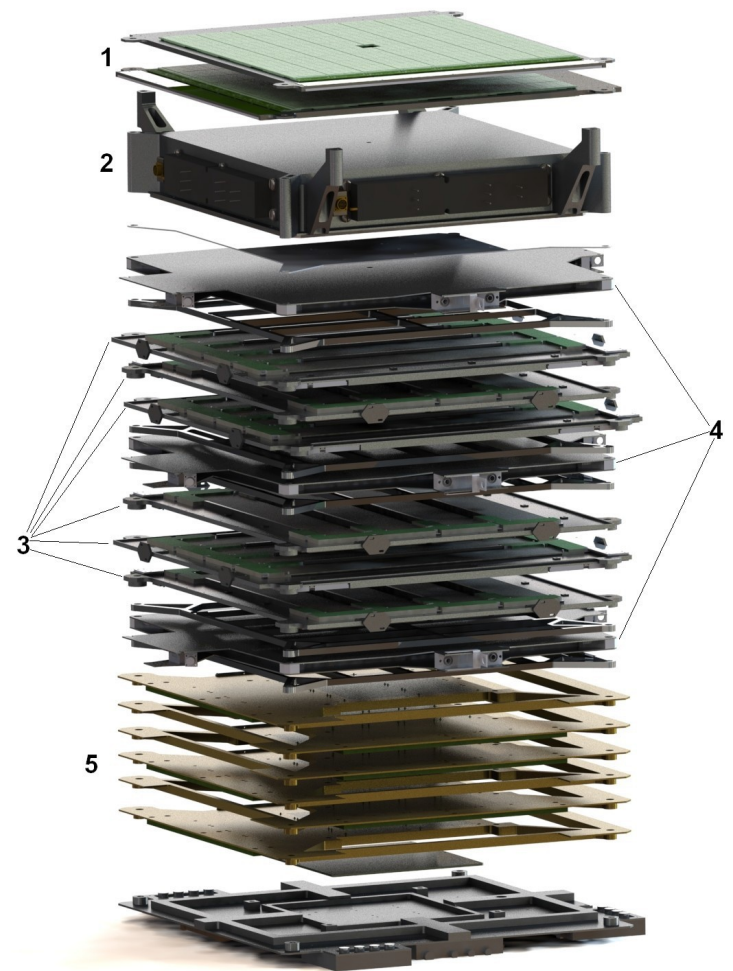
The number of secondary particles with high pseudorapidity after the first interaction increases logarithmically along increasing of the primary energy of particle

The energies are reconstructed by S-parameter -  
$$S = \Sigma(I_i * \ln^2(2H/x_i))$$

# The NUCLEON apparatus

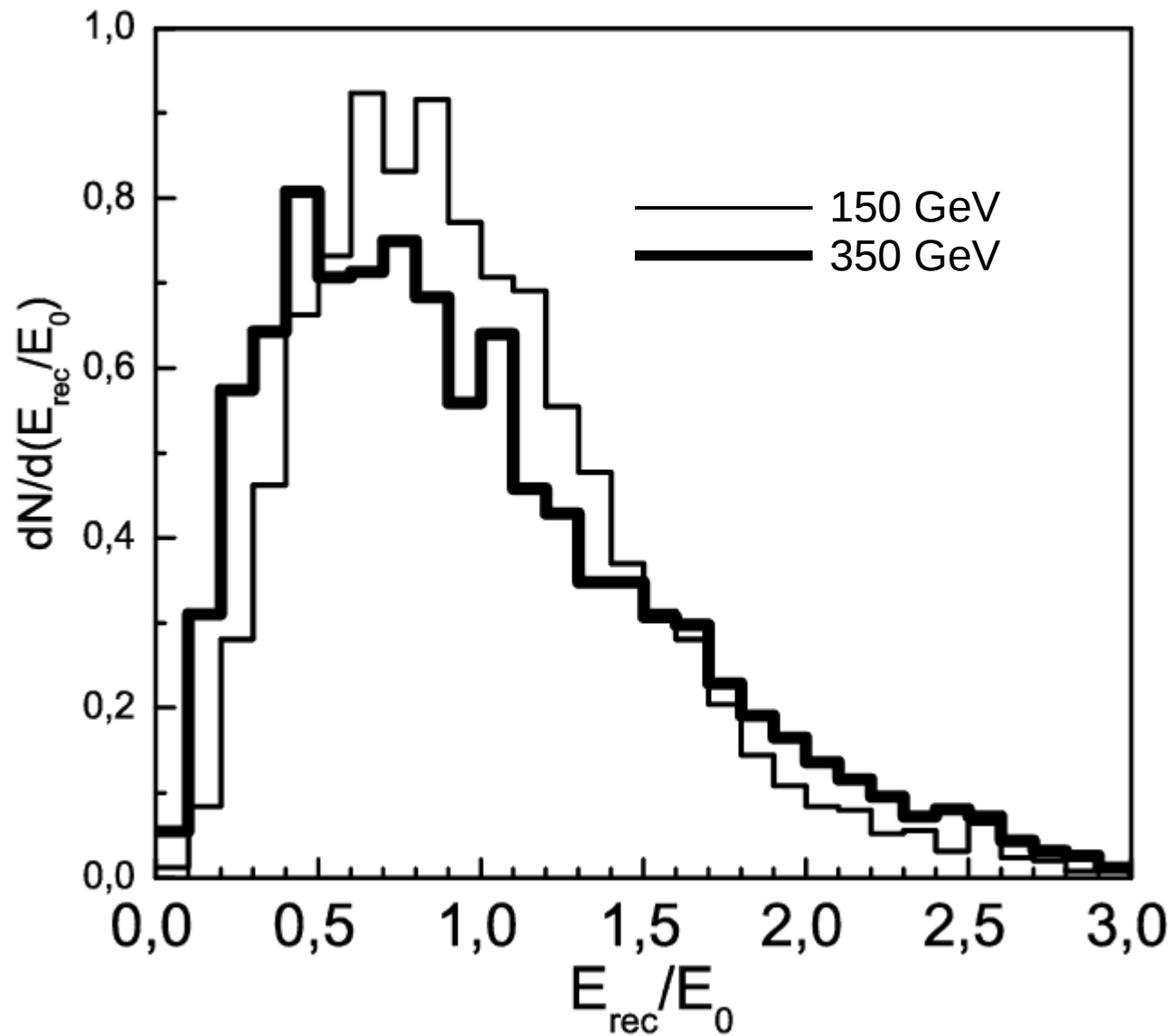


- (1) System of charge measurements – four planes of pad silicon detectors ( $1.5 \times 1.5 \text{ cm}^2$ );
- (2) Carbon target of 0.25 proton interaction lengths;
- (3) KLEM tracker – six planes of microstrip silicon detectors (0.4mm step) with tungsten between them ( $\sim 2 \text{ mm}$  each,  $\sim 3$  X-lengths summary).  
Active square  $500 \times 500 \text{ mm}^2$ .  
KLEM geometrical factor  $0.24 \text{ m}^2 \text{ sr}$ .
- (4) Trigger system – tree double scintillator planes;
- (5) Ionization calorimeter (IC) – six planes of tungsten absorber ( $\sim 8 \text{ mm}$  each,  $\sim 12$  X-lengths summary) with silicon strip detectors (1mm step).  
Active square  $250 \times 250 \text{ mm}^2$ .  
IC geometrical factor  $0.06 \text{ m}^2 \text{ sr}$

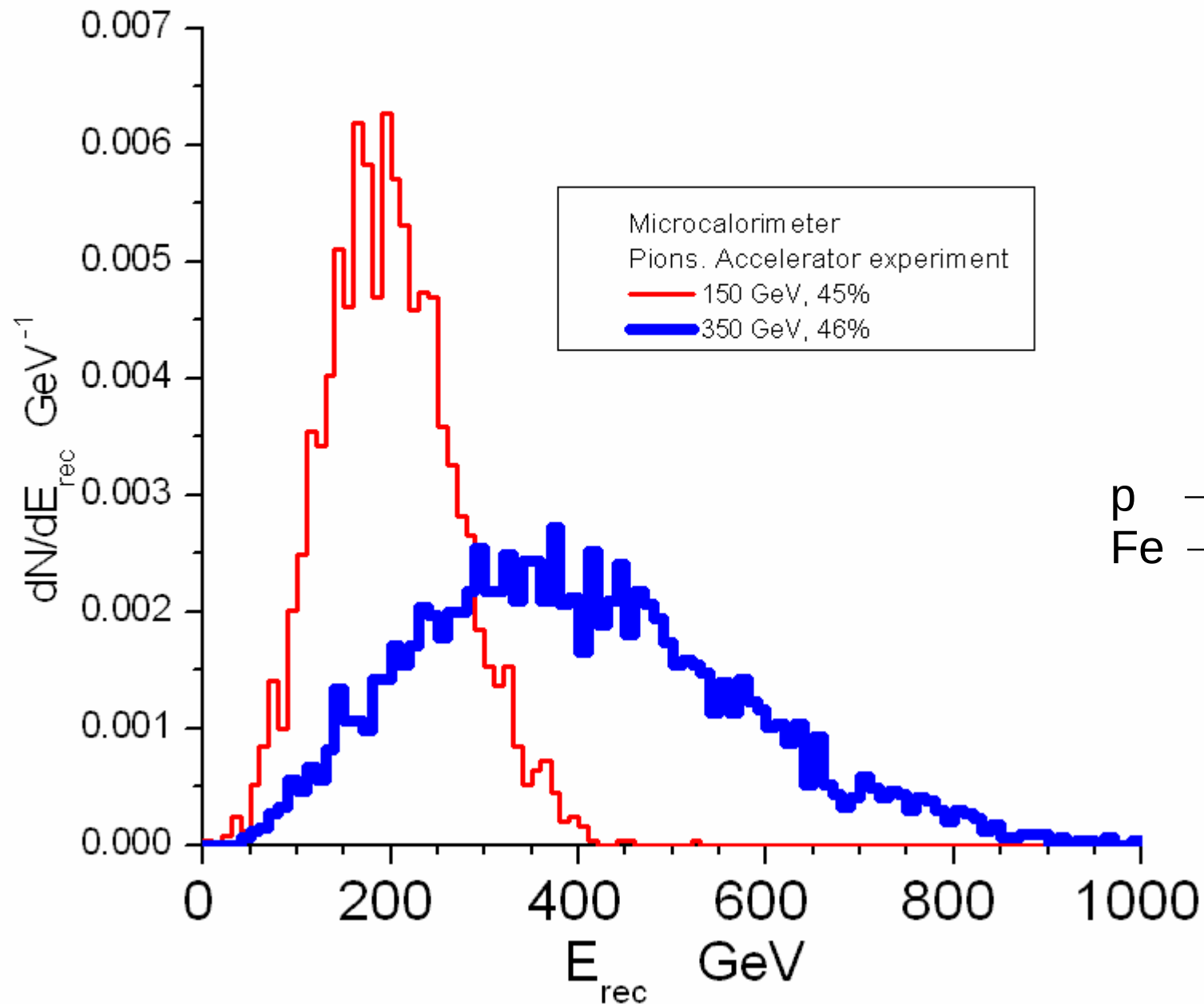


10604 independent  
electronic channels in total

## KLEM, CERN test for $\pi^-$ , resolution $\sim 60\%$



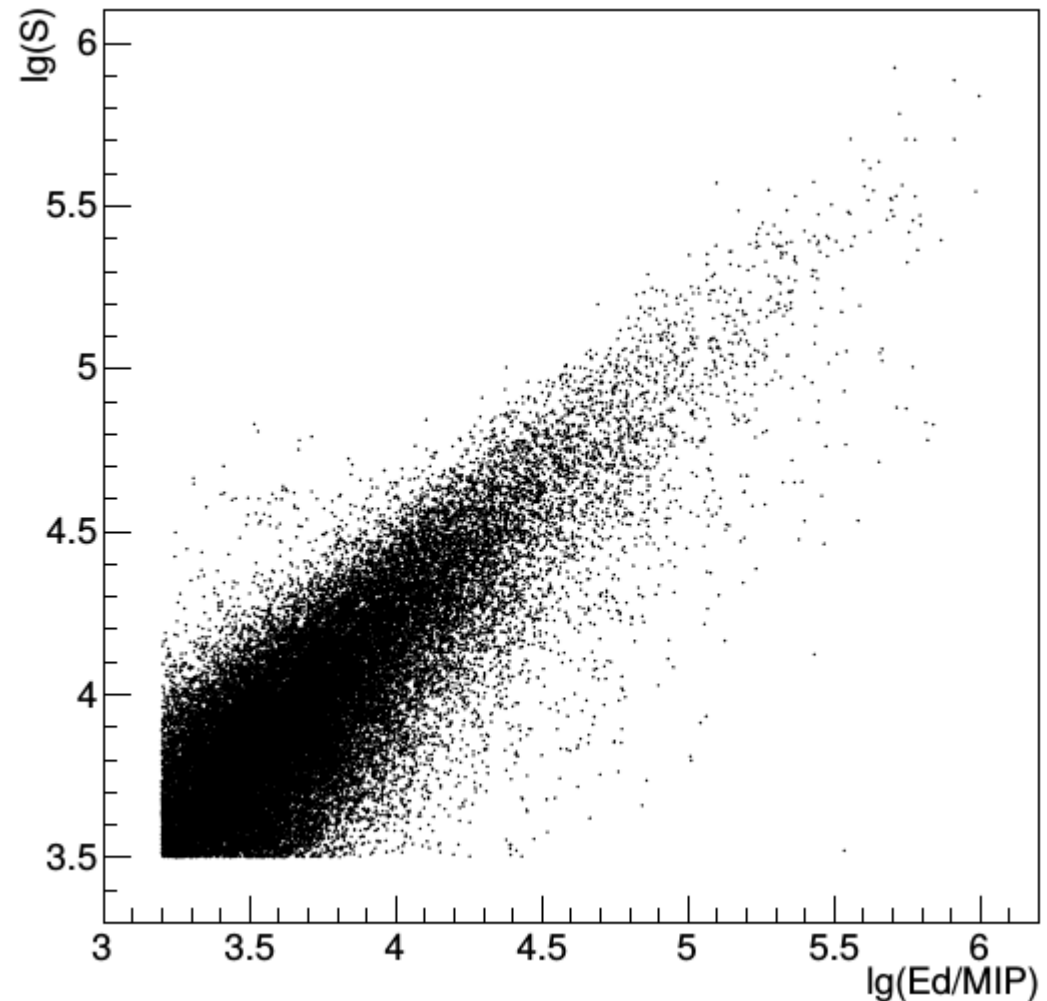
# Calorimeter, CERN test for $\pi^-$ , resolution $\sim 45\%$



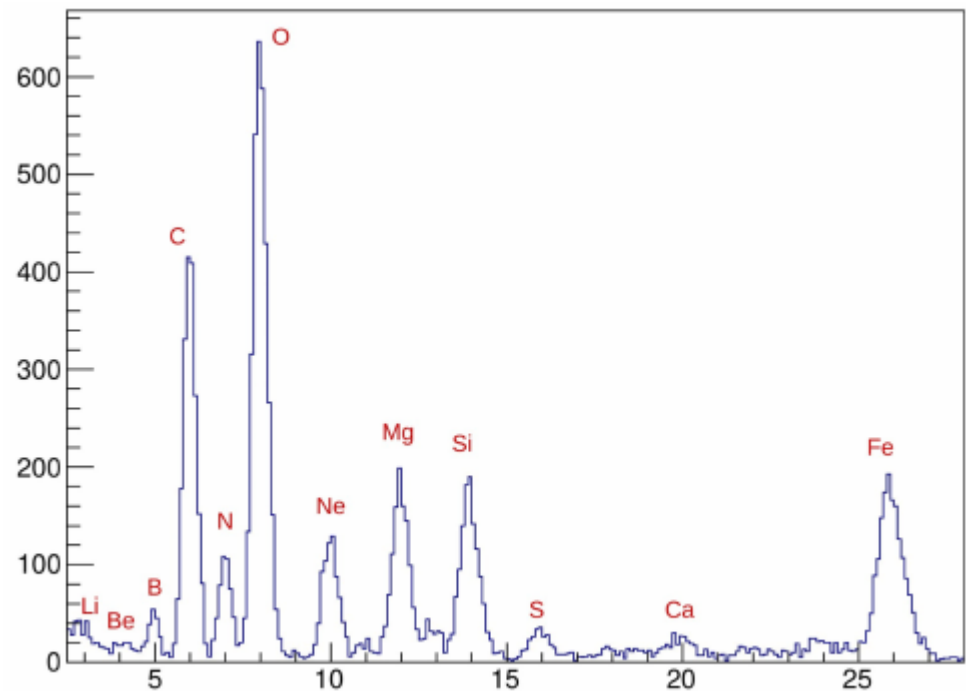
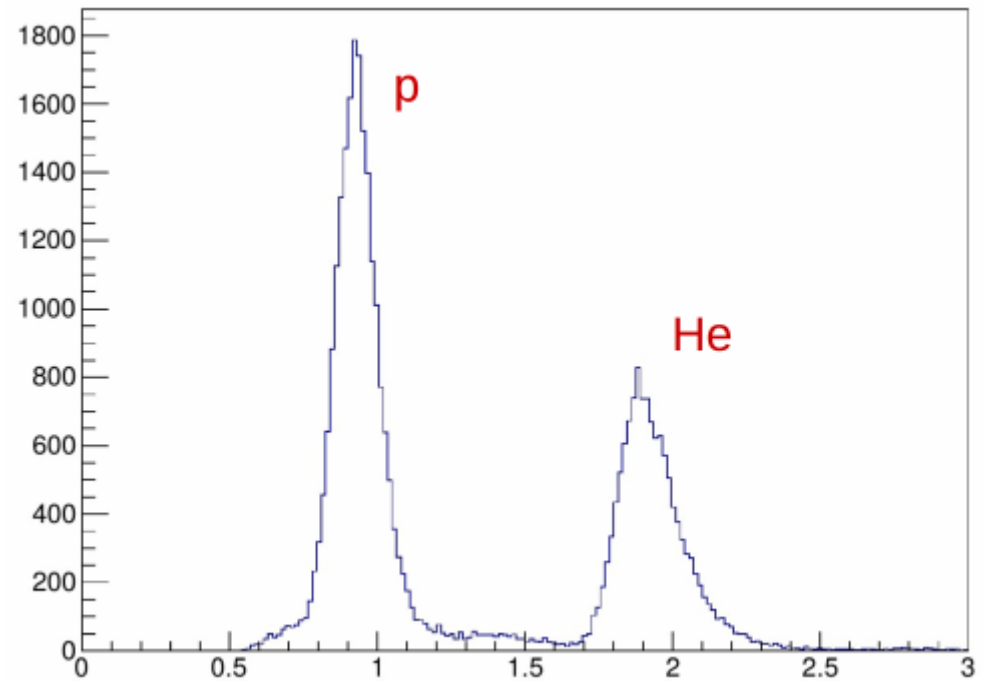
# Model-independent evidence that the KLEM method may be used to measure energy of particles

Correlation of the  
calorimeter energy  
deposit ( $E_d$ )  
and KLEM  
parameter ( $S$ )  
~90%

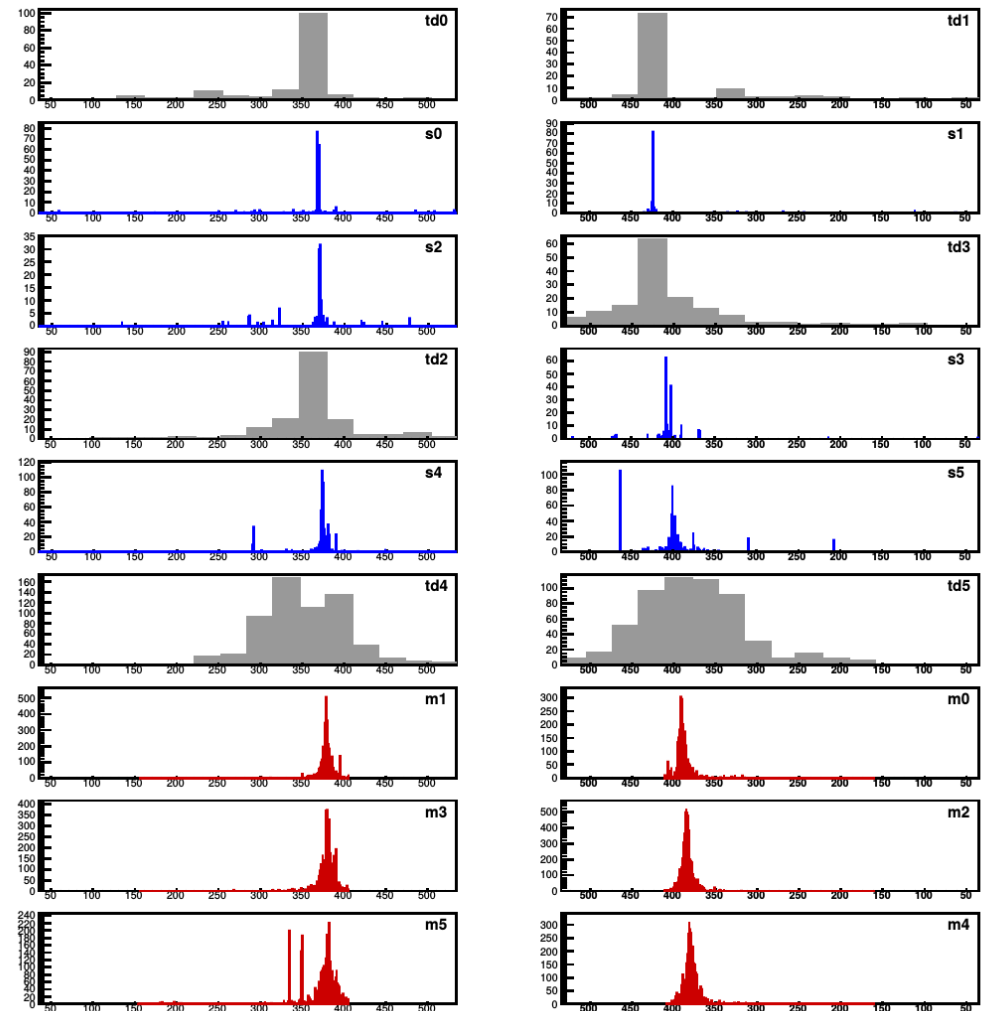
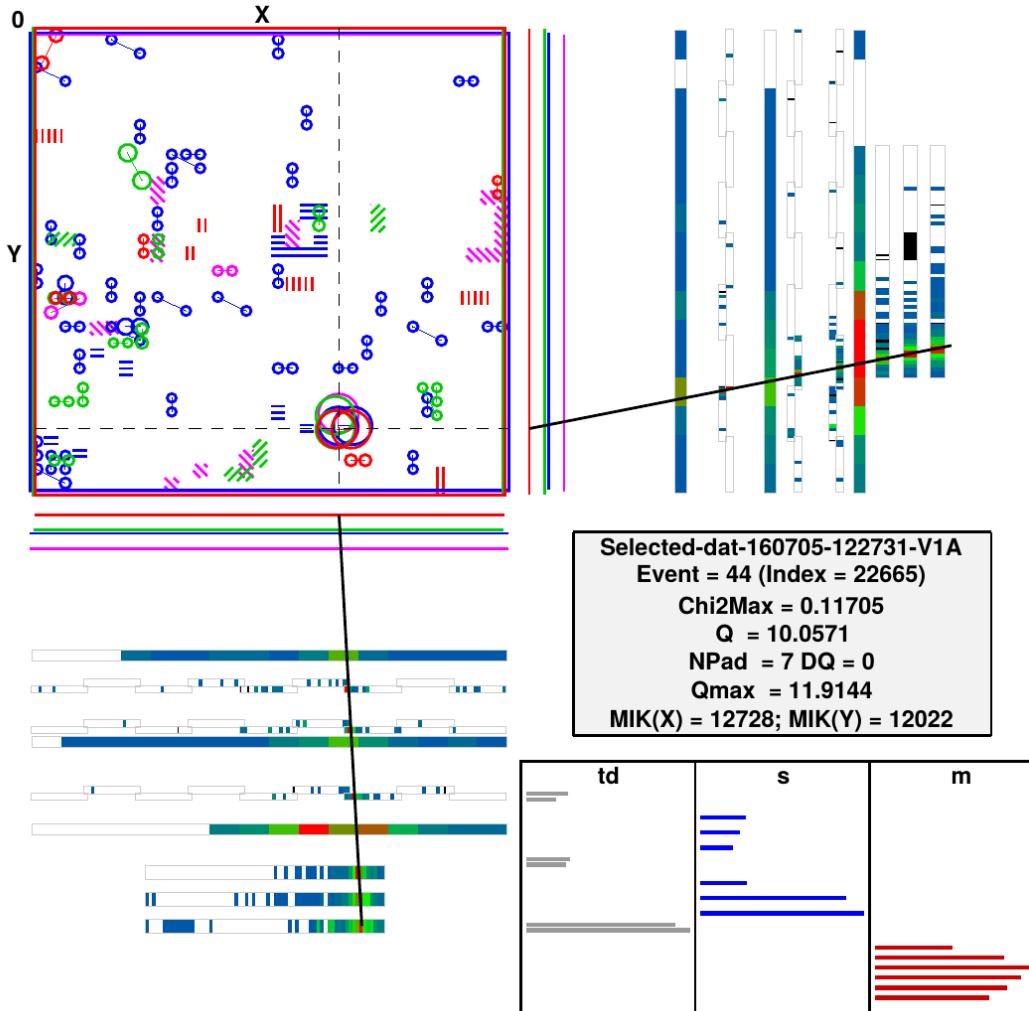
This correlation is a  
model-independent  
result



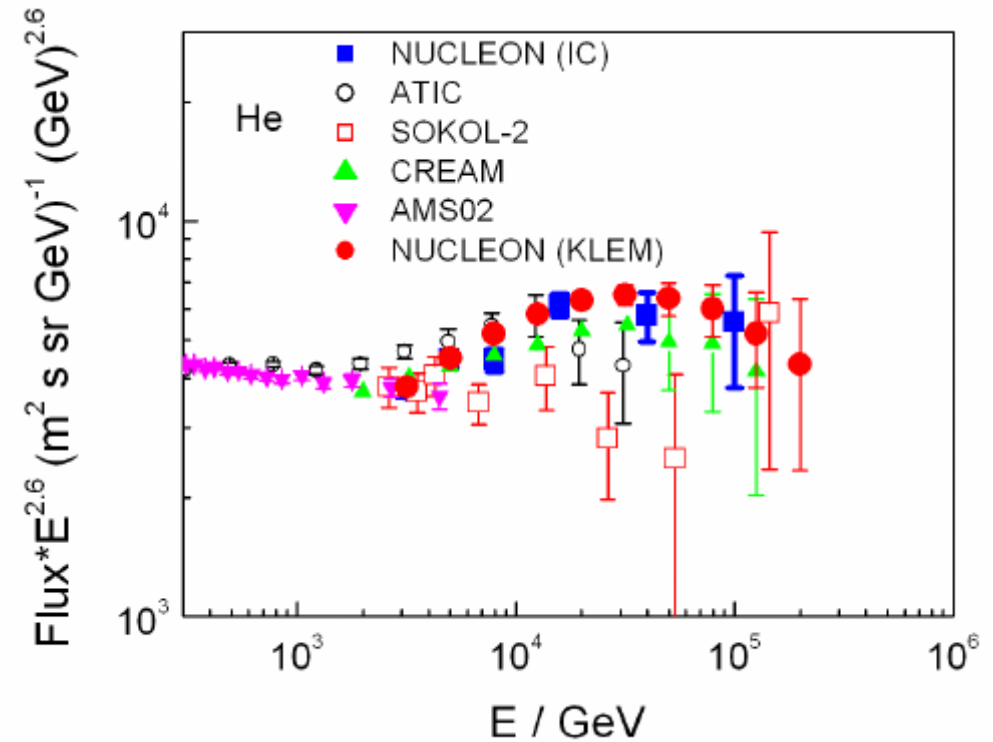
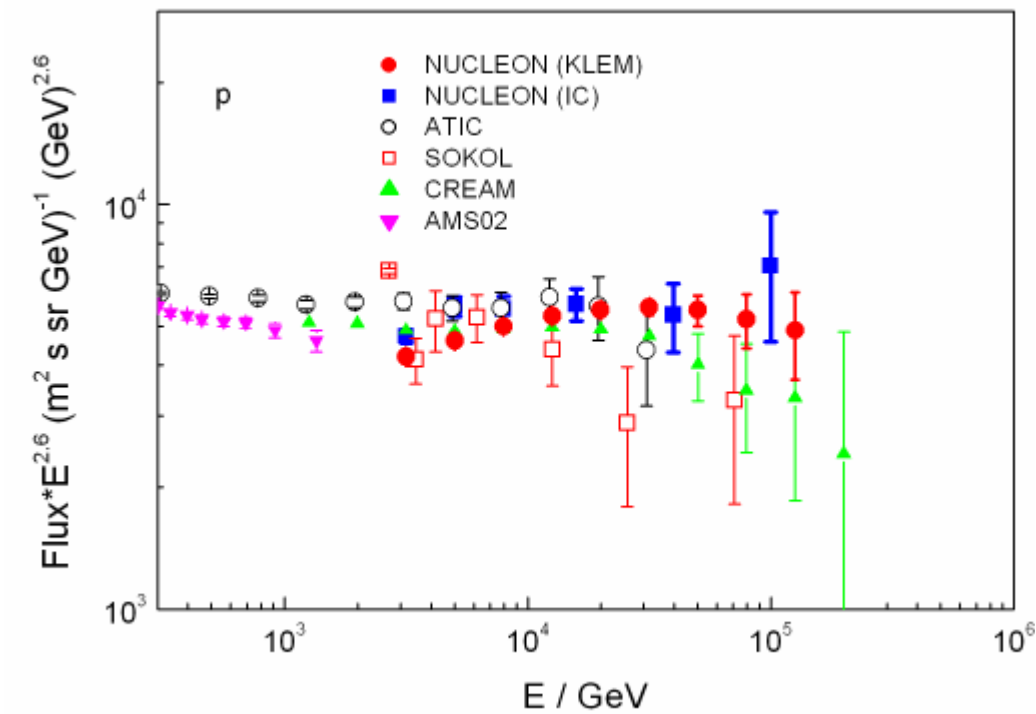
**Charge resolution of four  
silicon planes detector  
~0.2 charge units**



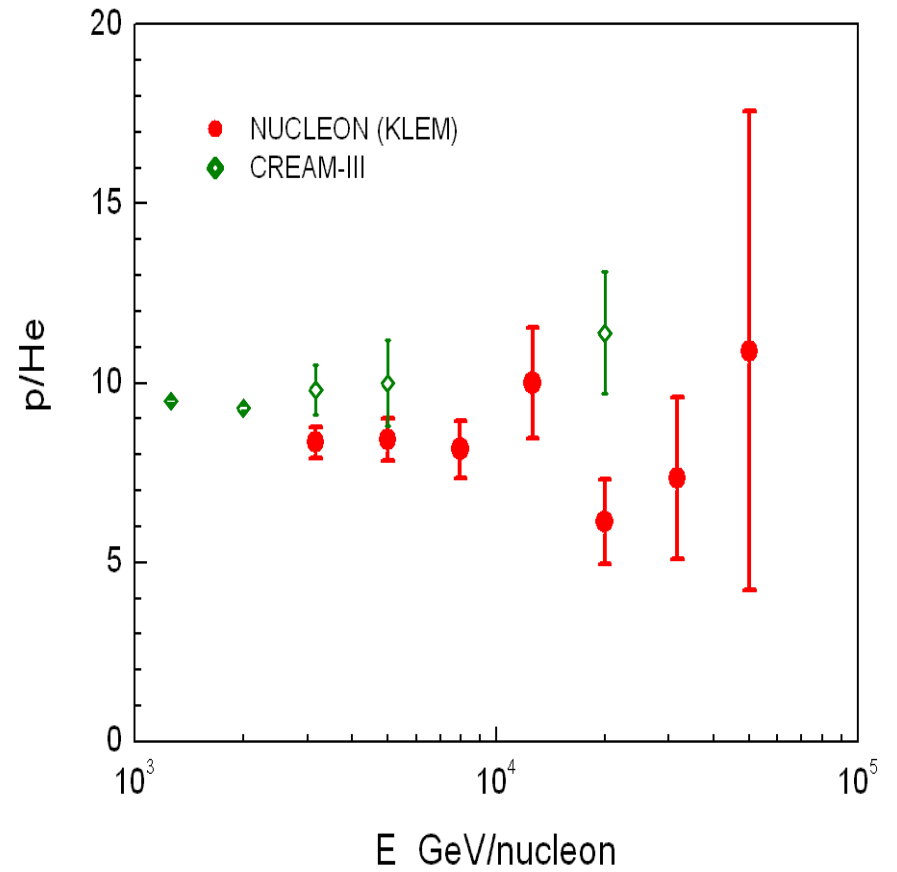
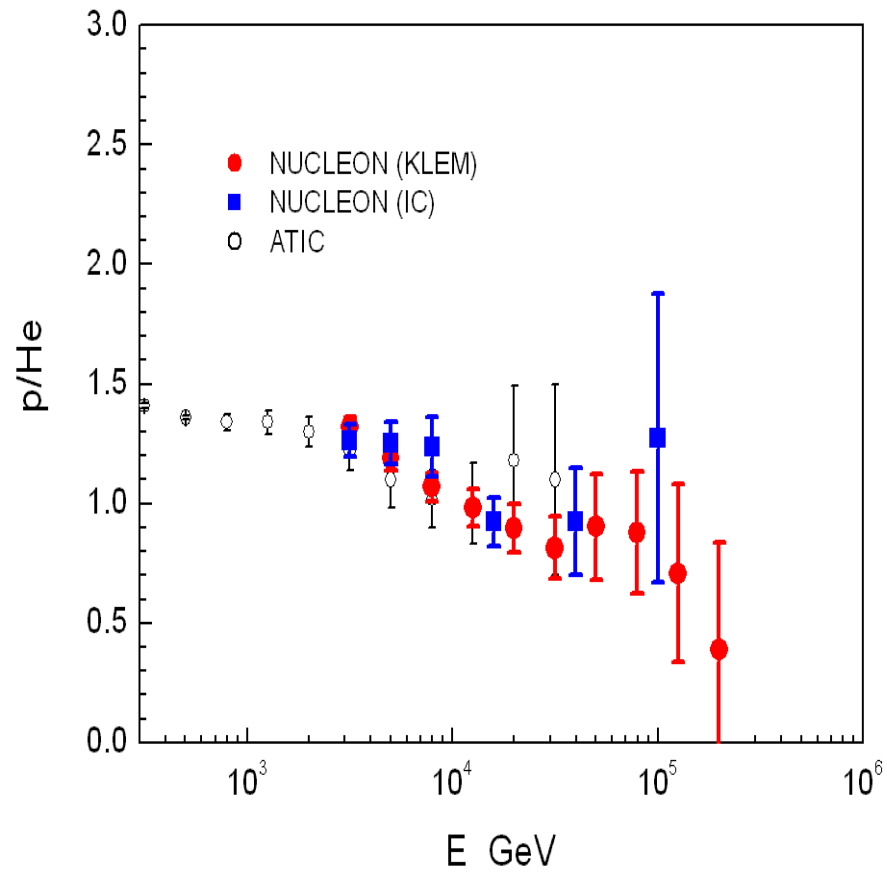
## An example of an event «portrait»



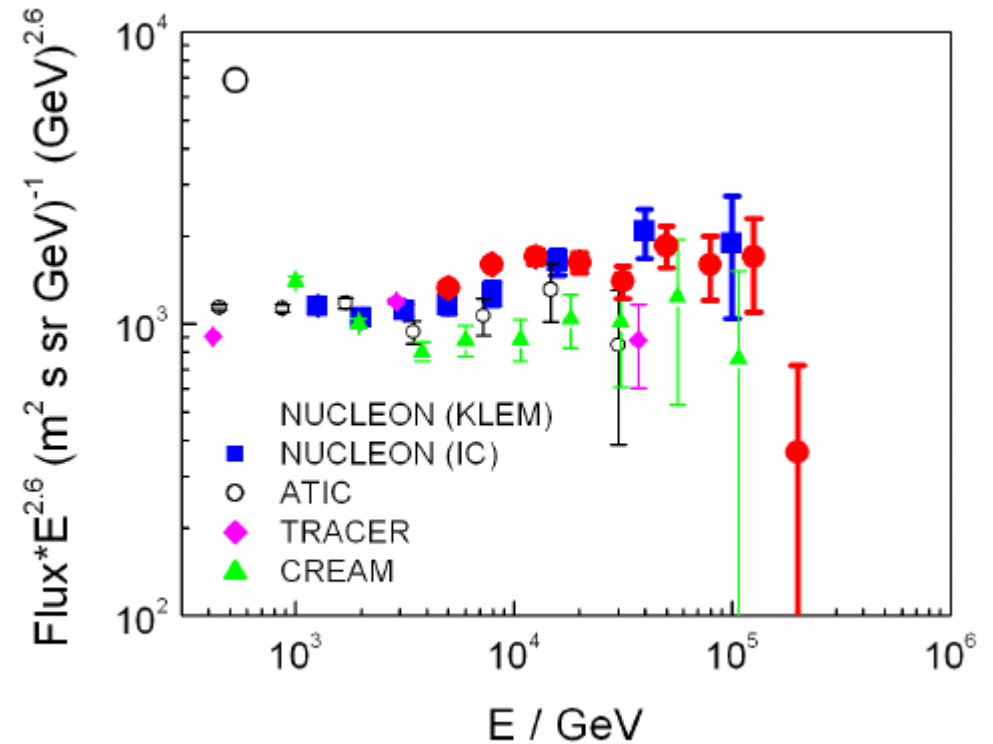
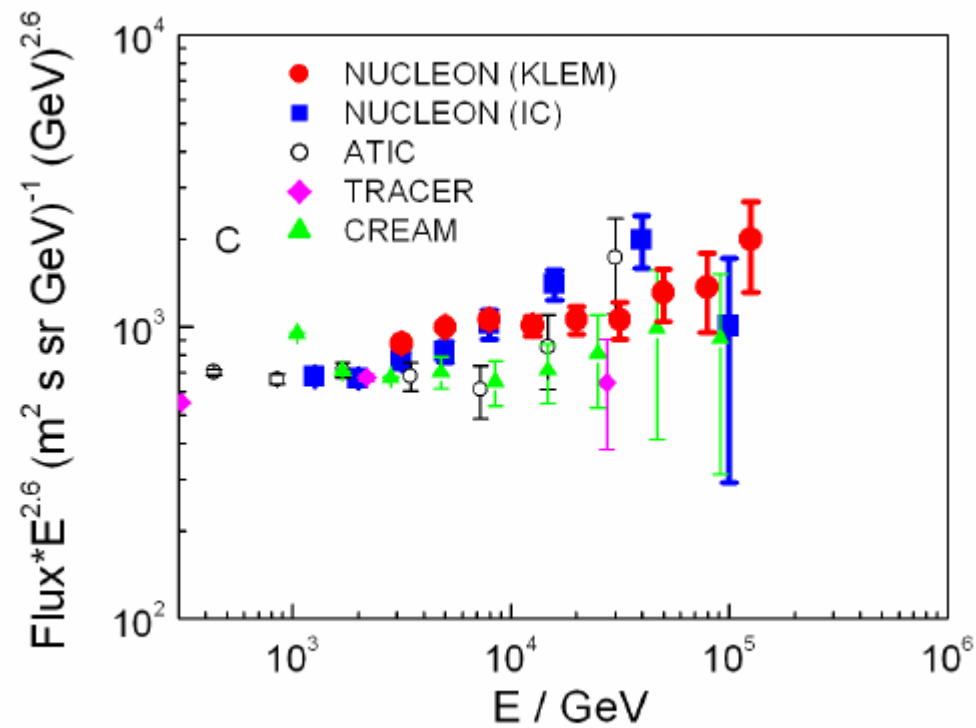
# Protons and Helium



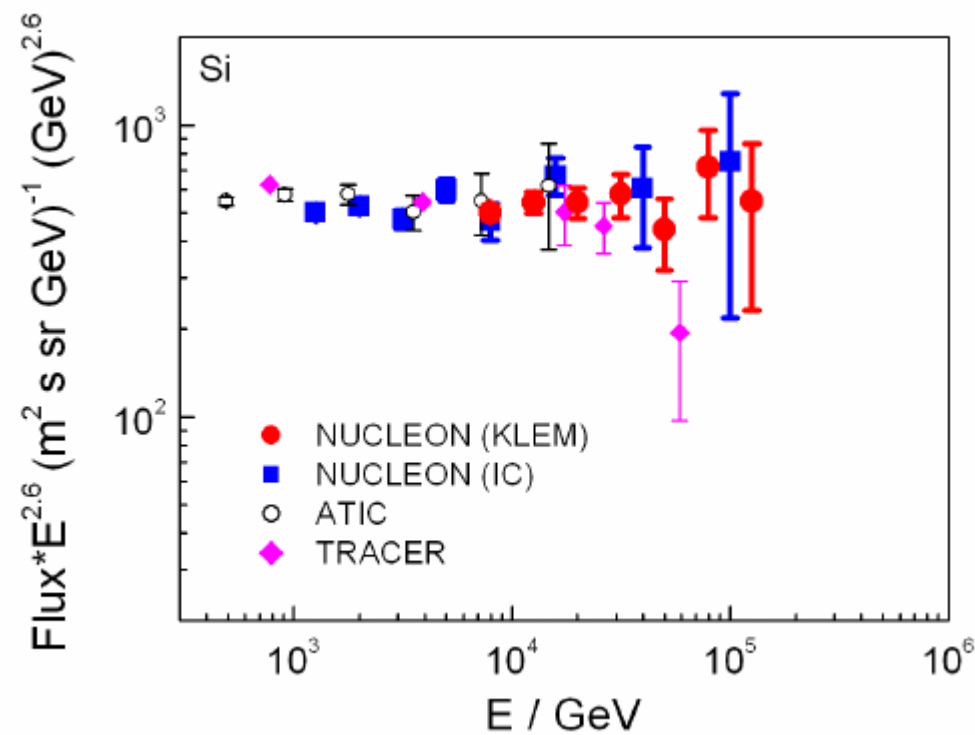
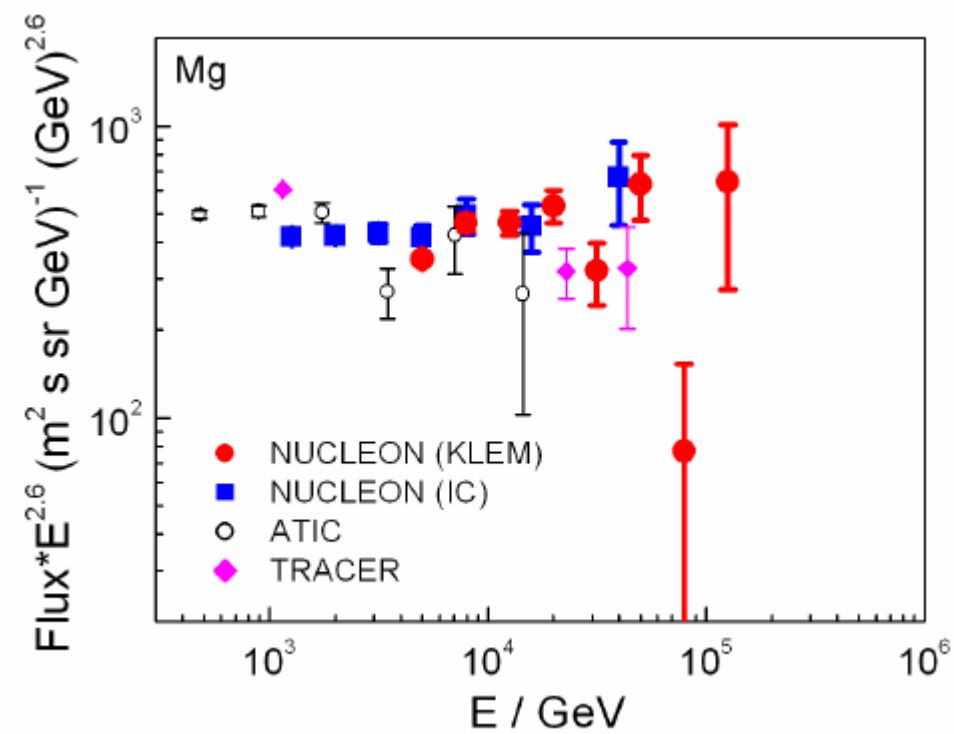
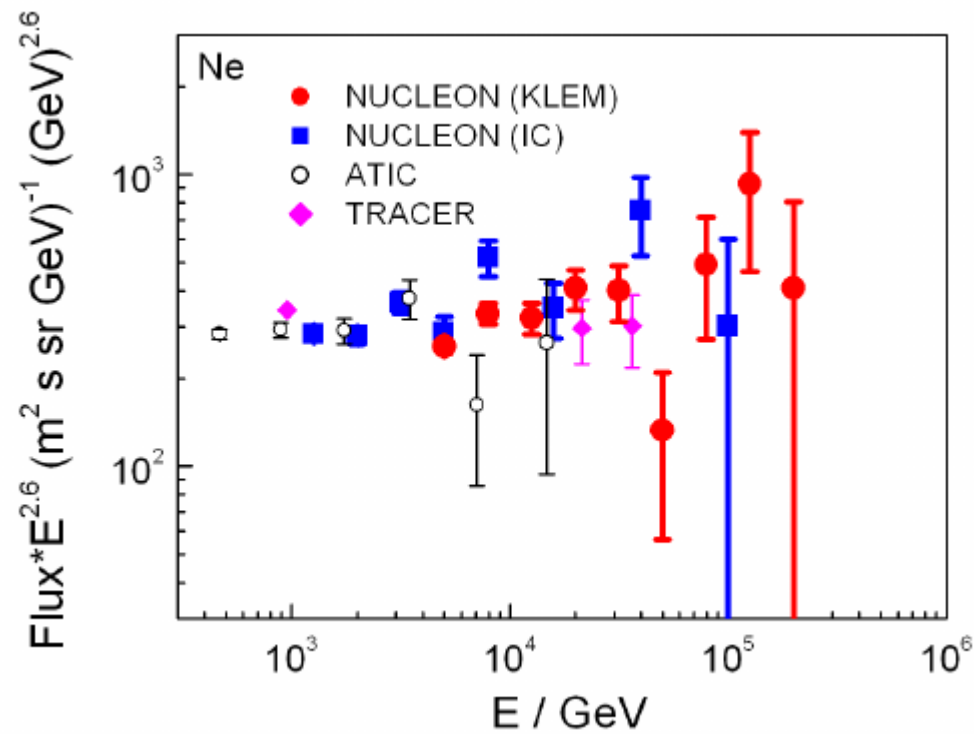
# p/He ratio



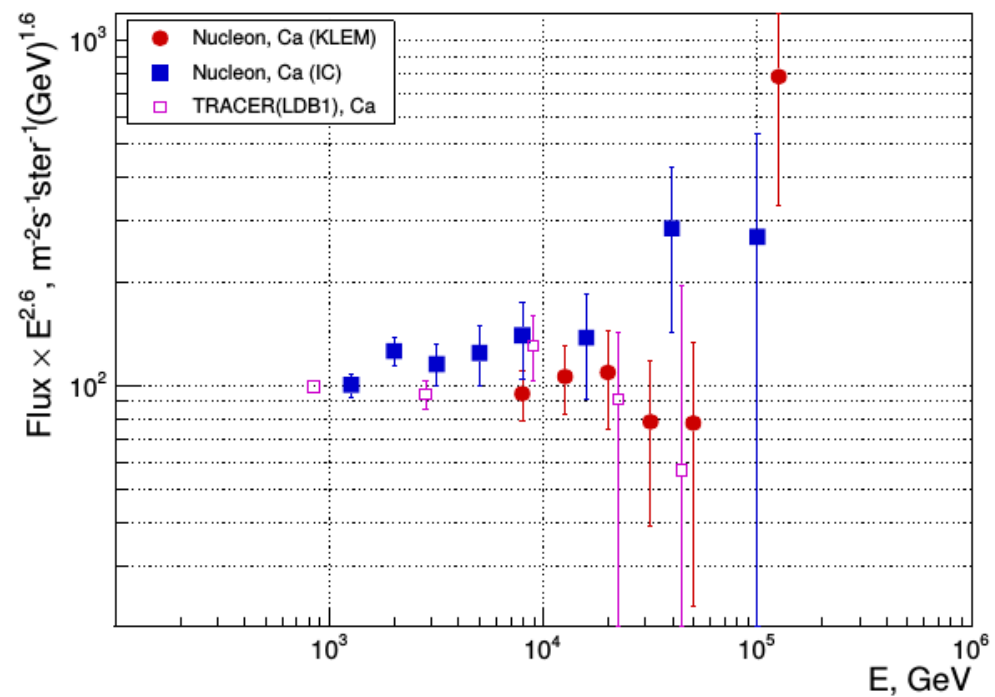
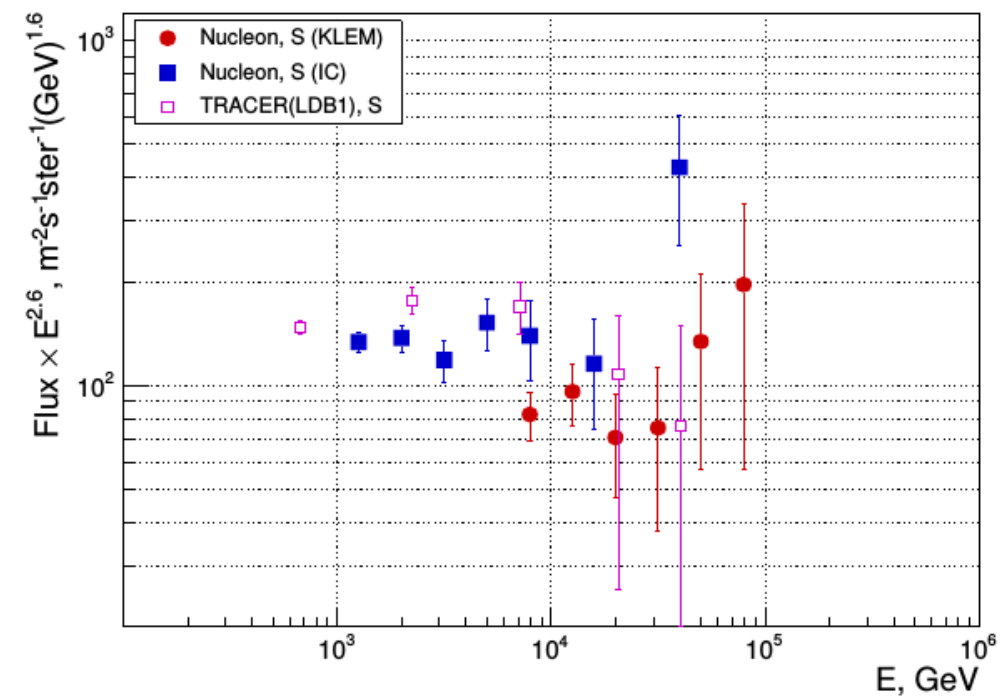
## Carbon and Oxygen: hard ( $\gamma \approx 2.4$ ) above $\sim 3$ TeV



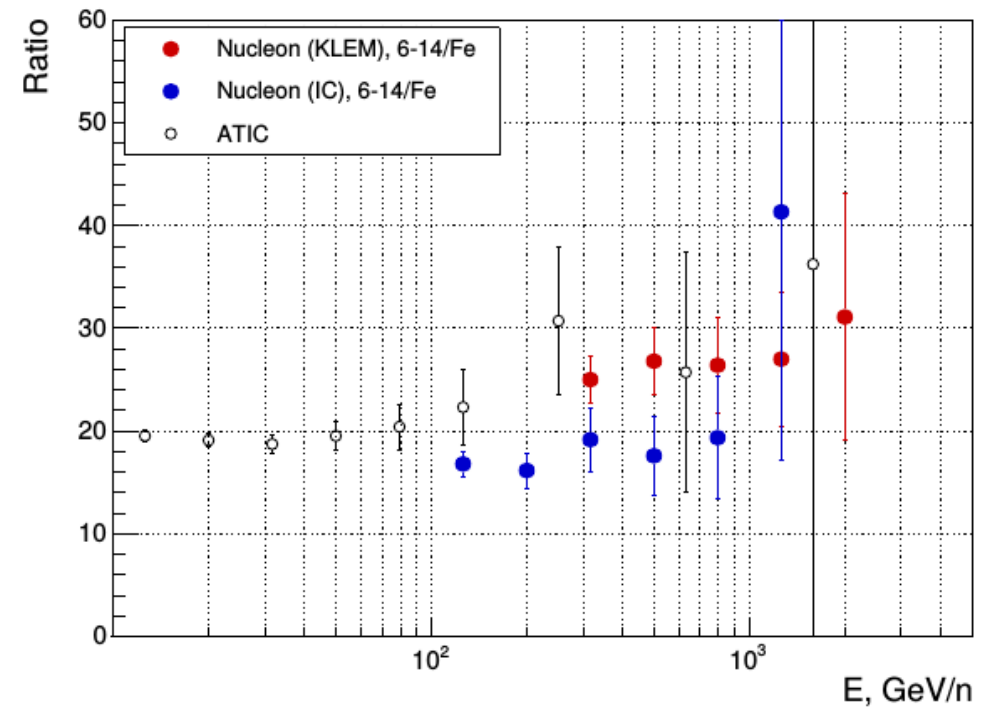
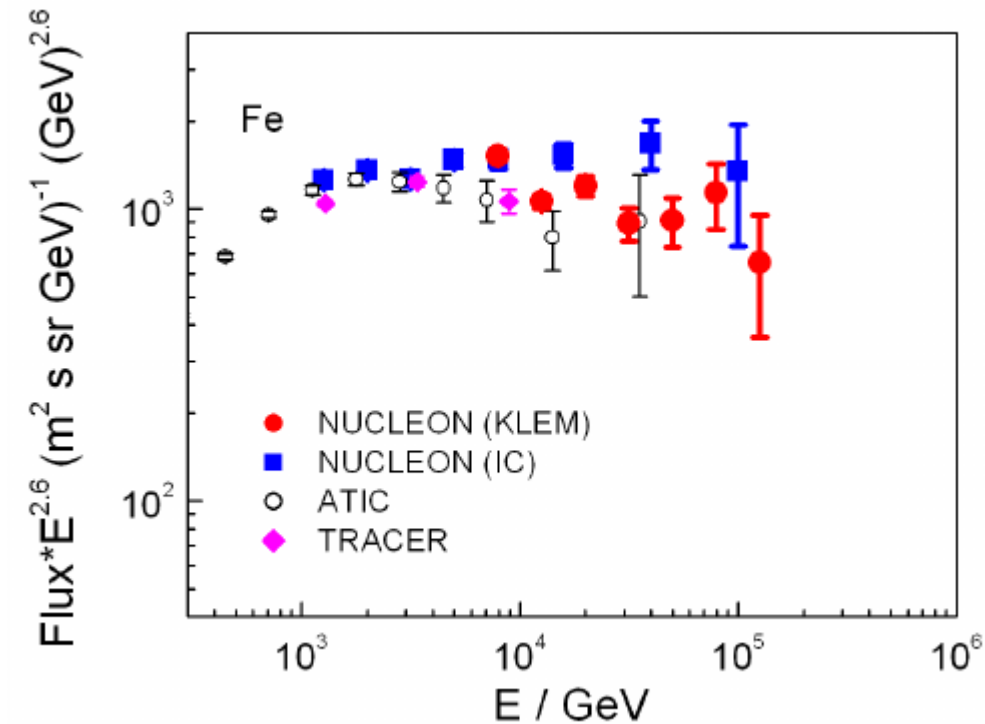
# Ne-Mg-Si



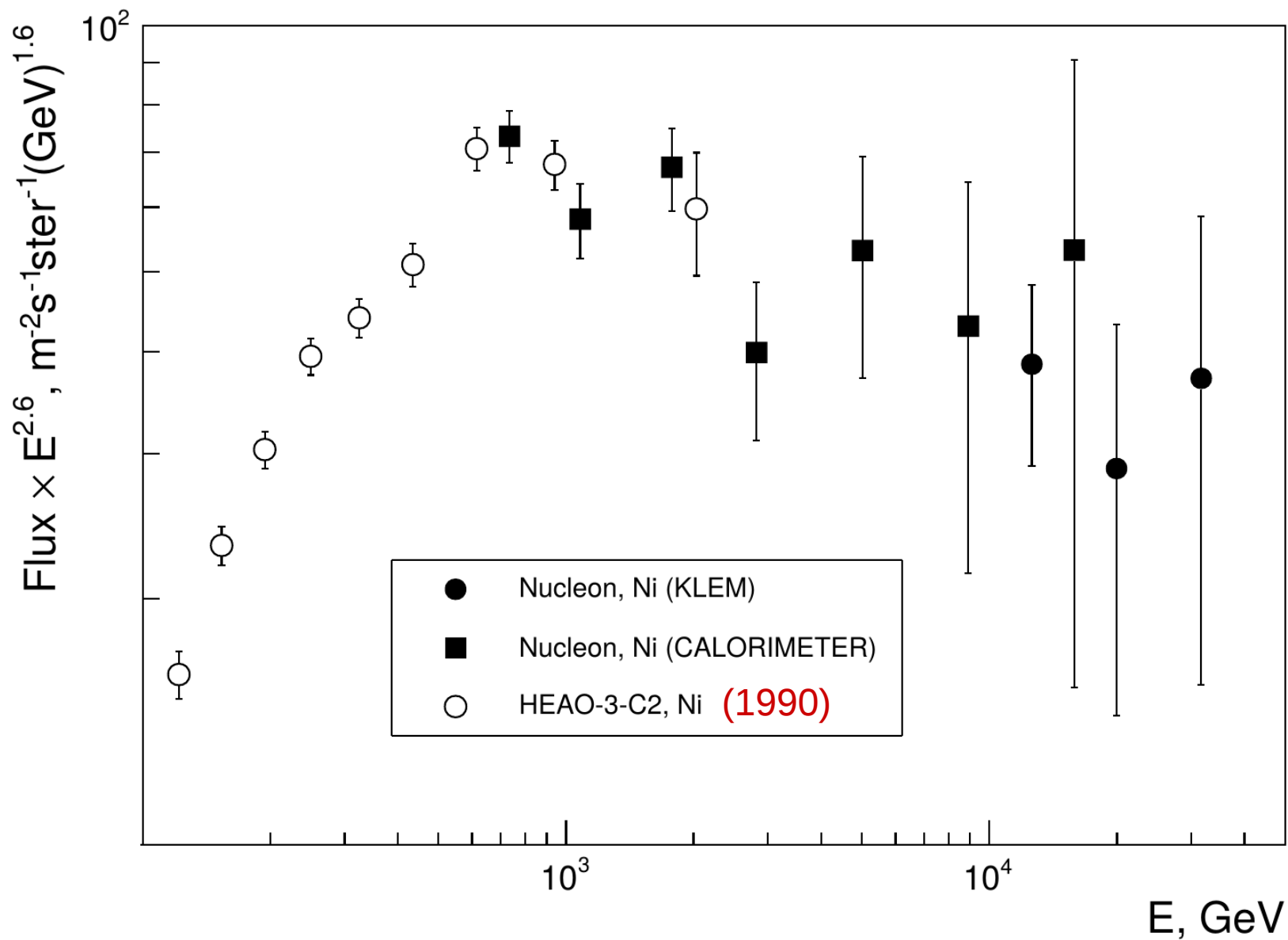
# S and Ca - hints of complicated behavior, more statistics are needed



# Iron spectrum ( $\gamma \approx 2.6$ ) - softer, than the spectra of other heavy nuclei? ( $Z = 6-14$ )

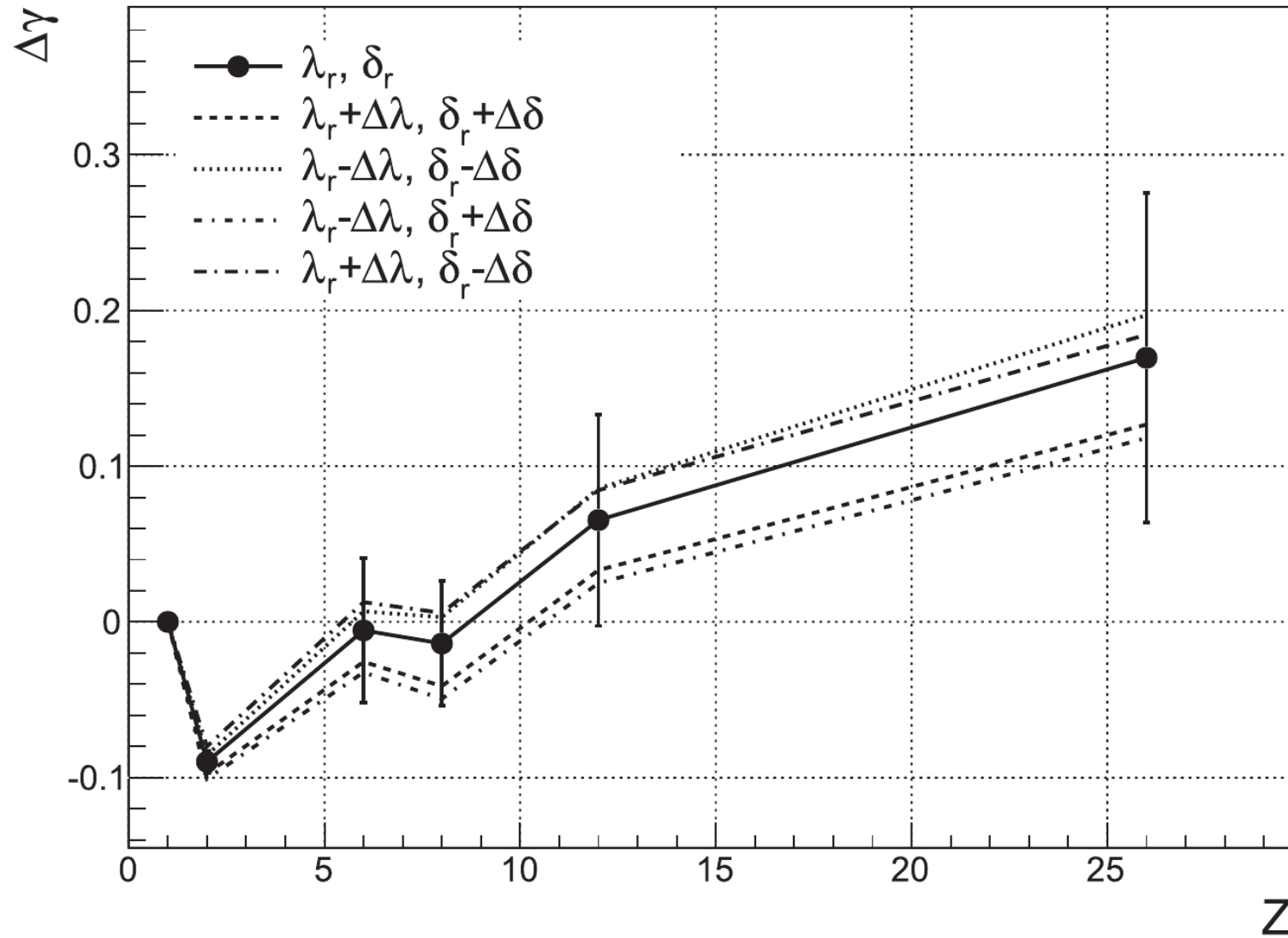


# Nickel (Z=28) spectrum. $\gamma = 2.83 \pm 0.09$

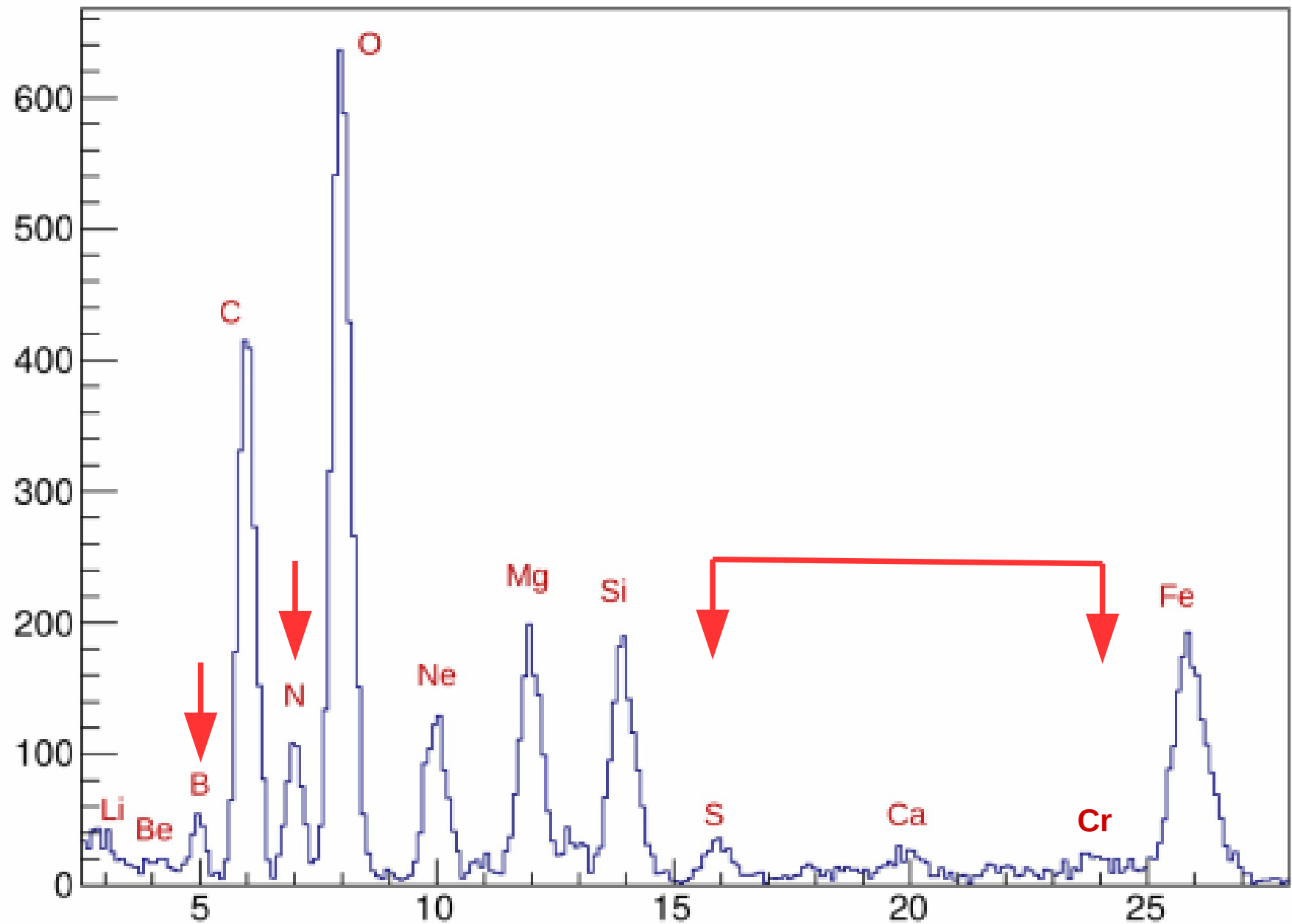


## ATIC: the spectra of nuclei in source are softer for higher charges

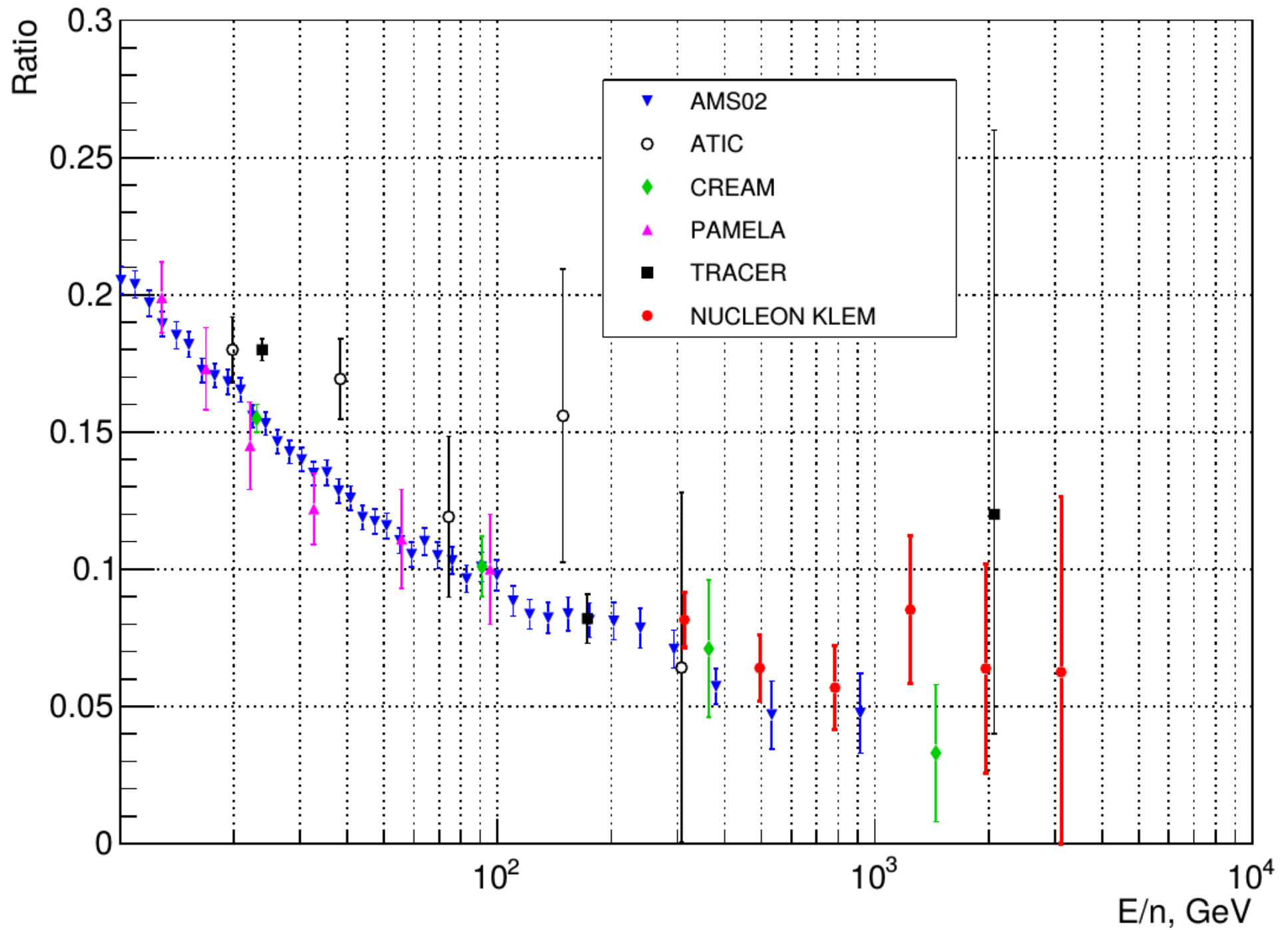
The Astrophysical Journal, V.837, 77 (2017)



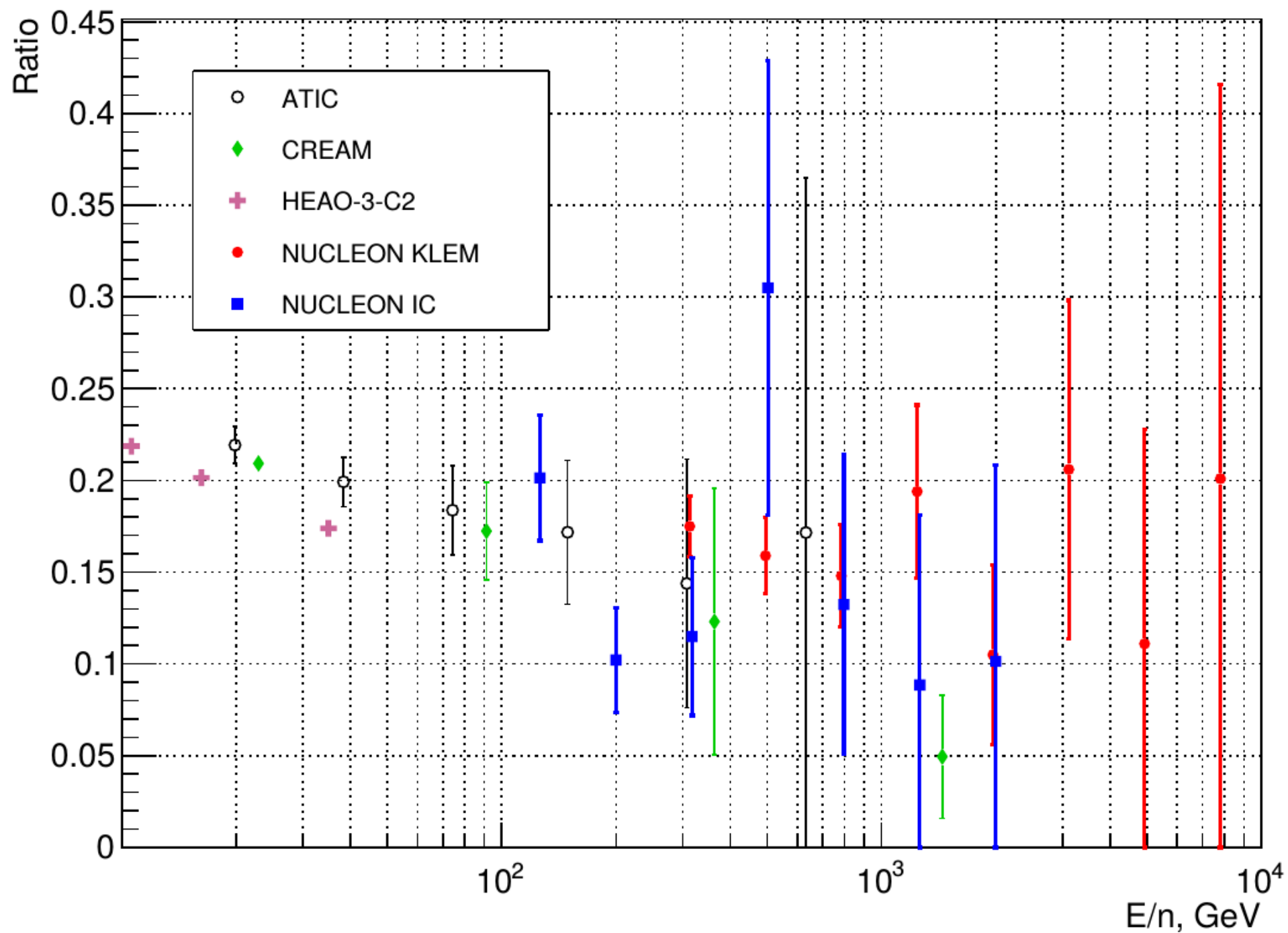
# Spectra of secondary nuclei



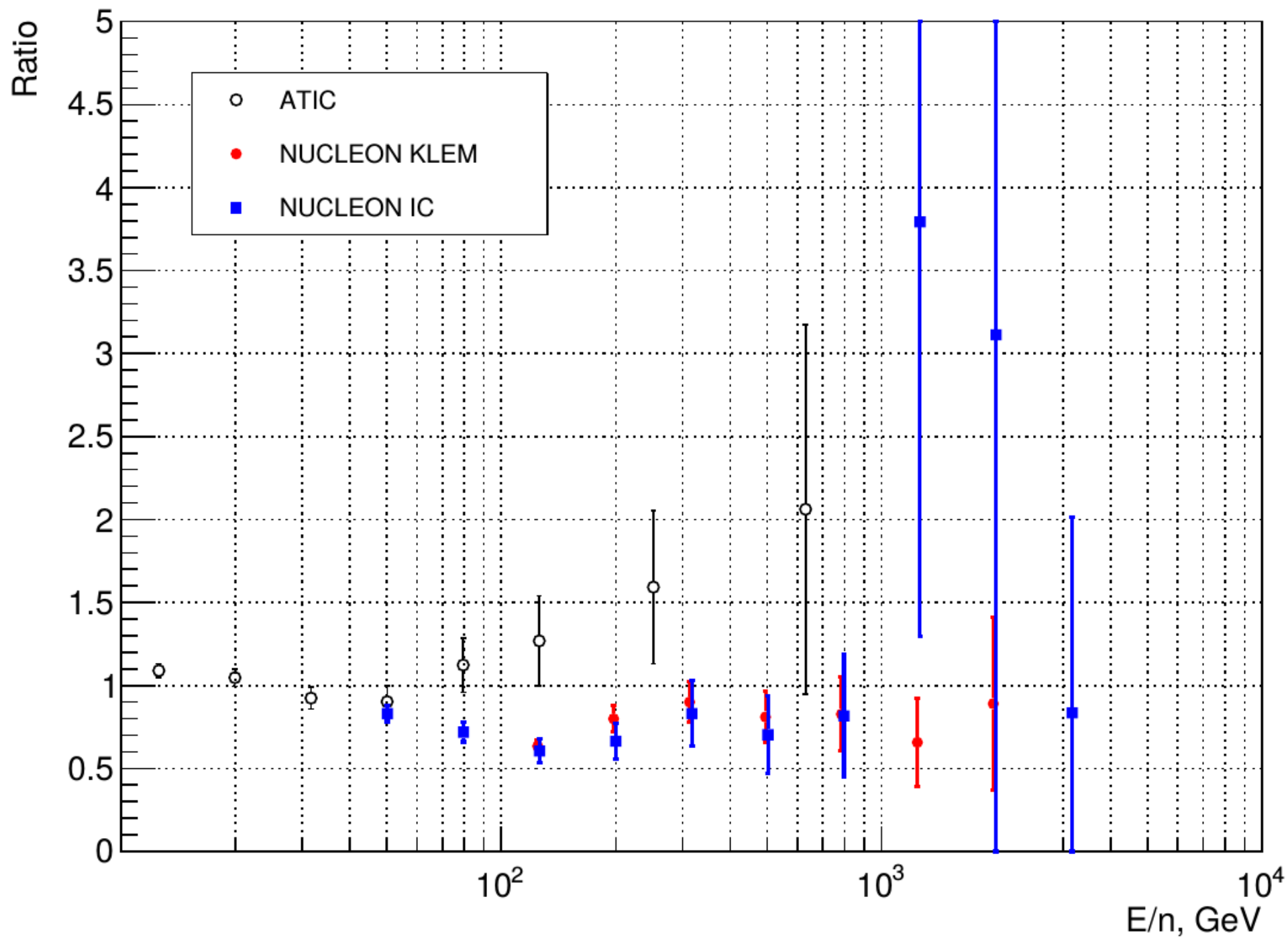
# B to C ratio



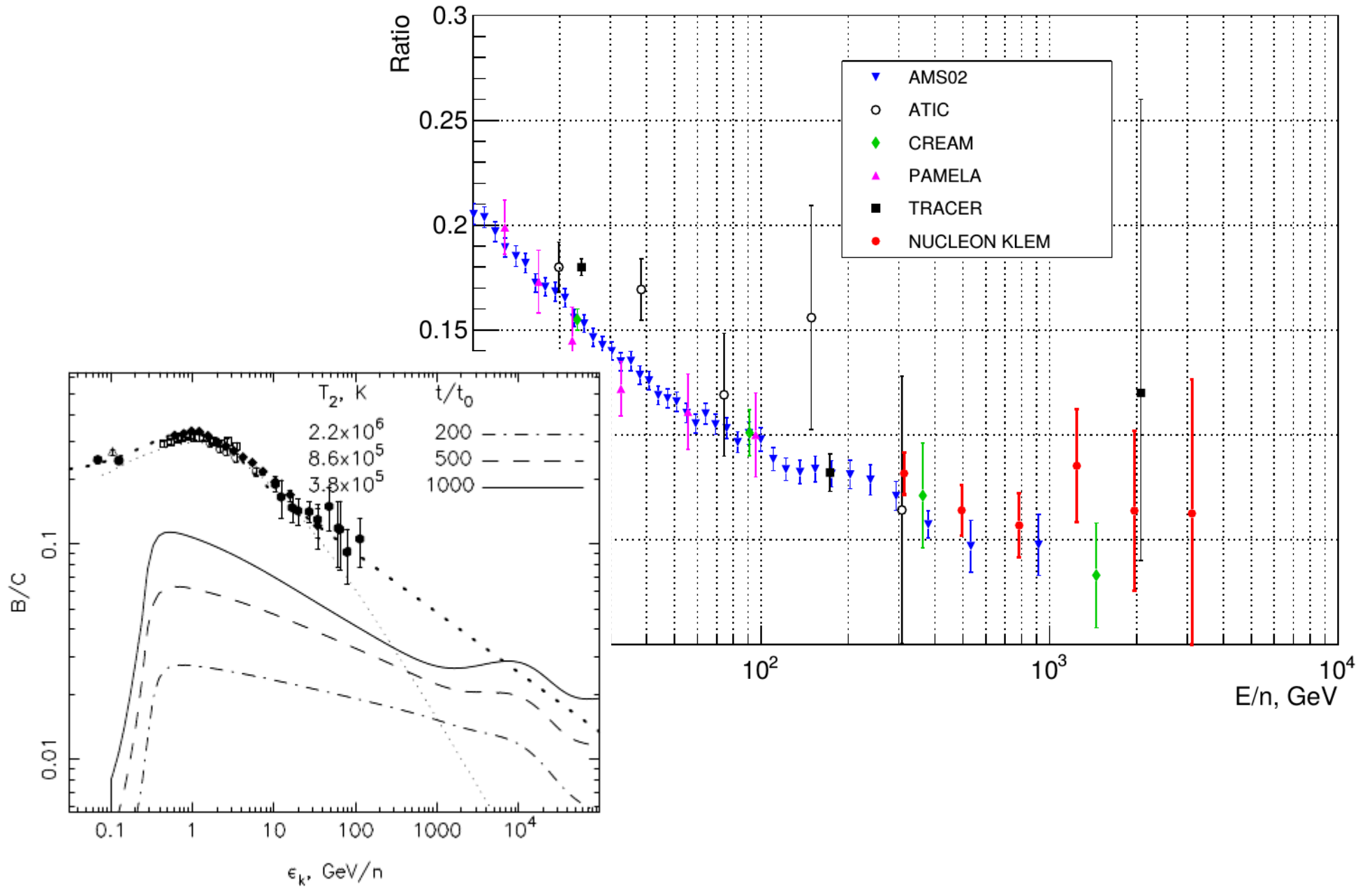
# N to O ratio



# SubFe to Fe ratio



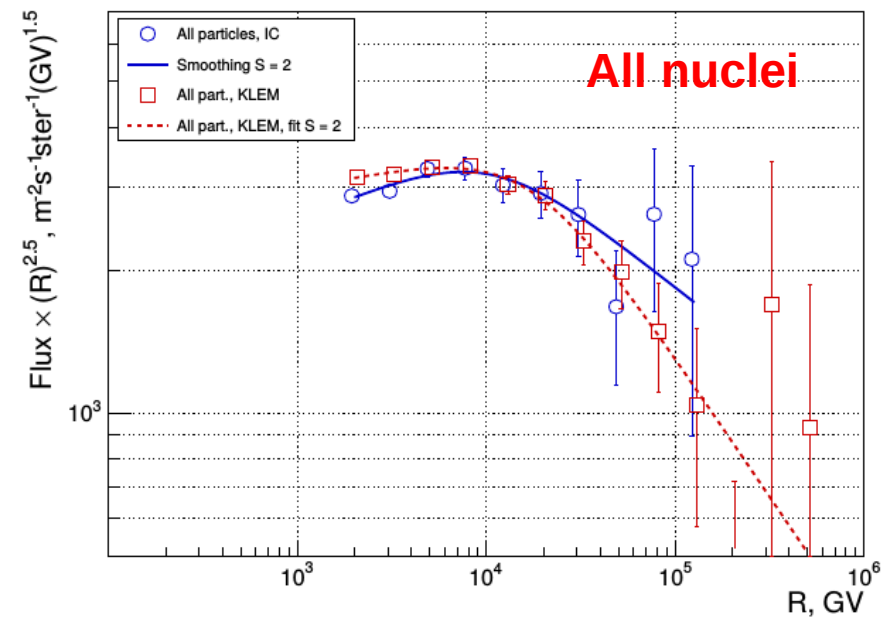
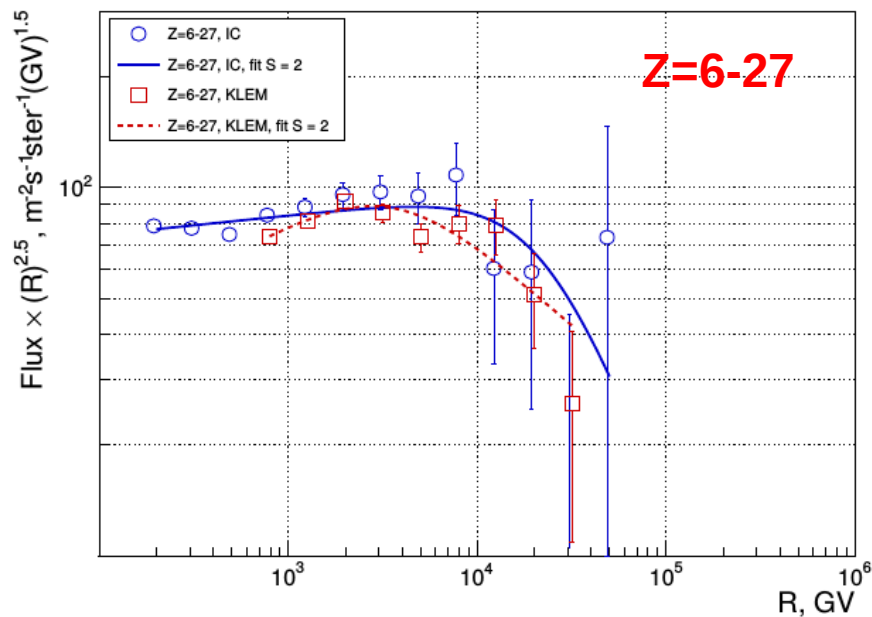
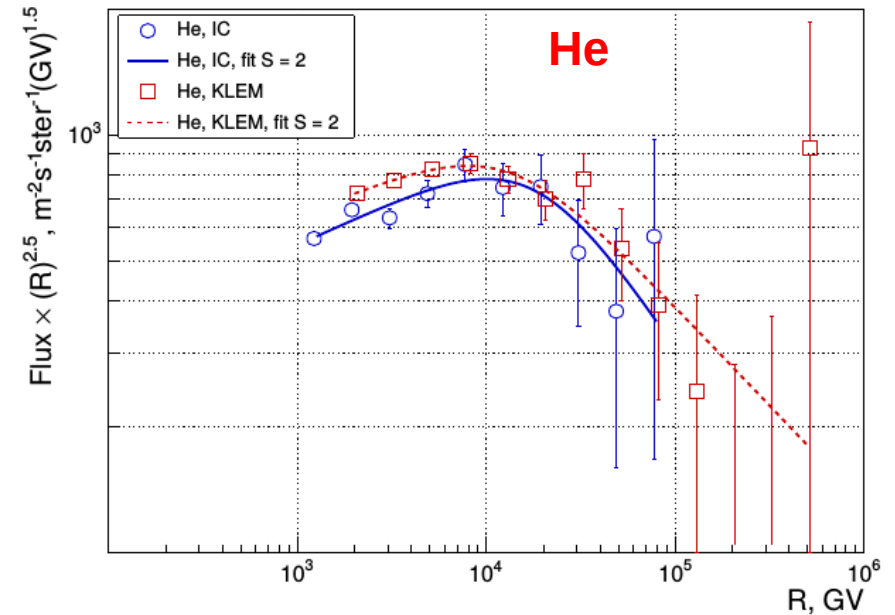
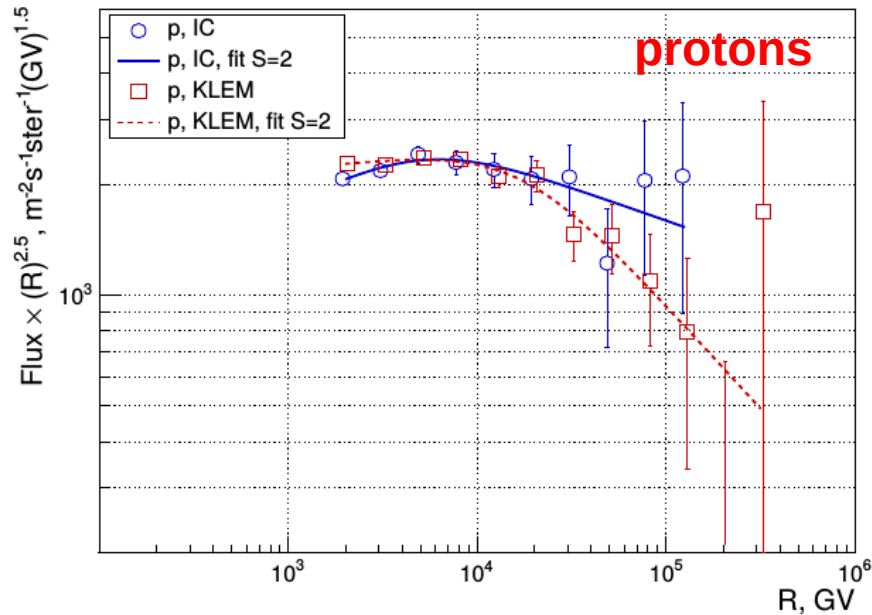
# B to C ratio



E. G. Berezhko, L. T. Ksenofontov, V. S. Ptuskin, V. N. Zirakashvili, H. J. Voelk.  
 Astron.Astrophys. 410 (2003) 189-198 (arXiv:astro-ph/0308199v1).

# A new universal cosmic-ray knee near the magnetic rigidity 10 TV

JETP Lett. V.108, No 1, P. 5 (2018) [arXiv:1805:07119]



# Statistical significance of the 10 TV - knee: Monte Carlo method 10000 MC tests

**S = 2 - smooth link**

**S =  $\infty$  - point- like link.**

| nucl.<br>group | $S = 2$<br>percents( $\sigma$ ) | $S = \infty$<br>percents( $\sigma$ ) | single $\gamma$<br>percents( $\sigma$ ) |
|----------------|---------------------------------|--------------------------------------|-----------------------------------------|
| 1              | 2                               | 3                                    | 4                                       |
| IC             |                                 |                                      |                                         |
| p              | 99.31(2.70)                     | 98.49(2.43)                          | 53.17(0.73)                             |
| He             | 99.85(3.25)                     | 99.45(2.78)                          | 90.13(1.65)                             |
| 6 $\div$ 27    | 95.94(2.05)                     | 97.93(2.31)                          | 90.00(1.64)                             |
| All            | 99.97(3.62)                     | 99.89(3.26)                          | 91.07(1.70)                             |
| KLEM           |                                 |                                      |                                         |
| p              | >99.99(3.9)                     | >99.99(3.9)                          | 98.67(2.47)                             |
| He             | >99.99(3.9)                     | >99.99(3.9)                          | 99.59(2.93)                             |
| 6 $\div$ 27    | >99.99(3.9)                     | >99.99(3.9)                          | >99.99(3.9)                             |
| All            | >99.99(3.9)                     | >99.99(3.9)                          | >99.99(3.9)                             |

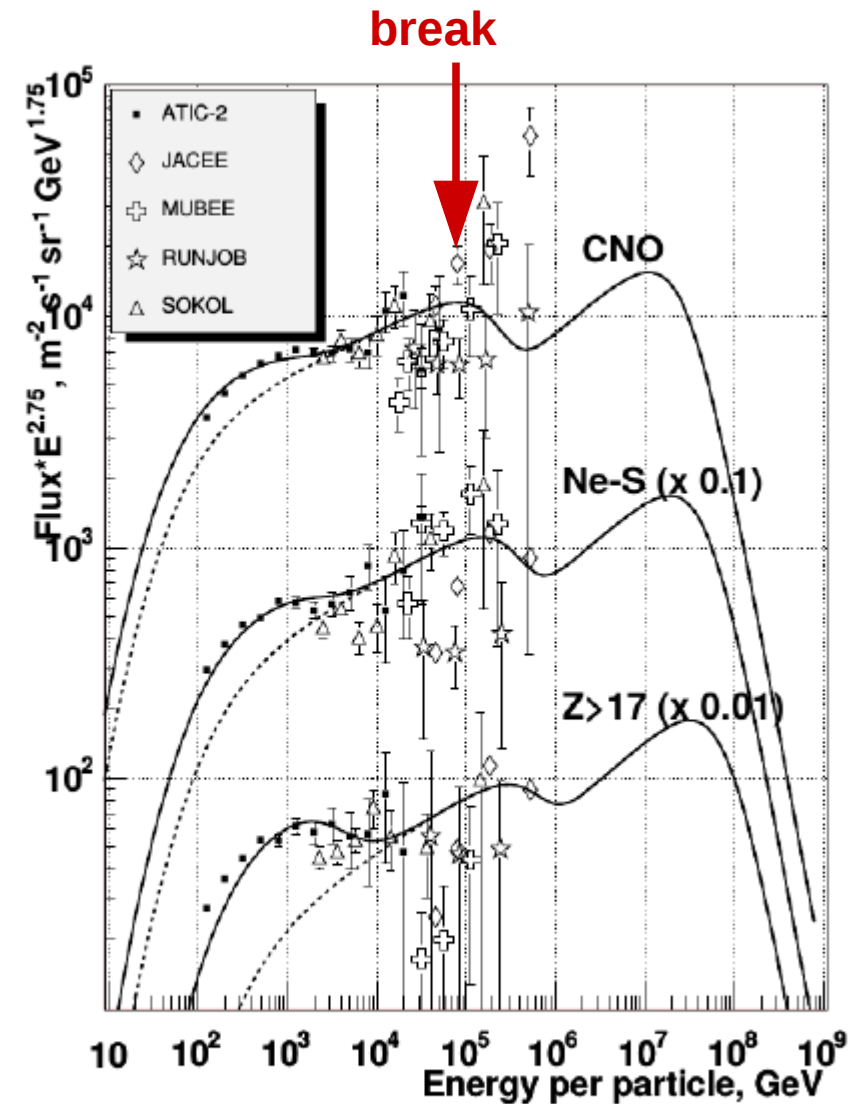
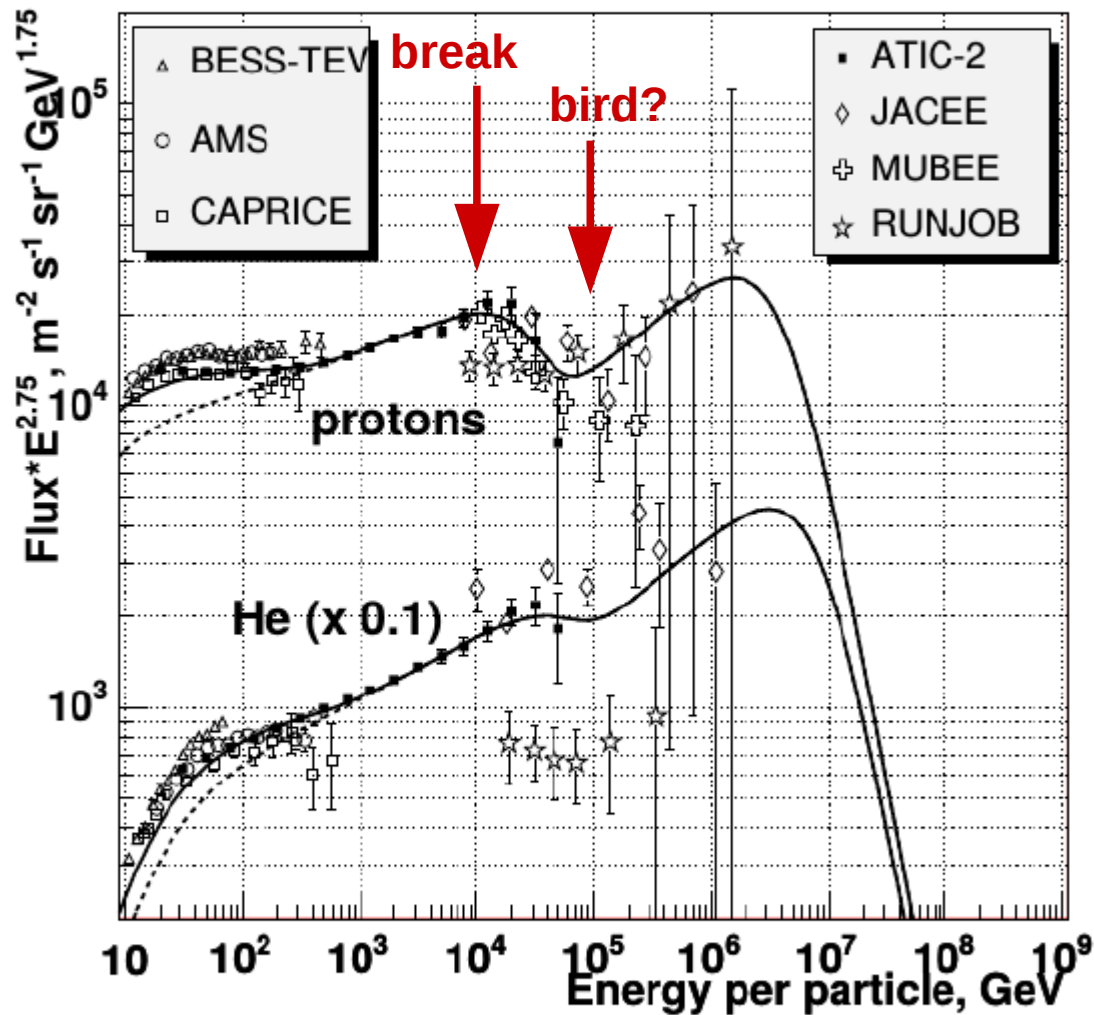
## **The simplest possible explanation of 10TV-knee:**

10TV is a limit of acceleration of some kind of CR source - a nearby one or generic.

### **Question:**

How so soft spectra may be linked with the usual knee (3 PeV) region?

Zatsepin & Sokolskaya model predicts breaks  
near  $R = 10$  TV both  
in spectra of protons and helium,  
and in spectra of heavy nuclei



V.I. Zatsepin and N.V. Sokolskaya.  
A&A, V.458, 2006, pp.1-5

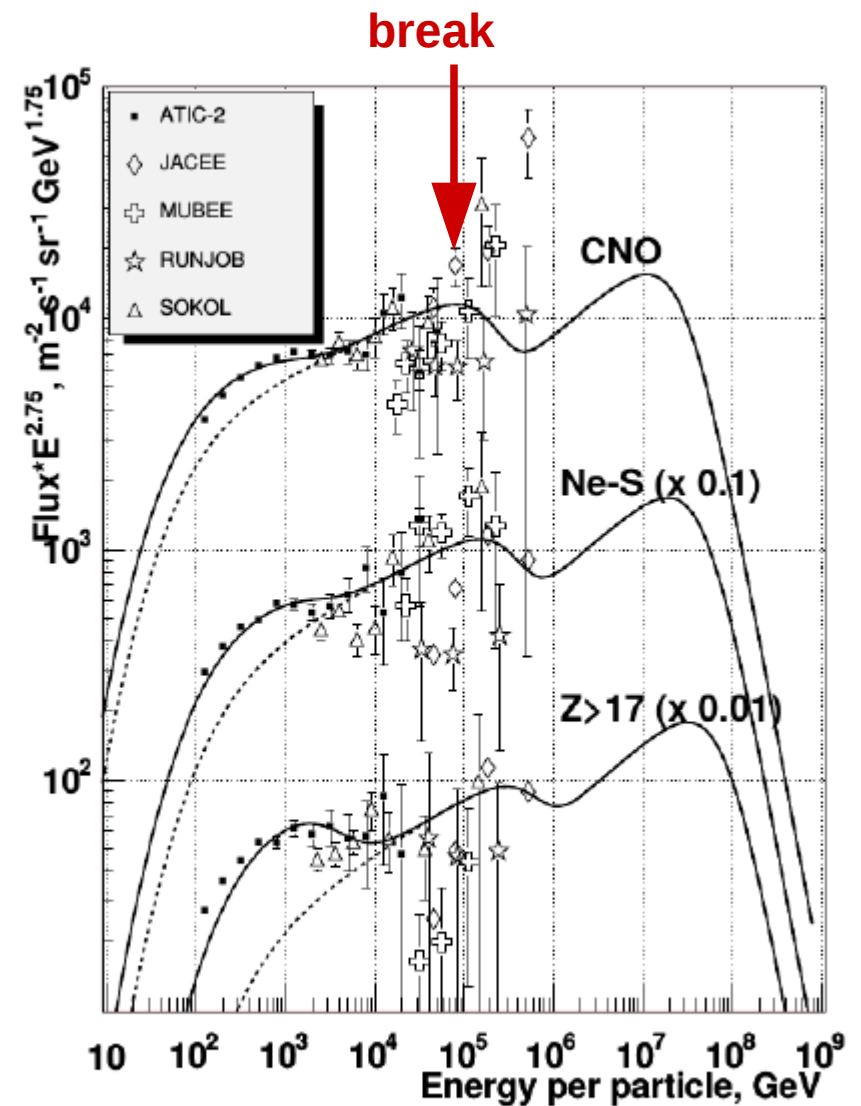
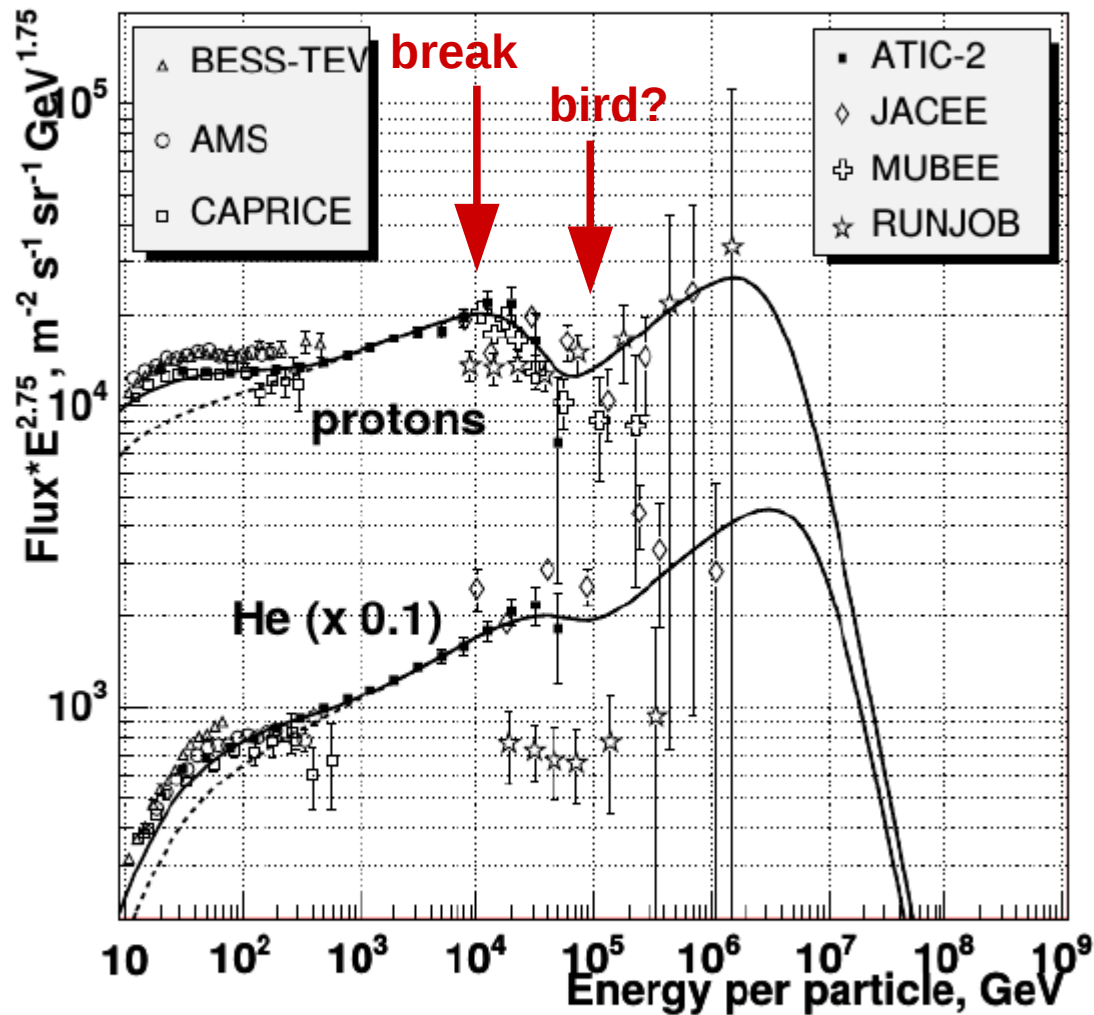
# Conclusions

- The preliminary analysis of the NUCLEON space experiment data gives indications of the existence of a number of features in the energy spectra of cosmic ray nuclei at energies from few TeV to  $\sim 100$  TeV (per particle).
- Secondary to primary ratios do not show decreasing at highest energies.
- A new universal cosmic-ray knee near 10 TV was observed with statistical significance more than 3.9 sigma.

**Thank you  
for attention!**



Zatsepin & Sokolskaya model predicts breaks  
near  $R = 10$  TV both  
in spectra of protons and helium,  
and in spectra of heavy nuclei



V.I. Zatsepin and N.V. Sokolskaya.  
A&A, V.458, 2006, pp.1-5

## All particle spectrum

